

Problem Assignment # 14

04/29/2021 due
05/10/2021**4.3.4. Classical atom**

Consider a classical electron in a circular orbit in a Coulomb potential, for which the virial theorem yields $\bar{V} = 2E$.

- a) Assuming that the electron continues to move on a circle as it radiates, calculate the average power radiated as a function of E .

hint: Note that the power is not a linear function of E , in contrast to the preceding two problems!

- b) Show that the electron reaches the nucleus in a finite amount of time. For a hydrogen atom, calculate that time if the initial orbit had a radius $r_0 = 10^{-8}$ cm.

(5 points)

4.3.5. Absence of dipole radiation

Show that a system of particles that all have the same ratio of charge to mass and are not subject to any external forces cannot emit either electric or magnetic dipole radiation.

(3 points)

4.3.6. Rotating dipole

An electric dipole moment \mathbf{d} rotates uniformly with angular velocity Ω in a plane. Find the radiated power per solid angle, and the total radiated power, averaged over one rotational period.

(5 points)

4.3.1.) a)

Wektor pol. Ld.: $V = -\frac{\alpha}{r}$

form $\vec{F} = \frac{-\alpha}{r^2} \hat{r} = m v$

$$\rightarrow \ddot{v} = \frac{-\alpha}{mr^2} \hat{r}$$

$$\rightarrow \underline{(\dot{v})^2} = \frac{\alpha^2}{m^2 r^4} = \frac{1}{2^2 m^2} (V(r))^4$$

hier wahr + valid nur + circular orbit ($\rightarrow \overline{V^4} = \bar{V}^4$)

$$\rightarrow \overline{(\dot{v})^2} = \frac{1}{2^2 m^2} \overline{V^4} = \frac{1}{2^2 m^2} 16 E^4$$

$$\rightarrow \underline{\overline{P}} = \frac{2e^2}{Jc^2} (\dot{v})^2 = \frac{2e^2}{Jc^2} \frac{16}{2^2 m^2} E^4 = \frac{32e^2}{Jc^2} \frac{1}{2^2 m^2} E^4$$

b) $\overline{P} = - \frac{d\bar{E}}{dt} = \lambda E^4 \quad \text{mit } \lambda = \frac{32e^2}{Jc^2} \frac{1}{2^2 m^2}$

$$\rightarrow \frac{d\bar{E}}{E^4} = -\lambda dt$$

$$\rightarrow -\frac{1}{3} \left(\frac{1}{\bar{E}^3} - \frac{1}{E_0^3} \right) = -\lambda t \quad \text{mit } \bar{E}_0 = E(t=0)$$

$$\rightarrow \frac{1}{\bar{E}^3} = \frac{1}{E_0^3} + 3\lambda t$$

$$\rightarrow \underline{\underline{\frac{E(t)}{(1+3\lambda E_0^3 t)^{1/3}}}} \quad \text{mit } \underline{\underline{E_0 < 0}}$$

\rightarrow Perihel mit km unter (i.e., $E = -\infty$) et dann

$$t^* = \frac{-1}{3E_0^3 \lambda}$$

Int $E_0 = \frac{1}{2} V_0 = -\frac{\alpha}{2r_0}$ will r_0 the stoch₁ radius

$$\Rightarrow \underline{\underline{t^*}} = \frac{8r_0^3}{3\alpha^2} \frac{3e^3}{32\epsilon^2} \frac{m^2}{\kappa} r_0^2 = \frac{c^3}{4e^2} \frac{m^2}{\kappa} r_0^3$$

Kondis: $\alpha = e^2$ $r_0 = 10^{-8} \text{ cm}$

$$\Rightarrow \underline{\underline{t^*}} = \frac{c^3}{4e^4} m^2 r_0^3 = \frac{(1 \times 10^{10})^3}{4 \times (4.8 \times 10^{-10})^4} (9.1 \times 10^{-28})^2 \times 10^{-24} \text{ s}$$

$$= 0.35 \times 10^{-10} \text{ s}$$

4.3.5.) Consider a set of charges e_k with mass m_k .

The electric dipole moment is

$$\vec{d} = \sum_k e_k \vec{x}_k = \sum_k \frac{e_k}{m_k} m_k \vec{x}_k$$

\rightarrow If $e_k/m_k = \text{const}$, then

(1) $\vec{d} = \text{const} \sum_k m_k \vec{x}_k = \text{const. } \vec{X}(t)$

will $\vec{X}(t)$ be center of mass. If $\vec{X}(t)$ moves

uniformly $\rightarrow \ddot{\vec{d}}(t) \propto \ddot{\vec{X}}(t) = 0 \rightarrow \underline{\text{no electric dipole radiation}}$

The magnetic dipole moment is

$$\begin{aligned} \vec{m} &= \frac{1}{2c} \sum_k e_k \vec{x}_k \times \vec{v}_k = \frac{1}{2c} \sum_k \frac{e_k}{m_k} \vec{x}_k \times m_k \vec{v}_k \\ &= \text{const.} \times \frac{1}{2c} \sum_k \vec{x}_k \times \vec{p}_k = \text{const.} \times \frac{1}{2c} \vec{L}(t) \end{aligned}$$

will $\vec{L} = \sum_k \vec{x}_k \times \vec{p}_k = \sum_k \vec{x}_k \times m_k \vec{v}_k$ be total angular momentum

If angular momentum is constant $\rightarrow \dot{\vec{L}}(t) = 0$

$\rightarrow \ddot{\vec{m}}(t) = 0 \rightarrow \underline{\text{no magnetic dipole radiation}}$

4.3.6.) Let \vec{d} lie in the x-y-plane.

$$\rightarrow d_x = d \omega s R t \quad \text{with } d = |\vec{d}|.$$

$$d_y = d \omega c R t$$

$$\rightarrow \ddot{\vec{d}}(t) = R^2 \vec{d}(t)$$

Let ϑ be the angle between the direction \hat{x} of the observer and the z-axis $\rightarrow (\hat{x} \times \vec{d})^2 = R^2 (\hat{x} \times \vec{d})^2$

Wieder $\hat{x} \times \vec{d}(t) = \begin{pmatrix} r \sin \vartheta \omega \\ r \sin \vartheta \omega \\ r \cos \vartheta \end{pmatrix} \times \begin{pmatrix} d_x(t) \\ d_y(t) \\ 0 \end{pmatrix} = \begin{pmatrix} -r \omega s \vartheta d_y(t) \\ r \omega s \vartheta d_x(t) \\ x d_y(t) - y d_x(t) \end{pmatrix}$

$$\begin{aligned} \rightarrow (\hat{x} \times \vec{d}(t))^2 &= r^2 \omega^2 d^2 \vec{d}^2 + x^2 d_y^2(t) + y^2 d_x^2(t) - 2xy d_x(t) d_y(t) \\ &= r^2 \omega^2 d^2 \vec{d}^2 + r^2 \sin^2 \vartheta \omega^2 d^2 \omega^2 R^2 t \\ &\quad + r^2 \sin^2 \vartheta \omega^2 d^2 \omega^2 R^2 t \\ &\quad - 2r^2 \sin^2 \vartheta \omega^2 d^2 \omega^2 R^2 t \sin \vartheta \cos \vartheta \end{aligned}$$

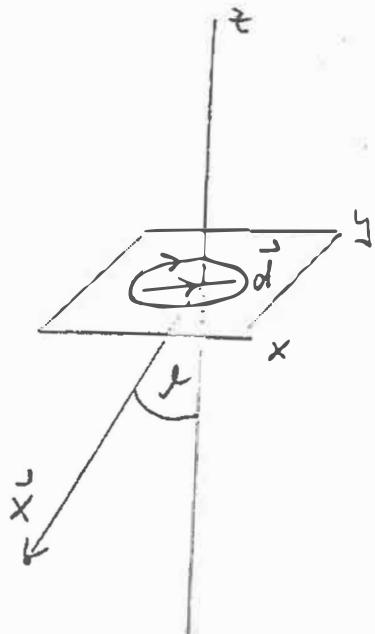
Average over one rotation period: $\bar{T} = 2\pi/\omega$:

$$\frac{\omega^2 R t}{\bar{T}} = \frac{1}{T} \int_0^T dt \omega^2 \frac{2\pi}{T} t = \frac{1}{2\pi} \int_0^{2\pi} dx \omega^2 x = \frac{1}{2}$$

$$\frac{\omega^2 R t}{\bar{T}} = \frac{1}{2}$$

$$\frac{\omega R t \omega R t}{\bar{T}} = 0$$

$$\begin{aligned} \rightarrow \frac{(\hat{x} \times \vec{d}(t))^2}{\bar{T}} &= \left(r^2 \omega^2 d^2 + \frac{1}{2} r^2 \omega^2 d^2 \right) d^2 = r^2 \left(\omega^2 d^2 + \frac{1}{2} - \frac{1}{2} \omega^2 d^2 \right) \\ &= \frac{1}{2} r^2 (1 + \omega^2 d^2) \end{aligned}$$



\rightarrow The power per wavelength, averaged over one rotation period, is

$$\begin{aligned} \overline{\frac{dP}{dR}} &= \frac{1}{4\pi c^3} \overline{(\hat{x} \cdot \vec{d}(t))^2} = \frac{R^4}{4\pi c^3} \overline{(\hat{x} \cdot \vec{d}(t))^2} \\ (1) \quad &= \frac{R^4 d^2}{4\pi c^3} \frac{1}{2} (1 + \cos^2 \vartheta) = \underline{\underline{\frac{R^4 d^2}{8\pi c^3} (1 + \cos^2 \vartheta)}} \end{aligned}$$

\rightarrow The total radiated power, averaged over one period, is

$$\begin{aligned} \underline{\underline{P}} &= \int dR \overline{\frac{dP}{dR}} = 2\pi \int_0^R d(R) (1 + \cos^2 \vartheta) \frac{R^4 d^2}{8\pi c^3} \\ (1) \quad &= \frac{R^4 d^2}{4c^3} 2 \left(1 + \frac{1}{2}\right) = \underline{\underline{\frac{3d^2}{32c^3} R^4}} \end{aligned}$$