

**Instructions:** Treat this exam as you would an open-book **3-hour** in-class exam. You can work on it during any contiguous 3-hour period of your choosing, but once you open the exam you must keep going; you must not stop and finish it later. Scan and email your solution to [dbelitz@uoregon.edu](mailto:dbelitz@uoregon.edu) when you are done, but no later than 5pm on Tuesday, June 8. Do not discuss the exam with anyone before then.

**Note:** A lot of points go into the discussion of the results.

### 1) Bremsstrahlung

As a simple model for an x-ray machine, consider a charged point particle that travels with constant velocity  $v_0$  for times  $t < 0$ , then gets decelerated at a constant rate  $-a$  (with  $a > 0$  an acceleration) for times  $t > 0$  until it comes to a stop at time  $t = T$ .

- a) Calculate, sketch, and discuss the radiated energy per frequency,

$$\frac{dU}{d\omega} = \frac{2e^2}{3\pi c^3} |\dot{v}(\omega)|^2$$

Which frequency range dominates the spectrum? What would you do to produce radiation with higher frequency?

- b) Calculate the total radiated power  $\mathcal{P}$  (keeping  $a$  fixed), and show that you recover the Larmor formula

$$\mathcal{P} = \frac{2e^2}{3c^3} a^2$$

for a uniformly accelerated point particle.

*hint:*  $\int_0^\infty dx \frac{\sin^2 x}{x^2} = \frac{\pi}{2}$

(15 points)

### 2) Pulsar

The pulsar SGR1806-20 is observed to have a rotational period  $T = 7.5$  s that increases at a rate  $\dot{T} = 8 \times 10^{-11}$ . As a very crude model of such a system, consider a rotating homogeneous neutron star with mass  $M = 3 \times 10^{33}$  g and radius  $R = 10^6$  cm that rotates about the  $z$ -axis and has a magnetic moment  $\mathbf{m}$  that lies in the  $xy$ -plane and rotates with the same period as the star as a whole.

- a) Assuming that the increase of the period is due to magnetic dipole radiation, express the magnitude  $m$  of the dipole moment in terms of the parameters given above.

*hint:* Remember that the rotational kinetic energy of a homogeneous sphere is  $U = I\Omega^2/2$ , with  $\Omega$  the angular velocity and  $I = 2MR^2/5$  the moment of inertia of the sphere.

- b) From the expression for the magnetic field generated by a magnetic dipole moment that we derived in ch.2 §3.7 corollary 1, estimate the maximum magnetic field at the surface of the neutron star. (Take  $r = R$  and don't worry about the validity of the far-field approximation.) Determine the maximum magnetic field in units of Gauss and compare with the field the surface of the earth ( $\approx 0.5$  G).

(9 points)