

## Chapter 2 Static solutions of Maxwell's equations

### §1 Poisson's equation

#### 1.1 Electrostatics

Consider M's eqs for static fields: (2)  $\vec{\nabla} \times \vec{E} = 0$

$$(3) \vec{\nabla} \cdot \vec{E} = 4\pi \rho$$

remark: (1) (2) and (3) written  $\vec{E}$  as  $\vec{E} = -\vec{\nabla}\phi$ , (4)  $\vec{\nabla} \cdot \vec{E} = 4\pi \rho$  written  $\vec{\nabla} \cdot \vec{E}$  as  $\vec{\nabla} \cdot \vec{E} = 4\pi \rho$ .  $\rightarrow$  For static fields,  $\vec{E}$  and  $\vec{\nabla}$  decouple.

cf § 1.4  $\rightarrow$  A static  $\vec{E}$ -field is determined by  $\phi$  alone:

$$\boxed{\vec{E}(\vec{x}) = -\vec{\nabla}\phi(\vec{x})} \quad (*)$$

remark: (2) M eq. (2) is automatically satisfied:

$$(\vec{\nabla} \times \vec{E})_i = -(\vec{\nabla} \times \vec{\nabla}\phi)_i = -\epsilon_{ijk} \partial_j \partial_k \phi = 0$$

proposition 1: The electrostatic potential  $\phi$  obeys Poisson's eq.

$$\boxed{\vec{\nabla}^2 \phi(\vec{x}) = -4\pi \rho(\vec{x})} \quad (**)$$

where  $\vec{\nabla}^2 \equiv \Delta \equiv \partial_i \partial^i$  is the Laplace operator

corollary. In vacuum,  $\phi$  obeys the Laplace eq.

$$\boxed{\vec{\nabla}^2 \phi(\vec{x}) = 0}$$

remark: (2') Solutions of Laplace's eq. are called harmonic fcts.

(3)  $\phi(\vec{x}) = \cos t$ ,  $\phi(\vec{x}) = x$ ,  $\phi(\vec{x}) = y$ , and  $\phi(\vec{x}) = z^2 - \frac{1}{2}(x^2 + y^2)$  are all harmonic fcts.

(4) A harmonic fct. can have no extrema, except at infinity.

## 1.2 Magnetostatics

$$\mathbf{d} \perp \mathbf{j} \rightarrow \vec{\mathbf{A}} = \vec{\nabla} \times \vec{\mathbf{A}}$$

remark: (1) This is always true, but is particular for static fields.

(2) M-eq. (1) is automatically fulfilled, via  $\vec{\nabla} \cdot \vec{\nabla} \times \vec{\mathbf{A}} = 0$ .

proposition 1: The Abelian vector potential  $\vec{\mathbf{A}}$  obeys

$$\boxed{\vec{\nabla}^2 \vec{\mathbf{A}}(\vec{\mathbf{x}}) = -\frac{4\pi}{c} \vec{\mathbf{j}}(\vec{\mathbf{x}})} \quad (*)$$

proof:  $(\vec{\nabla} \times \vec{\nabla} \times \vec{\mathbf{A}})_i = \epsilon_{ijk} \partial_j \epsilon_{klm} \partial_l A_m = \epsilon_{kij} \epsilon_{klm} \partial_j \partial_l A_m$   
 $= (\delta_{ik} \delta_{jl} - \delta_{il} \delta_{jk}) \partial_j \partial_l A_m = -\partial_j \partial_j A_i + \partial_i \partial_l A_l$   
 $= -\vec{\nabla}^2 A_i + \partial_i (\vec{\nabla} \cdot \vec{\mathbf{A}})$

Problem 2  $\rightarrow$  We can always choose Lorenz gauge, which makes  $\vec{\nabla} \cdot \vec{\mathbf{A}} = 0$

$$\rightarrow \frac{4\pi}{c} \vec{\mathbf{j}} = \vec{\nabla} \times \vec{\mathbf{A}} = \vec{\nabla} \times \vec{\nabla} \times \vec{\mathbf{A}} = -\vec{\nabla}^2 \vec{\mathbf{A}} \quad \square$$

remark: (1) The components  $(\phi, \vec{\mathbf{A}})$  of the static electromagnetic potential obey Poisson's eq. with  $-\frac{4\pi}{c}$  times the components  $(\rho, \vec{\mathbf{j}})$  of the 4-current as the inhomogeneity.

(4) Poisson's eq. is linear  $\rightarrow$  The most general solution is a particular solution plus the most general solution of the Laplace eq. (see ch 5.1 roughly 2).

(5) §1.1 remark (4)  $\rightarrow$  The def solution of Laplace's eq. that vanishes at infinity is the zero solution!

$\rightarrow$  In a finite system, there is def one physical solution of Poisson's eq.

(6) This statement was formulated in a bit more detail in the next section.

## §2 Solutions of Poisson's equation

### 2.1 The general solution of Poisson's equation

proposition: Every Fourier transformable solution of Poisson's eq. is uniquely determined by the inhomogeneity  $f$  via

$$\boxed{\varphi(\vec{x}) = \int \frac{d\vec{k}}{(2\pi)^3} e^{i\vec{k}\vec{x}} \frac{4\pi}{k^2} \hat{f}(\vec{k})} \quad (*)$$

proof: §1.1 (22)  $\rightarrow -k^2 \hat{\varphi}(\vec{k}) = -4\pi \hat{f}(\vec{k})$

$$\rightarrow \hat{\varphi}(\vec{k}) = 4\pi \hat{f}(\vec{k}) / k^2$$

Fourier backtransf.  $\rightarrow (*)$   $\square$

remark: (1) Thanks to the theory developed in §10 <sup>(2.2.14)</sup>, the class of solutions that can be constructed in this way is large!

(2) §1.2 remark (5) follows immediately from Fourier theory

$$\Delta \varphi(\vec{x}) = 0 \Leftrightarrow -k^2 \hat{\varphi}(\vec{k}) = 0 \Leftrightarrow \varphi(\vec{x}) = 0 \quad \forall \vec{x} \neq 0$$

$$\Leftrightarrow \varphi(\vec{x}) \equiv \text{const.}$$

(3) All of  $u_0$  is constant with §1.2 remark (4).

### 2.2 The Coulomb potential

Consider a point charge:  $\rho(\vec{x}) = e \delta(\vec{x})$  where  $\delta(\vec{x}) := \delta(x)\delta(y)\delta(z)$

Answer: The electrostatic potential resulting from a point charge is

the Coulomb potential

$$\boxed{\varphi(\vec{x}) = \frac{e}{r}} \quad \text{with } r = |\vec{x}|$$

proof:  $\hat{\rho}(\vec{x}) = e \rightarrow \hat{\varphi}(\vec{x}) = 4\pi e / r^2$

610 Problem 3.5.  $\rightarrow \underline{\underline{\varphi(\vec{x}) = e/r}}$   $\circ$

Remark: (1) We have now derived the work's potential (from a least-action principle) that we had postulated in PHYS 63.

Woolley: The electric field of a point charge is

$$\underline{\underline{\vec{E}(\vec{x}) = e \vec{x} / r^3}}$$



$\vec{E} \equiv \hat{E}(\vec{x})$

proof:  $\underline{\underline{\vec{E}(\vec{x}) = -\vec{\nabla} \varphi(\vec{x}) = -\vec{\nabla} \cdot \frac{e}{r} = e \frac{\vec{x}}{r^3}}}$   $\circ$

etc.

$\hat{r} \equiv \vec{x} / r$

Remark: (2) The  $\vec{E}$ -field of a point charge is purely radial and isotropic.

## 2.2 Poisson's formula

Proposition: Let  $\rho(\vec{x})$  be a charge distribution whose Fourier coeffs exist. Then

$$\underline{\underline{\varphi(\vec{x}) = \int d\vec{y} \frac{\rho(\vec{y})}{|\vec{x} - \vec{y}|}}} \quad \underline{\underline{\text{Poisson's formula}}}$$

proof: § 2.1 (2)  $\rightarrow \varphi(\vec{x})$  is the Fourier coefficients of the product

$$\varphi_{\vec{x}} = v_{\vec{x}}^c \int \vec{x} \quad \text{with } v_{\vec{x}}^c = 4\pi / r^2$$

that the Fourier coefficients of  $4\pi / r^2$  is  $1/|\vec{x}|$ , and that of  $\int \vec{x}$  is  $\rho(\vec{x})$ , and by the convolution theorem (610 d2) (or (1.11))

$$\underline{\underline{\varphi(\vec{x}) = \int d\vec{y} v_c(\vec{x} - \vec{y}) \rho(\vec{y}) = \int d\vec{y} \frac{1}{|\vec{x} - \vec{y}|} \rho(\vec{y})}} \quad \circ$$

Remark: (1) For  $\rho(\vec{x}) = e \delta(\vec{x})$  we recover the result...

Table 16.

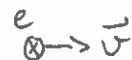
Field of  
ing + disk

Problem 11

spherical  
charge  
distributions

Problem 13

d-dim work's problem



## 2.4 The field of a uniformly moving charge

Consider a charge  $e$  that moves with a constant velocity  $\vec{v}$  with respect to an observer.



Let  $cs'$  be the inertial frame in which the charge is at rest.

$$\text{§ 2.2} \rightarrow \varphi'(\vec{x}') = e/r' \quad \text{and} \quad \vec{A}'(\vec{x}') = (\varphi'(\vec{x}'), 0)$$

Let  $cs$  be the inertial frame of the observer, and let  $\vec{v} = (v, 0, 0)$ .

$cs$  and  $cs'$  are related by a Lorentz boost in  $x$ -direction.

$$\text{cf § 4.1} \rightarrow x' = \gamma(x - vt), \quad y' = y, \quad z' = z \quad \gamma = \frac{1}{\sqrt{1 - v^2/c^2}}$$

$$\begin{aligned} \text{and } \varphi &= \gamma \varphi' = \gamma \frac{e}{r'} = \gamma \frac{e}{\sqrt{x'^2 + y'^2 + z'^2}} \\ &= \frac{e}{\gamma \left[ \gamma^2 (x - vt)^2 + y^2 + z^2 \right]^{1/2}} = \frac{e}{\left[ (x - vt)^2 + (1 - v^2/c^2)(y^2 + z^2) \right]^{1/2}} \end{aligned}$$

$\rightarrow$  The scalar potential due to the moving charge is

$$\boxed{\varphi(\vec{x}, t) = e/R(\vec{x}, t)} \quad \text{where} \quad \boxed{R(\vec{x}, t) := \frac{\left[ (x - vt)^2 + (1 - v^2/c^2)(y^2 + z^2) \right]^{1/2}}{1 - v^2/c^2}} \\ = \frac{r'(\vec{x}, t)}{\gamma^2}$$

and the vector potential is

$$\boxed{\vec{A}(\vec{x}, t) = \gamma \frac{\vec{v}}{c} \varphi' = \frac{\vec{v}}{c} \varphi(\vec{x}, t) = \frac{e\vec{v}}{cR(\vec{x}, t)}}$$

Now consider the fields. In  $cs'$  we have

$$\underline{\vec{E}'(\vec{x}') = e\vec{x}'/r'^3}, \quad \underline{\vec{J}'(\vec{x}') = 0}$$

$$\begin{aligned} \text{cf § 4.2} \rightarrow E_x &= E_x' = \frac{ex'}{r'^3}, \quad E_y = \gamma E_y' = \gamma \frac{ey'}{r'^3}, \quad E_z = \gamma E_z' = \gamma \frac{ez'}{r'^3} \\ &= \frac{e}{\gamma^2} \frac{x - vt}{(R(\vec{x}, t))^3}, \quad = \frac{e}{\gamma^2} \frac{y}{(R(\vec{x}, t))^3}, \quad = \frac{e}{\gamma^2} \frac{z}{(R(\vec{x}, t))^3} \end{aligned}$$

let  $\vec{R} = (x-ut, y, z)$

$\vec{E}(\vec{x}, t) = \frac{e}{r^2} \frac{\vec{R}(t)}{(R^*(\vec{x}, t))^2}$

electric field  
seen by the  
observer

$\Delta \perp \S 4.2 \rightarrow \vec{D}_x = \vec{D}_x' = 0, \vec{D}_y = -\gamma \frac{v}{c} E_t' = -\frac{v}{c} \vec{E}_t$

$\vec{D}_z = \gamma \frac{v}{c} E_y' = \frac{v}{c} E_y$

$\vec{D}(\vec{x}, t) = \frac{1}{c} \vec{v} \times \vec{E}(\vec{x}, t)$

magnetic field seen  
by the observer

discussion of  $\vec{E}(\vec{x}, t)$ :

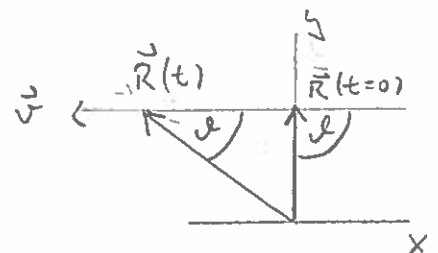
let  $\theta$  be the angle between  $\vec{v}$  and  $\vec{R}$

$\rightarrow \frac{\sqrt{y^2+z^2}}{|\vec{R}|} = \sin \theta \rightarrow y^2+z^2 = R^2 \sin^2 \theta$

$\rightarrow (R^*)^2 = R^2 - \frac{v^2}{c^2} (y^2+z^2) = R^2 [1 - \frac{v^2}{c^2} \sin^2 \theta]$

$\vec{E}(\vec{x}, t) = \frac{e}{r^2} \frac{\vec{R}(\vec{x}, t)}{R^3(\vec{x}, t)} \frac{1}{[1 - \frac{v^2}{c^2} \sin^2 \theta(t)]^{3/2}}$

$\vec{R}(\vec{x}, t) = (x-ut, y, z)$



$\rightarrow$  For fixed distance  $R$  from the charge,  $\vec{E}$  is maximal for  $\theta = 0, \pi$ , i.e., in the direction of the motion. This maximal value is

$E_{||} = \frac{e}{R^2} (1 - v^2/c^2)$

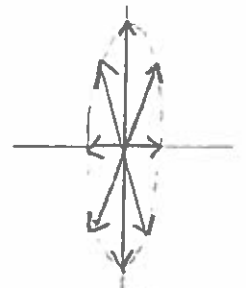
$\vec{E}$  is maximal for  $\theta = \pm \frac{\pi}{2}$ , i.e., in the direction perpendicular to the motion.

The maximal value is

$E_{\perp} = \frac{e}{R^2} \frac{1}{\sqrt{1-v^2/c^2}}$

$\rightarrow$  The field is no longer

isotropic, but squashed in the direction of the motion.



remark (I) Alternative point of view: 4-current in  $cs'$  is

$$j'^{\mu} = (c\rho'(x'), \vec{j}'(x')) \quad \text{with } \rho'(x') = e\delta(\vec{x}') \delta(t'), \vec{j}' = 0$$

$\rightarrow$  Observer in  $cs$  sees

charge density:  $\rho(\vec{x}, t) = \gamma \rho'(\vec{x}', t') = \gamma e \delta(\gamma(x-vt)) \delta(\gamma t) \delta(t)$   
 $= e \delta(x-vt) \delta(t) \delta(y)$

current density:  $\vec{j}(\vec{x}, t) = \gamma \frac{\vec{v}}{c} c\rho' = \vec{v} \rho = e\vec{v} \delta(x-vt) \delta(t) \delta(y)$

Now when D's eyes for this time-dependent 4-current. This is clearly equivalent, but much harder to do!

### 2.5 Electrostatic interaction

Consider a charge density  $\rho(\vec{x})$ .

proposition: The energy of the electric field produced by  $\rho(\vec{x})$  is

$$U = \frac{1}{2} \int d\vec{x} d\vec{y} \rho(\vec{x}) \frac{1}{|\vec{x}-\vec{y}|} \rho(\vec{y})$$

proof:  $U \stackrel{1.6}{=} \int d\vec{x} \vec{E}(\vec{x}) \cdot \vec{E}(\vec{x}) = \frac{1}{2\epsilon_0} \int d\vec{x} \vec{E}(\vec{x}) \cdot \vec{\nabla} \phi(\vec{x})$   
 $= \frac{1}{2\epsilon_0} \int d\vec{x} \underbrace{\vec{\nabla} \cdot (\vec{E}(\vec{x}) \phi(\vec{x}))}_{= \int d\vec{y} \vec{E} \cdot \vec{\nabla} \phi = 0 \text{ for } \phi \rightarrow \infty} + \frac{1}{2\epsilon_0} \int d\vec{x} \phi(\vec{x}) \vec{\nabla} \cdot \vec{E}(\vec{x}) \stackrel{\vec{\nabla} \cdot \vec{E} = 4\pi\rho}{=} \frac{1}{2} \int d\vec{x} \phi(\vec{x}) \rho(\vec{x})$   
 $= \frac{1}{2} \int d\vec{x} d\vec{y} \rho(\vec{x}) \frac{1}{|\vec{x}-\vec{y}|} \rho(\vec{y})$

Week 5  
 Problem 5+6-1  
 #16, 17, 18  
 + Take-home  
 Midterm

remark: (I) Let  $\rho(\vec{x})$  be composed of  $N$  localized

charge distributions:  $\rho(\vec{x}) = \sum_{\lambda=1}^N \rho^{(\lambda)}(\vec{x})$

$\rightarrow U = \frac{1}{2} \sum_{\lambda, \lambda'} \int d\vec{x} d\vec{y} \rho^{(\lambda)}(\vec{x}) \frac{1}{|\vec{x}-\vec{y}|} \rho^{(\lambda')}(\vec{y})$



$$= \sum_k U^{(k)} + \sum_{k \neq l} U^{(k,l)}$$

where 
$$U^{(k)} := \frac{1}{2} \int d\vec{x} d\vec{y} \rho^{(k)}(\vec{x}) \frac{1}{|\vec{x}-\vec{y}|} \rho^{(k)}(\vec{y})$$

"self energy" of the charge distribution

$$U^{(k,l)} := (1 - \delta_{k,l}) \frac{1}{2} \int d\vec{x} d\vec{y} \rho^{(k)}(\vec{x}) \frac{1}{|\vec{x}-\vec{y}|} \rho^{(l)}(\vec{y})$$

"electrostatic interaction" of localized charge densities via a Coulomb interaction

(2) Consider charged point particles:  $\rho^{(k)}(\vec{x}) = e_k \delta(\vec{x} - \vec{x}^{(k)})$

$$\Rightarrow U^{(k,l)} = (1 - \delta_{k,l}) \frac{1}{2} \frac{e_k e_l}{|\vec{x}^{(k)} - \vec{x}^{(l)}|} \quad \text{Coulomb interaction}$$

$U^{(k)}$  does not exist!

(3) The concept of a point charge leads to an infinite self energy and makes no sense within classical electrodynamics. Only the interaction energy of point charges is physically meaningful.

(4) Estimate the smallest spatial extension  $r_0$  of a charge  $e$  with mass  $m$  that still makes sense:

$$e^2/r_0 \approx mc^2 \quad \Rightarrow \quad r_0 = e^2/mc^2$$

For electrons:  $r_0^e = e^2/m_e c^2 = 2.8 \times 10^{-13} \text{ m}$

"classical electron radius"

least experimental upper limit on the radius of the electron:  $r_e < 10^{-20} \text{ m}$  (!)



## 2.6 The law of Biot and Savart

proposition 1: A stationary current density distribution  $\vec{j}(\vec{x})$  leads to a vector potential

$$\vec{A}(\vec{x}) = \frac{1}{c} \int d\vec{y} \frac{\vec{j}(\vec{y})}{|\vec{x} - \vec{y}|}$$

note 19  
inhaltlich 19.

proof: § 1.2  $\rightarrow$  Each component of  $\vec{A}$  obeys Poisson's eq  $\rightarrow$  The solution for each component is given by Poisson's formula  $\square$

remark: (1) § 1.2 proof of prop. 1  $\rightarrow$  This is true in Lorenz gauge,  $\vec{\nabla} \cdot \vec{A} = -\frac{1}{c} \dot{\rho}$

proposition 2: The magnetic field generated by a stationary current density is

$$\vec{B}(\vec{x}) = -\frac{1}{c} \int d\vec{y} \frac{(\vec{x} - \vec{y}) \times \vec{j}(\vec{y})}{|\vec{x} - \vec{y}|^3} \quad \text{law of Biot \& Savart}$$

proof:  $\vec{B}(\vec{x}) = \vec{\nabla} \times \vec{A}(\vec{x})$ , and

$$\left( \vec{\nabla}_x \times \frac{\vec{j}(\vec{y})}{|\vec{x} - \vec{y}|} \right)_i = \epsilon_{ijk} \partial_j \frac{j_k(\vec{y})}{\left( \sum_l (x_l - y_l)^2 \right)^{3/2}} = \epsilon_{ijk} j_k(\vec{y}) \frac{1}{2} \frac{2(x_j - y_j)}{|\vec{x} - \vec{y}|^3}$$

$$= -\epsilon_{ijk} (x_j - y_j) j_k(\vec{y}) \frac{1}{|\vec{x} - \vec{y}|^3} = \frac{((\vec{x} - \vec{y}) \times \vec{j}(\vec{y}))_i}{|\vec{x} - \vec{y}|^3}$$

remark: (2) Notice the analogy between electrostatics and magnetostatics.

(3) See § 2.7 below for a discussion of the concept of a stationary current density.

## 2.7 Magnetostatische Interaktion

Wir betrachten eine unendlich dünne Stromverteilung  $\vec{j}(\vec{x})$ .

Proposition 1: Die Energie des magnetischen Feldes, produziert durch  $\vec{j}(\vec{x})$ ,

$$U = \frac{1}{2c^2} \int d\vec{x} d\vec{y} \vec{j}(\vec{x}) \frac{1}{|\vec{x}-\vec{y}|} \vec{j}(\vec{y})$$

Lemma:  $\vec{\nabla} \cdot (\vec{A} \times \vec{B}) = \vec{B} \cdot (\vec{\nabla} \times \vec{A}) - \vec{A} \cdot (\vec{\nabla} \times \vec{B})$

Proof:  $\partial_i \epsilon_{ijk} A_j B_k = \epsilon_{ijk} (\partial_i A_j) B_k + \epsilon_{ijk} A_j \partial_i B_k$   
 $= B_k \epsilon_{kij} \partial_i A_j - \epsilon_{jik} A_j \partial_i B_k$

Proof of Prop:  $U = \frac{1}{2\mu_0} \int d\vec{x} \vec{B}^2(\vec{x}) = \frac{1}{2\mu_0} \int d\vec{x} \vec{B} \cdot (\vec{\nabla} \times \vec{A})$

Lemma:  $\frac{1}{2\mu_0} \int d\vec{x} \vec{A}(\vec{x}) \cdot (\vec{\nabla} \times \vec{B}(\vec{x})) + \frac{1}{2\mu_0} \int d\vec{x} \vec{\nabla} \cdot (\vec{A}(\vec{x}) \times \vec{B}(\vec{x}))$

$\vec{\nabla} \cdot \vec{B} = \frac{4\pi}{c} \vec{j} \Rightarrow \frac{1}{2\mu_0} \int d\vec{x} \vec{A}(\vec{x}) \cdot \frac{4\pi}{c} \vec{j}(\vec{x}) + \frac{1}{2\mu_0} \int d\vec{y} \cdot \underbrace{(\vec{A}(\vec{x}) \times \vec{B}(\vec{x}))}_{\rightarrow 0 \text{ für } V \rightarrow \infty}$

$$= \frac{1}{2c} \int d\vec{x} \vec{j}(\vec{x}) \cdot \frac{1}{c} \int d\vec{y} \vec{j}(\vec{y}) \frac{1}{|\vec{x}-\vec{y}|}$$

$$= \frac{1}{2c^2} \int d\vec{x} d\vec{y} \vec{j}(\vec{x}) \frac{1}{|\vec{x}-\vec{y}|} \vec{j}(\vec{y}) \quad \square$$

Remark: (1) Für lokalisiert und distribuierte  $\vec{j}(\vec{x}) = \sum_k \vec{j}^{(k)}(\vec{x})$

we agree distinguish between

$U^{(k)} = \frac{1}{2c^2} \int d\vec{x} d\vec{y} \vec{j}^{(k)}(\vec{x}) \frac{1}{|\vec{x}-\vec{y}|} \vec{j}^{(k)}(\vec{y})$  magnetostatische Selbstenergie

and

$U^{(k, \lambda)} = (1 - \delta_{k, \lambda}) \frac{1}{2c^2} \int d\vec{x} d\vec{y} \vec{j}^{(k)}(\vec{x}) \frac{1}{|\vec{x}-\vec{y}|} \vec{j}^{(\lambda)}(\vec{y})$  magnetostatische Interaktion

$\rightarrow U = \sum_k U^{(k)} + \sum_{k \neq \lambda} U^{(k, \lambda)}$

## §3 Multipole expansion for static fields

### 3.1 The electric dipole moment

Q: Given a localized charge distribution  $\rho(\vec{y})$ , what are the potential  $\varphi(\vec{x})$  and the field  $\vec{E}(\vec{x})$ , at a point far from the charges?

Let  $\rho(\vec{y}) \equiv 0$  for  $|\vec{y}| > r_0$ , let  $|\vec{x}| = r \gg r_0$ , and write

$$\frac{1}{|\vec{x} - \vec{y}|} = \frac{1}{\sqrt{r^2 - 2\vec{x} \cdot \vec{y} + y^2}} = \frac{1}{r} \left( 1 - 2 \frac{\vec{x} \cdot \vec{y}}{r^2} + \frac{y^2}{r^2} \right)^{-1/2}$$

$= O(r_0/r) \quad = O(r_0^2/r^2)$

$$= \frac{1}{r} \left( 1 + \frac{\vec{x} \cdot \vec{y}}{r^2} + O(r_0^3/r^3) \right)$$

Potential is found (3.2.3)  $\rightarrow$

$$\varphi(\vec{x}) = \int d\vec{y} \frac{\rho(\vec{y})}{|\vec{x} - \vec{y}|} = \int d\vec{y} \rho(\vec{y}) \frac{1}{r} \left[ 1 + \frac{\vec{x} \cdot \vec{y}}{r^2} + O(r_0^3/r^3) \right]$$

$$= \frac{1}{r} \int d\vec{y} \rho(\vec{y}) + \frac{\vec{x}}{r^2} \cdot \int d\vec{y} \vec{y} \rho(\vec{y}) + O(1/r^3)$$

proposition: For large distances  $r$  from the localized charge distribution the scalar potential has the form

$$\varphi(\vec{x}) = \frac{Q}{r} + \frac{\vec{d} \cdot \vec{x}}{r^2} + O(1/r^3)$$

where  $Q = \int d\vec{y} \rho(\vec{y})$  is the total charge

and  $\vec{d} = \int d\vec{y} \vec{y} \rho(\vec{y})$  is the electric dipole moment

of the charge distribution.



Remark: (1) Analogous results hold for the potential of a localized mass distribution, see PHYS 2655.

If you ever get confused (2) but this, consider a distribution of point charges  $q_i$  at locations  $\vec{x}_i$ :

$$\rho(\vec{y}) = \sum_k q_k \delta(\vec{y} - \vec{x}_k)$$

shifted as means shifted

large locations:  $\vec{x}_k' = \vec{x}_k + \vec{c}$

hence

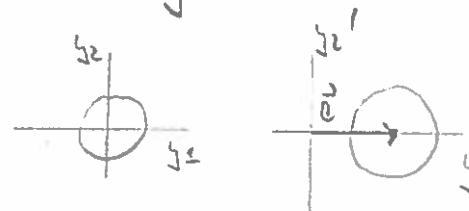
$$\rho'(\vec{y}) = \sum_k q_k \delta(\vec{y} - \vec{x}_k - \vec{c})$$

If  $Q=0$ , then the dipole moment is independent of the origin of the coordinate system: let  $\vec{x}' = \vec{x} + \vec{c}$  will  $\vec{c} = \text{const}$ . Then is the new coordinate system we have

$$\rho'(\vec{y}) = \rho(\vec{y} - \vec{c})$$

$$\rightarrow \vec{d}' = \int d\vec{y} \vec{y} \rho'(\vec{y})$$

$$= \int d\vec{y} \vec{y} \rho(\vec{y} - \vec{c}) = \int d\vec{y} (\vec{y} + \vec{c}) \rho(\vec{y}) = \vec{d} + \vec{c} Q = \vec{d}$$



Woolley: The field at large distances is

$$\vec{E}(\vec{x}) = Q \frac{\vec{x}}{r^3} + \frac{3(\hat{x} \cdot \vec{d}) \hat{x} - \vec{d}}{r^3} + O(1/r^4) \quad \hat{x} = \vec{x}/|\vec{x}|$$

proof:  $\vec{E} = -\vec{\nabla} \phi$  and  $-\vec{\nabla} Q/r = Q \vec{x}/r^3$ , see §2.2

$$\vec{\nabla} \frac{\vec{d} \cdot \vec{x}}{r^3} = \frac{1}{r^3} \vec{\nabla} (\vec{d} \cdot \vec{x}) + (\vec{d} \cdot \vec{x}) \vec{\nabla} \frac{1}{r^3}$$

$$= \frac{1}{r^3} \vec{d} + (\vec{d} \cdot \vec{x}) \left(-\frac{3}{r^2}\right) \frac{1}{r^3} \vec{x} = \frac{\vec{d}}{r^3} - 3 \frac{1}{r^3} (\vec{d} \cdot \vec{x}) \hat{x}$$

Remark: (2) For  $Q=0$ , the leading contribution to the field falls off as  $1/r^3$ .

(4) Obviously, this expansion can be continued, with the next term being the quadrupole moment (a rank-2 tensor, see PHYS 2655). However, it is advantageous to introduce a more general concept.

† ad Problem 20

Week 6

...

Lemma 1: "completeness"  
 Any piecewise continuous and not differentiable fct  
 $f(x) : [-1, 1] \rightarrow \mathbb{R}$  can be expanded in Legendre polynomials.

$$f(x) = \sum_{l=0}^{\infty} f_l P_l(x)$$

where the coefficients are given by

$$f_l = \frac{2l+1}{2} \int_{-1}^1 dx f(x) P_l(x)$$

3.2 Legendre functions, and spherical harmonics (for proofs, see Hall books)

def. 1: The polynomials of degree  $l$  defined by

$$P_l(x) := \frac{1}{2^l l!} \left( \frac{d}{dx} \right)^l (x^2 - 1)^l \quad l = 0, 1, 2, \dots$$

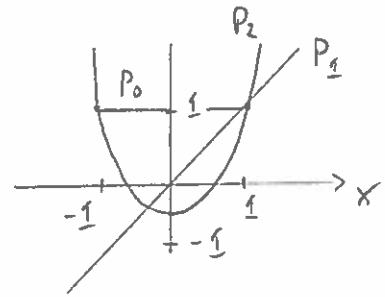
are called Legendre polynomials.

remark: (1) The first few Legendre polynomials are

$$P_0(x) = 1$$

$$P_1(x) = x$$

$$P_2(x) = \frac{1}{2}(3x^2 - 1)$$



(2) The  $P_l(x)$  have the following properties:

- (a)  $P_l(1) = 1 \neq 0$
- (i)  $P_l(-x) = (-1)^l P_l(x)$  parity
- (ii)  $(1-x^2)P_l''(x) - 2xP_l'(x) + l(l+1)P_l(x) = 0$  diff. eq.
- (iii)  $(2l+1)P_{l+1}(x) = (2l+1)xP_l(x) - lP_{l-1}(x)$  recursion relation
- (iv)  $\int_{-1}^1 dx P_l(x)P_{l'}(x) = \delta_{ll'} \frac{2}{2l+1}$  orthogonality

(3) The Legendre polynomials are a member of a more general family called orthogonal polynomials. See, e.g., Szegő's book.

def. 2: The functions

$$P_l^m(x) := \frac{(-1)^m}{2^l l!} (1-x^2)^{m/2} \left( \frac{d}{dx} \right)^{l+m} (x^2 - 1)^l \quad \begin{matrix} l = 0, 1, 2, \dots \\ m = -l, -l+1, \dots, l-1 \end{matrix}$$

are called associated Legendre functions

remark: (4)  $P_l^0(x) = P_l(x)$

(5) For fixed  $l$ , there are  $2l+1$   $P_l^m$ .

2/15/17  
 →

(6) The first few  $P_\ell^m(x)$  are

$$P_0^0(x) = P_0(x) = 1$$

$$P_1^0(x) = P_1(x) = x, \quad P_1^1(x) = -\sqrt{1-x^2}, \quad P_1^{-1}(x) = \frac{1}{2}\sqrt{1-x^2}$$

(7) The  $P_\ell^m$  have the properties

(i)  $P_\ell^m(x=\pm 1) = 0$  for  $m \neq 0$  zeros

(ii)  $P_\ell^{-m}(x) = (-1)^m \frac{(\ell-m)!}{(\ell+m)!} P_\ell^m(x)$  symmetry

(iii)  $\frac{d}{dx} \left[ (1-x^2) \frac{d}{dx} P_\ell^m(x) \right] + \left[ \ell(\ell+1) - \frac{m^2}{1-x^2} \right] P_\ell^m(x) = 0$  ODE

(iv)  $\int_{-1}^1 dx P_\ell^m(x) P_{\ell'}^m(x) = \delta_{\ell\ell'} \frac{2}{2\ell+1} \frac{(\ell+m)!}{(\ell-m)!}$  orthogonality

(v)  $(\ell+1-m) P_{\ell+1}^m(x) = (\ell+1) P_\ell^m(x) - (\ell+m) P_{\ell-1}^m(x)$  recurrence relation

(vi)  $(2\ell+1)\sqrt{1-x^2} P_\ell^m(x) = P_{\ell-1}^{m+1}(x) - P_{\ell+1}^{m+1}(x)$

def. 3: Consider a unit sphere. let  $\mathcal{R} = (\theta, \varphi)$

be a point on the sphere, and let  $\zeta = \cos\theta$

( $-1 \leq \zeta \leq 1$ ). The the  $\ell$ -valued functions defined on the sphere:

$$\underline{Y}_{\ell m}(\mathcal{R}) = \left[ \frac{(2\ell+1)(\ell-m)!}{4\pi(\ell+m)!} \right]^{1/2} e^{im\varphi} P_\ell^m(\zeta)$$

are called spherical harmonics.

(7) Different books define the normalization differently!

remark: (8)  $\underline{Y}_{00}(\mathcal{R}) = \frac{1}{\sqrt{4\pi}}$

$$\underline{Y}_{\ell 0}(\mathcal{R}) = \sqrt{\frac{2\ell+1}{4\pi}} \cos\theta, \quad \underline{Y}_{\ell, \pm\ell}(\mathcal{R}) = \pm \sqrt{\frac{2}{4\pi}} e^{\pm i\varphi} \sin^\ell\theta$$

(9) The  $\underline{Y}_{\ell m}$  have the properties

(i)  $\underline{Y}_{\ell m}^*(\mathcal{R}) = (-1)^m \underline{Y}_{\ell, -m}(\mathcal{R})$  complex conjugation

(ii)  $-i \frac{\partial}{\partial \varphi} \underline{Y}_{\ell m}(\mathcal{R}) = m \underline{Y}_{\ell m}(\mathcal{R})$

ODEs

$$\Delta \underline{Y}_{\ell m}(\mathcal{R}) = -\ell(\ell+1) \underline{Y}_{\ell m}(\mathcal{R})$$

DO!



where  $\Delta = \frac{1}{r^2} \frac{\partial}{\partial r} r^2 \frac{\partial}{\partial r} + \frac{1}{r^2} \frac{\partial^2}{\partial \varphi^2} - \frac{\partial}{\partial r} (1-r^2) \frac{\partial}{\partial r} + \frac{1}{1-r^2} \frac{\partial^2}{\partial \varphi^2}$  is the angular part of the Laplacian operator in spherical coordinates.

$$(iii) \int dR \underline{Y}_{\ell m}^*(R) \underline{Y}_{\ell' m'}(R) = \delta_{\ell\ell'} \delta_{mm'}$$
 orthonormality

Known 2: Any piecewise continuous and well-behaved differentiable fct. on the sphere  $f(R)$  can be expanded in terms of spherical harmonics:

$$f(R) = \sum_{\ell, m} f_{\ell m} \underline{Y}_{\ell m}(R)$$

and the coefficients can give by

$$f_{\ell m} = \int dR f(R) \underline{Y}_{\ell m}^*(R)$$

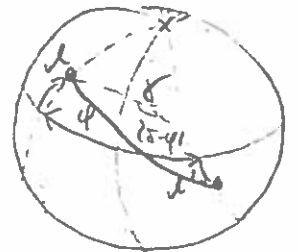
Remark: (10) This is often referred to by saying "the  $\underline{Y}_{\ell m}$  form a complete set on the sphere".

Proposition: Addition theorem

Let  $R = (\vartheta, \varphi)$  and  $R' = (\vartheta', \varphi')$

and let  $\gamma$  be the angle between the two points:

$$\cos \gamma = \cos \vartheta \cos \vartheta' + \sin \vartheta \sin \vartheta' \cos(\varphi - \varphi')$$



Then

$$P_\ell(\cos \gamma) = \frac{4\pi}{2\ell+1} \sum_{m=-\ell}^{\ell} \underline{Y}_{\ell m}^*(R') \underline{Y}_{\ell m}(R)$$

For  $\gamma = 0$  we have  $R = R'$  and  $P_\ell(1) = 1$

$$\Rightarrow \sum_{m=-\ell}^{\ell} |\underline{Y}_{\ell m}(R)|^2 = \frac{2\ell+1}{4\pi}$$
 "sum rule"

Problem 21

spherical polynomials

Problem 22

associated Legendre fcts

Problem 23

spherical harmonics

Warning:



### 3.3 Operation of the Laplace operator in spherical coordinates

Consider the Laplace operator

$$\nabla^2 = \Delta = \frac{1}{r} \frac{\partial^2}{\partial r^2} r + \frac{1}{r^2} \Delta \quad \text{with} \quad \Delta = \frac{1}{\sin \vartheta} \frac{\partial}{\partial \vartheta} \sin \vartheta \frac{\partial}{\partial \vartheta} + \frac{1}{\sin^2 \vartheta} \frac{\partial^2}{\partial \varphi^2}$$

see §3.2 remark (A)

$$\Rightarrow \Delta f(r, \vartheta, \varphi) = \left( \frac{1}{r} \frac{\partial^2}{\partial r^2} r + \frac{1}{r^2} \Delta \right) f(r, \vartheta, \varphi) = \underbrace{\frac{1}{r} \frac{\partial^2}{\partial r^2} r f}_{\text{acts on } r \text{ only}} + \underbrace{\frac{1}{r^2} \Delta f}_{\text{acts on } \vartheta, \varphi \text{ only}}$$

Theorem: The differential eq.

$$\boxed{[-\Delta + V(r)] \psi(r, \vartheta, \varphi) = a(r, \vartheta, \varphi)} \quad (*) \quad \text{for the fct. } \psi(\vec{x})$$

is solved by  $\psi(r, \vartheta, \varphi) = \sum_{l, m} \frac{1}{r} u_{lm}(r) \zeta_{lm}(\vartheta, \varphi)$

where  $u_{lm}(r)$  is the solution of the ODE

$$\boxed{\left( -\frac{d^2}{dr^2} + V_l(r) \right) u_{lm}(r) = r a_{lm}(r)} \quad (**)$$

where  $V_l(r) = V(r) + \frac{l(l+1)}{r^2}$

and  $a_{lm}(r) = \int d\vartheta d\varphi a(r, \vartheta, \varphi) \zeta_{lm}^*(\vartheta, \varphi)$  (\*)

Remark: (1) The Poisson eq. has the form (\*)

(2) This theorem is also very useful in QM.

proof of theorem: ansatz:  $\psi(r, \vartheta, \varphi) = \frac{1}{r} \sum_{lm} u_{lm}(r) \zeta_{lm}(\vartheta, \varphi)$

$$(*) \Rightarrow -\frac{1}{r} \frac{\partial^2}{\partial r^2} r \sum_{lm} u_{lm}(r) \zeta_{lm}(\vartheta, \varphi) - \frac{1}{r^2} \sum_{lm} u_{lm}(r) \Delta \zeta_{lm}(\vartheta, \varphi) + \frac{V(r)}{r} \sum_{lm} u_{lm}(r) \zeta_{lm}(\vartheta, \varphi) = a(r, \vartheta, \varphi)$$

$= -\frac{l(l+1)}{r} \sum_{lm} u_{lm}(r) \zeta_{lm}(\vartheta, \varphi) = a(r, \vartheta, \varphi)$   
 by §3.2 (1a)

response

proof: Consider of 2.7 with  $V(r) = 0$ ,  $c(r, \theta, \varphi) = 0$

$$\rightarrow \partial_r^2 u_{\text{em}}(r) = \frac{l(l+1)}{r^2} u_{\text{em}}(r)$$

ansatz:  $u_{\text{em}}(r) = r^n \rightarrow n(n-1) = l(l+1)$

$$\rightarrow n^2 - n - l(l+1) = 0$$

$$\rightarrow n = \frac{1}{2} (1 \pm \sqrt{1 + 4l(l+1)}) = \frac{1}{2} (1 \pm (2l+1)) = \begin{cases} l+1 \\ -l \end{cases}$$

$\rightarrow$  The two linearly independent solutions are  $\rightarrow$

$$\varphi_{\text{em}}^-(\vec{x}) = \text{const} \times \frac{1}{r} r^{l+1} \underline{Y}_{lm}(\Omega) = \text{const} \times r^l \underline{Y}_{lm}(\Omega)$$

$$\text{and } \varphi_{\text{em}}^+(\vec{x}) = \text{const} \times \frac{1}{r} r^{-l} \underline{Y}_{lm}(\Omega) = \text{const} \times \frac{1}{r^{l+1}} \underline{Y}_{lm}(\Omega)$$

3.2 known  $\rightarrow$  any monochromy will be solved  $\phi(r, R)$  can be expanded in spherical harmonics:  $\phi(r, R) = \sum_{lm} c_{lm}(r) Y_{lm}(R)$  with  $c_{lm}(r)$  given by (+).

$$\rightarrow \sum_{lm} \left[ -\frac{1}{r} \partial_r^2 u_{lm}(r) + \frac{l(l+1)}{r^2} u_{lm}(r) + \frac{V(r)}{r} u_{lm}(r) \right] Y_{lm}(R) = \sum_{lm} c_{lm}(r) Y_{lm}(R)$$

$$\rightarrow \underline{\underline{\left[ -\partial_r^2 + \left( V(r) + \frac{l(l+1)}{r^2} \right) \right] u_{lm}(r) = r c_{lm}(r) \quad \square}}$$

### 3.4 Expansion of harmonic fct in spherical harmonics

besides harmonic fct, i.e., solutions of

$$\boxed{\Delta \phi(\vec{x}) = 0} \quad (*)$$

and even that  $\phi$  is twice continuously differentiable.

proposition: The most general solution of (\*) has the form

$$\boxed{\phi(\vec{x}) = \sum_{lm} \left[ \varphi_{lm}^+(\vec{x}) + \varphi_{lm}^-(\vec{x}) \right]}$$

where  $\varphi_{lm}^+(\vec{x}) = g_{lm}^+ \frac{1}{r^{l+1}} Y_{lm}(R)$

$\varphi_{lm}^-(\vec{x}) = g_{lm}^- r^l Y_{lm}(R)$

with constant coefficients  $g_{lm}^{\pm}$ .

remark: (1)  $\varphi_{lm}^+(\vec{x} \rightarrow 0) \rightarrow \infty \forall l$ ,  $\varphi_{lm}^-(\vec{x} \rightarrow \infty) \rightarrow \infty \forall l > 0$

$\rightarrow$  The only harmonic fct. that is finite at  $r=0$  and at  $r=\infty$  is the constant  $l=0$  contribution, and thus any one that is finite at  $r=0$  and zero at  $r=\infty$  is

### 3.5 Multipole expansion of the electrostatic potential

$$\text{lemma: } \frac{1}{|\vec{x} - \vec{x}'|} = \frac{1}{r_+} \sum_{l=0}^{\infty} \left(\frac{r_-}{r_+}\right)^l \frac{4\pi}{2l+1} \sum_{m=-l}^l Y_{lm}(\Omega) Y_{lm}(\Omega')$$

$$\text{where } \vec{x} = (r, \Omega), \vec{x}' = (r', \Omega'), \quad r_+ = \max(r, r') \\ r_- = \min(r, r')$$

proof: let  $\cos \gamma = \vec{x} \cdot \vec{x}' / rr'$

$$\rightarrow |\vec{x} - \vec{x}'| = \sqrt{r^2 - 2rr' \cos \gamma + r'^2}$$

$$\text{1st con: } r > r' \rightarrow \frac{1}{|\vec{x} - \vec{x}'|} = \frac{1}{r} \frac{1}{\left[1 - 2\frac{r'}{r} \cos \gamma + \left(\frac{r'}{r}\right)^2\right]^{1/2}} = \frac{1}{r} \left[1 - 2\frac{r'}{r} \cos \gamma + \left(\frac{r'}{r}\right)^2\right]^{-1/2}$$

$$\text{§ 3.2 lemma 1} \\ = \frac{1}{r} \sum_{l=0}^{\infty} f_l \left(\frac{r'}{r}\right) P_l(\cos \gamma)$$

$$\text{§ 3.2 prop.} \\ = \frac{1}{r} \sum_{l=0}^{\infty} \frac{4\pi}{2l+1} f_l \left(\frac{r'}{r}\right) \sum_{m=-l}^l Y_{lm}(\Omega') Y_{lm}(\Omega)$$

Remaining question: What is  $f_l(r_-/r_+)$ ?

§ 3.4  $\rightarrow \frac{1}{|\vec{x} - \vec{x}'|}$  is a harmonic fun. for  $r > r'$ , i.e.

$$\Delta_{\vec{x}} \frac{1}{|\vec{x} - \vec{x}'|} = \Delta_{\vec{x}} \frac{1}{r} = \frac{1}{r} \partial_r^2 r = 0$$

Furthermore,  $\frac{1}{|\vec{x} - \vec{x}'|} = O(1/r)$  for  $r \rightarrow \infty$

§ 3.4  $\rightarrow \frac{1}{r} f_l \left(\frac{r'}{r}\right) = \frac{1}{r} \left(\frac{r'}{r}\right)^l c_l$  will work w/ w/out  $c_l$

Put  $\gamma = 0 \rightarrow \frac{1}{|\vec{x} - \vec{x}'|} = \frac{1}{r} \frac{1}{1 - r'/r} = \frac{1}{r} \sum_{l=0}^{\infty} \left(\frac{r'}{r}\right)^l \rightarrow \underline{c_l = 1}$

$$\rightarrow \underline{f_l \left(\frac{r_-}{r_+}\right) = \left(\frac{r_-}{r_+}\right)^l}$$

2nd con:  $r' > r$  analogous

proposition: The electrostatic potential of a localized charge distribution  $\rho(\vec{x})$  ( $\rho(\vec{x})=0$  for  $|\vec{x}|>r_0$ ) can be written, for  $|\vec{x}|>r_0$ ,

$$\varphi(\vec{x}) = \sum_{l,m} \frac{Q_{lm}}{r^{l+1}} \left(\frac{4\pi}{2l+1}\right)^{1/2} Y_{lm}(\theta, \phi)$$

where the  $Q_{lm} = \left(\frac{4\pi}{2l+1}\right)^{1/2} \int_0^\infty dr r^{2+l} \int d\Omega \rho(r, \Omega) Y_{lm}^*(\theta, \phi)$

• on the multipole moments of the charge distribution

proof: § 2.3  $\leadsto \varphi(\vec{x}) = \int d^3y \frac{\rho(\vec{y})}{|\vec{y}-\vec{x}|} \stackrel{|\vec{x}|>r_0}{=} \int d^3y \rho(\vec{y}) \frac{1}{r} \sum_{l=0}^{\infty} \left(\frac{y}{r}\right)^l \frac{4\pi}{2l+1} \sum_{m=-l}^l Y_{lm}(\theta, \phi) Y_{lm}^*(\theta, \phi)$

$$= \sum_{l,m} \frac{1}{r^{l+1}} Y_{lm}(\theta, \phi) \frac{4\pi}{2l+1} \int d^3y y^{2+l} \int d\Omega_y \rho(\vec{y}) Y_{lm}^*(\theta, \phi)$$

remark: (1) The  $l=0$  moment is

$$Q_{00} = \sqrt{\frac{4\pi}{3}} \int_0^\infty dr r^2 \int d\Omega \rho(r, \Omega) \frac{1}{\sqrt{4\pi}} = Q \quad \text{total charge}$$

and the  $l=1$  moments are

$$Q_{1m} = \sqrt{\frac{4\pi}{3}} \int_0^\infty dr r^2 \int d\Omega \rho(r, \Omega) \sqrt{\frac{3}{4\pi}} \left[ \delta_{m,0} \cos\theta - \delta_{m,1} \frac{1}{\sqrt{2}} e^{-i\phi} \sin\theta + \delta_{m,-1} \frac{1}{\sqrt{2}} e^{+i\phi} \sin\theta \right]$$

$$\rightarrow Q_{10} = \int_0^\infty dr r^2 \int d\Omega r \cos\theta \rho(r, \Omega) = \int d^3x x_3 \rho(\vec{x}) = d_3$$

$$Q_{1\pm 1} = \sqrt{\frac{4\pi}{3}} \int_0^\infty dr r^2 \int d\Omega \rho(r, \Omega) r \frac{\sqrt{3}}{2} e^{-i\phi} (-1) (1-\cos^2\theta)$$

$$= \frac{-1}{\sqrt{2}} \int_0^\infty dr r^2 \int d\Omega \rho(r, \Omega) r \sin^2\theta [\cos\phi - i \sin\phi] = \frac{-1}{\sqrt{2}} (d_1 - i d_2)$$

$$Q_{1,-1} = \frac{1}{\sqrt{2}} (d_1 + i d_2) \quad d_1 = \frac{1}{\sqrt{2}} (Q_{1,-1} - Q_{1,1})$$

$$d_2 = \frac{-i}{\sqrt{2}} (Q_{1,-1} + Q_{1,1})$$

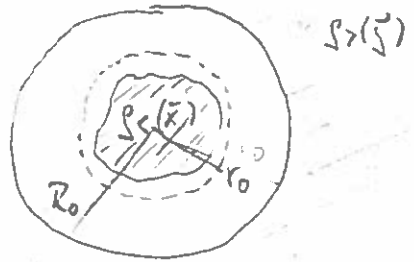
### 3.6 Multipole expansion of the electrostatic interaction

We have a charge density  $\rho_<(\vec{x})$  confined to a region  $R_<$  inside a sphere of radius  $r_0$ . Let  $\rho_<(\vec{x})$

be subject to a field generated by

a charge density  $\rho_>(\vec{y})$  confined to a region  $R_>$  outside a sphere with

radius  $R_0$ .  $\rho_>(\vec{y}) \rightarrow$  The electrostatic interaction energy of this system is



$$\begin{aligned} U &= \frac{1}{2} \int_{R_<} d\vec{x} \rho_<(\vec{x}) \int_{R_>} d\vec{y} \frac{1}{|\vec{x}-\vec{y}|} \rho_>(\vec{y}) + \frac{1}{2} \int_{R_>} d\vec{x} \rho_>(\vec{x}) \int_{R_<} d\vec{y} \frac{1}{|\vec{x}-\vec{y}|} \rho_<(\vec{y}) \\ &= \int_{R_<} d\vec{x} \rho_<(\vec{x}) \int_{R_>} d\vec{y} \frac{1}{|\vec{x}-\vec{y}|} \rho_>(\vec{y}) = \int_{R_<} d\vec{x} \rho_<(\vec{x}) \underline{\underline{\phi_>(\vec{x})}} \end{aligned}$$

where

$$\underline{\underline{\phi_>(\vec{x})}} = \int_{R_>} d\vec{y} \frac{1}{|\vec{x}-\vec{y}|} \rho_>(\vec{y}) \quad \text{is the potential generated by the charges in the region } R_>.$$

If  $R_0 \gg r_0$ ,  $\phi_>(\vec{x})$  will vary slowly within  $R_<$   $\rightarrow$  Taylor expand

$$\phi_>(\vec{x}) = \phi_>(\vec{x}=0) + \vec{x} \cdot \vec{\nabla} \phi_> \Big|_{\vec{x}=0} + \frac{1}{2} x_i x_j \frac{\partial^2}{\partial x_i \partial x_j} \phi_> \Big|_{\vec{x}=0} + \dots$$

(1.1)  $\rightarrow \phi_>(\vec{x})$  obeys Laplace's eq.  $\forall \vec{x} \in R_<$

$$\rightarrow \delta_{ij} \frac{\partial^2}{\partial x_i \partial x_j} \phi_> \Big|_{\vec{x}=0} = 0$$

$$\rightarrow \phi_>(\vec{x}) = \phi_>(\vec{x}=0) + \vec{x} \cdot \vec{\nabla} \phi_> \Big|_{\vec{x}=0} + \frac{1}{2} (x_i x_j - \frac{\vec{x}^2}{3} \delta_{ij}) \frac{\partial^2}{\partial x_i \partial x_j} \phi_> \Big|_{\vec{x}=0} + \dots$$

def. 1: known by

$\phi_0 := \phi_>(\vec{x}=0)$  the potential  $\phi_>$  at the origin  
 $\vec{E} := -\vec{\nabla} \phi_>(\vec{x}=0)$  the field due to  $\phi_>$  at the origin  
 $\dots := \dots$  the field gradients at the origin

$$\rightarrow \varphi_>(\vec{x}) = \varphi_0 - \vec{x} \cdot \vec{E} + \frac{1}{2} (x_i x_j - \frac{\vec{x}^2}{3} \delta_{ij}) \varphi_{ij} + \dots$$

Now drop the  $\langle$  or  $\int \langle$  and the  $\rangle$  or  $\varphi_>$  and will

$$\underline{U} = \int d\vec{x} \rho(\vec{x}) \varphi(\vec{x}) = \varphi_0 \int d\vec{x} \rho(\vec{x}) - \vec{E} \cdot \int d\vec{x} \vec{x} \rho(\vec{x}) + \frac{1}{2} \varphi_{ij} \frac{1}{2} \int d\vec{x} (x_i x_j - \delta_{ij} \frac{\vec{x}^2}{3}) \rho(\vec{x})$$

$$\rightarrow \boxed{U = \varphi_0 Q - \vec{E} \cdot \vec{d} + \frac{1}{2} \varphi_{ij} Q_{ij} + \dots}$$

where  $\varphi_0, \vec{E}, \varphi_{ij}$  are the potential, electric field, and field gradient known due to  $\rho_>$  which is at the origin and  $Q, \vec{d}, Q_{ij}$  are the total charge, dipole moment, and quadrupole moments of  $\rho_<$ .

remark: (1) Alternatively, we can use §3.4 to expand

$$\varphi(\vec{x}) = \sum_{lm} \varphi_{lm}(\vec{x}) = \sum_{lm} q_{lm}^{-} r^l Y_{lm}(\Omega)$$

$$\rightarrow \underline{U} = \int d\vec{x} \rho(\vec{x}) \varphi(\vec{x}) = \int d\vec{x} \sum_{lm} q_{lm}^{-} r^l Y_{lm}(\Omega) \rho(\vec{x})$$

$$= \sum_{lm} q_{lm}^{-} \int_0^\infty dr r^{2+l} \int d\Omega Y_{lm}(\Omega) \rho(\vec{x})$$

$$\stackrel{\S 3.5}{=} \sum_{lm} q_{lm}^{-} \left( \frac{2l+1}{4\pi} \right)^{1/2} Q_{lm}$$

where the  $Q_{lm}$  are the multipole moments of the charge density  $\rho(\vec{x}) \equiv \rho_<(\vec{x})$  and the  $q_{lm}^{-}$  are the coefficients of the expansion of the harmonic potential  $\varphi(\vec{x}) \equiv \varphi_>(\vec{x})$  in spherical harmonics.

2/22/17

Wohl 25  
add due to  
Mittel Georges

Wohl 26  
Lager Sphäroid  
in Parallelepiped

Week 7

WS 17 (19, 20, 21, 22)

## 3.7 The magnetic moment

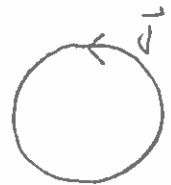
§2.6  $\rightarrow$  The law of Biot & Savart gives the magnetic field resulting from a stationary current distribution. This requires an interpretation as currents are produced by moving charges and hence are intrinsically time dependent.

def 1: By a stationary current density  $\vec{j}(\vec{x})$  we mean the time average taken over a time  $T$  that is very long compared to all microscopic time scales:

$$\vec{j}(\vec{x}) = \overline{\vec{j}(\vec{x}, t)} \equiv \frac{1}{T} \int_0^T dt \vec{j}(\vec{x}, t)$$

example: (1) Current in a wire loop.

$T \gg$  time it takes an electron to complete one revolution.



remark: (1) With this definition the 4<sup>th</sup> Maxwell eq. reduces to its static version upon time averaging, provided the electric field  $\vec{E}$  as a function of time is bounded:

$$\overline{\partial \vec{E}(\vec{x}, t) / \partial t} = \frac{1}{T} \int_0^T dt \frac{\partial \vec{E}}{\partial t} = \frac{1}{T} [\vec{E}(\vec{x}, T) - \vec{E}(\vec{x}, 0)] \xrightarrow{T \rightarrow \infty} 0$$

if  $\vec{E}(\vec{x}, t)$  is bounded.

$$\rightarrow \overline{-\frac{1}{c} \partial_t \vec{E} + \vec{\nabla} \times \vec{J}} = \boxed{\vec{\nabla} \times \vec{J} = \frac{4\pi}{c} \vec{j}}$$

Now consider the vector potential  $\vec{A}(\vec{x}) \equiv \overline{\vec{A}(\vec{x}, t)}$  at large distance from a localized static current density:  $\vec{j}(\vec{y}) = \sum_{\alpha} e_{\alpha} \vec{v}_{\alpha} \delta(\vec{y} - \vec{x}_{\alpha})$



$$\S 2.6 \leadsto \underline{\underline{\vec{A}(\vec{x})}} = \frac{1}{c} \int d\vec{J} \frac{\vec{J}(\vec{J})}{|\vec{x}-\vec{J}|} = \frac{1}{c} \sum_k \frac{e_k \vec{v}_k}{|\vec{x}-\vec{x}_k|}$$

$$\underline{\underline{\vec{A}(\vec{x})}} = \frac{1}{c} \sum_k \overline{e_k \vec{v}_k} \frac{1}{r} \left[ 1 + \frac{\vec{x} \cdot \vec{x}_k}{r^2} + \dots \right]$$

$$\sum_k \overline{e_k \vec{v}_k} = \frac{d}{dt} \sum_k e_k \vec{x}_k = 0 \text{ by remark (1)}$$

$$= \frac{1}{c} \frac{1}{r^3} \sum_k \overline{e_k \vec{v}_k (\vec{x}_k \cdot \vec{x})}$$

Wieder

$$\underline{\underline{\sum_k e_k \vec{v}_k (\vec{x}_k \cdot \vec{x})}} = \sum_k e_k \dot{\vec{x}}_k (\vec{x}_k \cdot \vec{x}) =$$

$$= \frac{1}{c} \frac{d}{dt} \sum_k e_k \vec{x}_k (\vec{x}_k \cdot \vec{x}) + \frac{1}{c} \sum_k e_k (\vec{v}_k (\vec{x}_k \cdot \vec{x}) - \vec{x}_k (\vec{v}_k \cdot \vec{x}))$$

time  
average  $\rightarrow$

$$0 + \frac{1}{c} \sum_k \overline{e_k (\vec{v}_k (\vec{x}_k \cdot \vec{x}) - \vec{x}_k (\vec{v}_k \cdot \vec{x}))}$$

$$\rightarrow \underline{\underline{\vec{A}(\vec{x})}} = \frac{1}{2c} \frac{1}{r^3} \sum_k \overline{e_k (\vec{v}_k (\vec{x}_k \cdot \vec{x}) - \vec{x}_k (\vec{v}_k \cdot \vec{x}))}$$

def. 2: The magnetic moment of the charges is defined as

$$\vec{m} := \frac{1}{2c} \sum_k \overline{e_k (\vec{x}_k \times \vec{v}_k)} = \frac{1}{2c} \int d\vec{x} \overline{\vec{x} \times \vec{J}(\vec{x})}$$

proposition: The vector potential for large distances is given by the magnetic moment via

$$\underline{\underline{\vec{A}(\vec{x})}} = \frac{1}{r^3} \vec{m} \times \vec{x}$$

proof:  $\vec{m} \times \vec{x} = \frac{1}{2c} \sum_k \overline{e_k (\vec{x}_k \times \vec{v}_k) \times \vec{x}} = \frac{1}{2c} \sum_k \overline{e_k (\vec{v}_k (\vec{x}_k \cdot \vec{x}) - \vec{x}_k (\vec{v}_k \cdot \vec{x}))}$

Wolley: The magnetic field for large distances is

$$\vec{A}(\vec{x}) = \frac{J(\hat{x} \cdot \vec{m}) \hat{x} - \vec{m}}{r^3} + O(1/r^4)$$

proof:  $\vec{A}_i = (\nabla \times \vec{A})_i = \epsilon_{ijk} \partial_j \frac{1}{r^3} \epsilon_{klm} m_l x_m = \partial_j \frac{1}{r^3} (m_i x_j - m_j x_i) = \frac{1}{r^3} (3m_i - m_i) \partial_j x_j + \dots$

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Proposition 2: If all of the moving charges have the same charge-to-mass ratio  $e/m \equiv e/m$ , and if the motion is nonrelativistic,  $v \ll c$ , then the magnetic moment is proportional to the angular momentum  $\vec{L}$  of the moving charges.

$$\vec{m} = \frac{e}{2mc} \vec{L} \quad (*)$$

proof:  $\vec{L} = \sum_k \vec{x}_k \times \vec{p}_k = \sum_k m_k \vec{x}_k \times \vec{v}_k$

$$\vec{m} = \frac{1}{2c} \sum_k e_k (\vec{x}_k \times \vec{v}_k) = \frac{1}{2c} \sum_k \frac{e_k}{m_k} m_k (\vec{x}_k \times \vec{v}_k) = \frac{e}{2mc} \vec{L}$$

Remark: (1) The proportionality factor  $\frac{e}{2mc}$  is called gyromagnetic ratio.

(2) (\*) holds for the orbital angular momentum  $\vec{L}$  of classical particles, but not for the magnetic moment related to the spin of quantum particles. For electrons,

$$\vec{m}_e = g \frac{e}{2mc} \vec{S}_e$$

with  $\vec{S}_e = \frac{1}{2} \hbar$  the spin of the electron and  $g = 2.0023 \dots$  the g-factor.

(4) The Dirac eq. yields  $g=2$ ;  $g \neq 2$  is due to loop corrections.

P50 f.p.

Note (my eyes only):

This is a very tricky and confusing point. Jackson discusses it, but in a disjointed way. In pp 186, 216 he has book for the magnetostatic case, and pp 142, 161 for the electrostatic case. LL II discuss the electrostatic case in §42, but they refer to discuss the magnetostatic analogy in Vol II! They do discuss the problem in Vol VIII, where they point out that the roles of the thermodynamic potentials they call  $\tilde{F}$  and  $\tilde{F}$ , respectively, is reversed in the magnetic case compared to the electric case. See Vol. VIII §§31, 32, and especially the remark at the end of §30 in my German edition.

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## 1.8 The magnetostatic energy of a unit distribution

Consider the magnetostatic energy of §1.6, i.e., a collection of localized unit densities  $j_j(\vec{r})$  (let on "ferromagnet") and another one,  $j_k(\vec{r})$ , (let on "ion").

§1.7  $\rightarrow$  the energy of the magnetic fields (let omit for them units is

$$\underline{U} = \frac{1}{c^2} \int d\vec{x} d\vec{r} \vec{j}_k(\vec{x}) \cdot \int d\vec{r} \frac{1}{|\vec{x}-\vec{r}|} \vec{j}_j(\vec{r})$$

$$\underline{U} \stackrel{1.6}{=} \frac{1}{c} \int d\vec{x} \vec{j}_k(\vec{x}) \cdot \vec{A}_j(\vec{x}) \quad \text{with } \vec{A}_j \text{ the potential generated by the units } \vec{j}_j.$$

Taylor expand  $\vec{A}_j(\vec{x})$  in energy to §1.6:

$$A_j^i(\vec{x}) = A_j^i(\vec{x}=0) + x_j \partial_j A_j^i(\vec{x}) \Big|_{\vec{x}=0} + \dots$$

$$\rightarrow \underline{U} = \frac{1}{c} \int d\vec{x} j_k^i(\vec{x}) A_j^i(\vec{x}=0) + \frac{1}{c} \int d\vec{x} j_k^i(\vec{x}) x_j \partial_j A_j^i(\vec{x}) \Big|_{\vec{x}=0} + \dots$$

$$\int d\vec{x} \vec{j}_k(\vec{x}) = \int d\vec{x} \sum_k e_k \vec{v}_k \delta(\vec{x}-\vec{x}_k) = \sum_k e_k \vec{v}_k = 0 \quad \text{by } \S 1.7$$

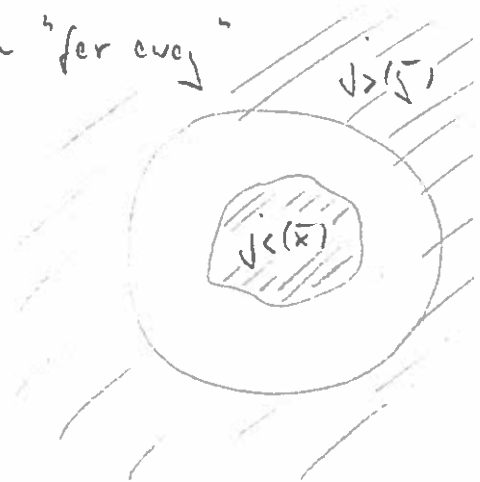
$$= \frac{1}{c} \int d\vec{x} \sum_k e_k v_k^i \delta(\vec{x}-\vec{x}_k) x_j \partial_j A_j^i(\vec{x}) \Big|_{\vec{x}=0} + \dots$$

$$= \frac{1}{c} \sum_k e_k v_k^i x_k^j \left( \partial_j A_j^i \right) \Big|_{\vec{x}=0} =$$

$$\approx \frac{1}{c^2} \sum_k e_k \left[ v_k^i x_k^j \left( \partial_j A_j^i \right) \Big|_{\vec{x}=0} - x_k^i v_k^j \left( \partial_j A_j^i \right) \Big|_{\vec{x}=0} \right]$$

$$= \frac{1}{c^2} \left( \partial_j A_j^i \right) \Big|_{\vec{x}=0} \sum_k e_k (v_k^i x_k^j - v_k^j x_k^i)$$

same eqn as in §1.7, p 48



Now consider

$$\begin{aligned} \underline{\vec{A}} \cdot \underline{\vec{m}} &= (\underline{\nabla} \times \underline{\vec{A}}) \cdot \underline{\vec{m}} = \epsilon_{ijk} \partial_j A_k \frac{1}{c} \sum_x e_x \epsilon_{ilm} x_l v_m \\ &= \frac{1}{c} (\partial_j A_k) \sum_x e_x (\delta_{je} \delta_{km} - \delta_{jm} \delta_{ke}) x_l v_l \\ &= \frac{1}{c} (\partial_j A_k) \sum_x e_x (x_k^j v_k^l - v_k^j x_k^l) \end{aligned}$$

$$\rightarrow \boxed{U = \underline{\vec{A}} \cdot \underline{\vec{m}} + (\text{quadrupole term})}$$

where  $\underline{\vec{A}}$  is the field due to  $\underline{\vec{j}}$ , evaluated at the origin  
 $\underline{\vec{m}}$  is the magnetic moment of the  $\underline{\vec{j}}$

remark: (1) This has the opposite sign of the dipole term in the corresponding electrostatic expansion, § 2.6!

(2) This is not the energy of a magnetic dipole with fixed moment  $\underline{\vec{m}}$  in an external field  $\underline{\vec{A}}$  (which we know is  $-\underline{\vec{m}} \cdot \underline{\vec{A}}$ ). Rather, it is the energy of the total field configuration resulting from the  $\underline{\vec{j}}$  and  $\underline{\vec{j}}_e$ , which includes the work that was done to get the currents  $\underline{\vec{j}}$  flowing. That is, the time dependent processes hidden in the time average necessary to see talk about a static limit during leads to a fundamental difference between the magnetic static energy and the electrostatic one.

† see also § 2.9

## 2.9 The energy of dipoles in external fields

To find the energy of a fixed magnetic dipole (e.g., a electron spin) in a magnetic field, consider the force exerted by a field  $\vec{B}(\vec{x})$  (produced by  $\vec{j}$ ) on a unit distribution  $\vec{j}$  ( $=\vec{j}^c$ )

cf. §§ 2.5c, 2.5  $\rightarrow$  the magnetic or Lorentz force on a single charge is  $\frac{e}{c} \vec{v} \times \vec{B}$

$\rightarrow$  the force on  $\vec{j}(\vec{x}) = \sum_k e_k \vec{v}_k \delta(\vec{x}_k - \vec{x})$  is

$$\vec{F}_{\text{mag}} = \frac{1}{c} \int d\vec{x} \vec{j}(\vec{x}) \times \vec{B}(\vec{x}) \stackrel{\text{Taylor}}{=} \frac{1}{c} \int d\vec{x} \vec{j}(\vec{x}) \times \left[ \vec{B}(\vec{x}=0) + (\vec{x} \cdot \nabla) \vec{B}(\vec{x}) + \dots \right]_{\vec{x}=0}$$

line: (1)  $\int d\vec{x} \vec{j}(\vec{x}) = 0$

(2)  $\int d\vec{x} (x_i j_j(\vec{x}) + x_j j_i(\vec{x})) = 0$

(3)  $\int d\vec{x} (\vec{a} \cdot \vec{x}) \vec{j}(\vec{x}) = -\frac{1}{2} \vec{a} \times \int d\vec{x} (\vec{x} \times \vec{j}(\vec{x}))$  with  $\vec{a} = \text{const}$

proof:  $\vec{j}(\vec{x}) = \sum_k e_k \vec{v}_k \delta(\vec{x} - \vec{x}_k)$

$\rightarrow$  (1)  $\int d\vec{x} \vec{j}(\vec{x}) = \sum_k e_k \vec{v}_k = \frac{d}{dt} \sum_k e_k \vec{x}_k = 0$  by § 2.7 mod. (1)

(2)  $\int d\vec{x} (x_i j_j + x_j j_i) = \sum_k e_k (x_k^i v_k^j + x_k^j v_k^i) = \frac{d}{dt} \sum_k e_k x_k^i x_k^j = 0$

by the same argument

(3)  $\int d\vec{x} (\vec{a} \cdot \vec{x}) \vec{j}(\vec{x}) = \int d\vec{x} a_j x_j j_i \stackrel{(1)}{=} \frac{1}{2} \int d\vec{x} a_j (x_j j_i - x_i j_j)$

$= -\frac{1}{2} \epsilon_{ijk} a_j \int d\vec{x} \epsilon_{klm} x_l j_m = -\frac{1}{2} (\vec{a} \times \int d\vec{x} (\vec{x} \times \vec{j}))_i$

$\rightarrow \vec{F}_{\text{mag}}^i = -\frac{1}{c} \left( \vec{B}(\vec{x}=0) \times \int d\vec{x} \vec{j}(\vec{x}) \right)_i + \frac{1}{c} \epsilon_{ijk} \int d\vec{x} j_j(\vec{x}) \left( \nabla_{\vec{x}=0} \vec{B}_k \right)_i \cdot \vec{x} + \dots$

$= 0$  by line (1)

line (2)  
 $= -\frac{1}{2c} \epsilon_{ijk} \left( \left( \nabla_{\vec{x}=0} \vec{B}_k \right) \times \int d\vec{x} (\vec{x} \times \vec{j}(\vec{x})) \right)_j + \dots$

§ 2.7 def 2

$$\underline{\underline{F}}_{\text{mag}} = \epsilon_{ijk} \left( (\nabla \vec{A})_{\vec{x}=0} \times \vec{m} \right)_j + \dots = \epsilon_{ijk} \left( (\vec{m} \times \nabla)_{\vec{x}=0} \vec{A}_k \right)_j + \dots$$

$$\begin{aligned} \rightarrow \underline{\underline{F}}_{\text{mag}} &= (\vec{m} \times \nabla)_{\vec{x}=0} \times \vec{A} + \dots = \nabla (\vec{m} \cdot \vec{A}) - \underbrace{\vec{m} (\nabla \cdot \vec{A})}_{=0} + \dots = \nabla (\vec{m} \cdot \vec{A}(\vec{x}))_{\vec{x}=0} \\ &= \underline{\underline{\nabla}} u \quad \text{with } u \text{ per § 2.8} \end{aligned}$$

That  $\underline{\underline{F}}$  is the gradient of minus the desired potential energy

$$\rightarrow \underline{\underline{F}}_{\text{mag}} = -\underline{\underline{\nabla}} \tilde{u} \quad \text{remark: (0) then we interpret } \tilde{u} \text{ as } \vec{x}\text{-dependent via } \vec{A}: \\ \tilde{u}(\vec{x}) = -\vec{m} \cdot \vec{A}(\vec{x}) \text{ at pt } \vec{x}=0 \text{ after taking the gradient.}$$

when  $\tilde{u} = -u = -\vec{m} \cdot \vec{A} + \dots$  is the potential energy of the magnetic dipole  $\vec{m}$  in the magnetic field  $\vec{A}$

remark: (1) In a QM context, with  $\vec{m}$  the magnetic moment of a spin, this is often referred to as the Zeeman energy

(2) For the corresponding electrostatic problem, the force exerted by an electric field  $\vec{E}$  on a charge distribution  $\rho(\vec{x})$  is given by

$$\underline{\underline{F}}_{\text{el}} = -\underline{\underline{\nabla}} u \quad \text{with } u \text{ per § 2.6, see Problem 27}$$

This is a fundamental difference between electrostatics and magnetostatics!

### Problem 27

Electric charges in an external field

### Week 8

Problem 8 (23, 24, 25)