

4.4.1. Radiation by an accelerated point particle

Show that the result of ch.4 §4.4 example (2) for the energy radiated by an accelerated point particle,

$$\frac{dU}{d\omega} = \frac{2}{3} \frac{e^2}{\pi c^3} |\dot{\mathbf{v}}(\omega)|^2$$

is consistent with the Larmor formula for the radiated power, ch.4 §3.3:

$$\mathcal{P}(t) = \frac{2e^2}{3c^3} (\dot{\mathbf{v}}(t))^2$$

(2 points)

4.4.2. Classical model of an atom

Consider the classical model of a radiating atom from ch.4 §4.5; i.e., a damped harmonic oscillator with charge e and equation of motion

$$\ddot{y} = -\omega_0^2 y - \gamma \dot{y}$$

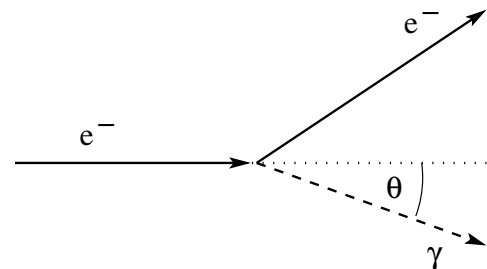
Do NOT assume the damping to be small and solve the equation of motion exactly.

- Assuming that the oscillator represents an electron, and that the spectrum is peaked somewhere in the range of visible light, use the approximate result from ch.4 §4.5 to show that one does not need to consider the case of the oscillator being overdamped.
- Calculate again the radiation spectrum $dU/d\omega$, and the total energy U radiated, exactly. Compare with the approximate result obtained in class. Show that the result self-consistently validates the assumption made in part a) above.

(8 points)

4.5.1. Čerenkov radiation

- Consider a particle with rest mass m_0 that travels through a medium with index of refraction n . The particle emits a photon that moves at an angle θ with respect to the particle's initial trajectory. Use relativistic kinematics to express θ in terms of the particle's initial energy and momentum, the wave number of the photon, and n .
- Show that to leading order for $c \rightarrow \infty$ one recovers the result of the nonrelativistic approximation used in ch. 4 §5.2.
- The core of a <https://www.youtube.com/watch?v=mgNwteP-6M> water-cooled reactor emits a 1.42 MeV electron that travels through the water ($n = 1.33$). Calculate θ for blue light ($\lambda = 4,000 \text{ \AA}$). How good or bad is the nonrelativistic approximation in this case?



(6 points)

4.4.1.) The Larmor formula gives for the total radiated energy

$$\underline{U = \int dt P(t) = \frac{2e^2}{3c^3} \int dt (\ddot{\vec{r}}(t))^2} \quad (*)$$

whence by 4.4.1. implies

$$\begin{aligned} \underline{U} &= \int_0^\infty du \, du/du = \int_0^\infty du \frac{2e^2}{3c^3} |\ddot{\vec{r}}(u)|^2 \\ &= \frac{2e^2}{3c^3} \frac{1}{\sigma} \int_0^\infty du |\ddot{\vec{r}}(u)|^2 \quad (**) \end{aligned}$$

①

Now let $f(t) \in \mathbb{R}$ be a fct. of time, and

$$f(u) = \int dt e^{iut} f(t) \quad \text{its Fourier transform.}$$

Then

$$\underline{\int_0^\infty du |f(u)|^2} = \int_0^\infty du f(u) f^*(u)$$

$$= \frac{1}{2} \int_{-\infty}^{\infty} du \int dt e^{iut} f(t) \int dt' e^{-iut'} f(t')$$

$$= \frac{1}{2} \int dt dt' f(t) f(t') \underbrace{\int du e^{iu(t-t')}}_{= 2\pi \delta(t-t')}$$

$$= \underline{\pi \int dt (f(t))^2}$$

①

$$\rightarrow \underline{(*) = (**)}$$

4.4.2.1 a) Eq. of motion: $\ddot{y} + \omega_0^2 y + \gamma \dot{y} = 0$

ansatz: $y(t) = e^{rt} \rightarrow r^2 + \gamma r + \omega_0^2 = 0$

$$\rightarrow r = \frac{1}{2} (-\gamma \pm \sqrt{\gamma^2 - 4\omega_0^2})$$

1st case: $\gamma < 2\omega_0$ underdamped

$$r = \pm i\omega_0 \sqrt{1 - \gamma^2/4\omega_0^2} - \frac{1}{2}\gamma =: \pm i\tilde{\omega}_0 - \frac{1}{2}\gamma$$

$$\rightarrow y(t) = c_1 e^{i\tilde{\omega}_0 t - \frac{1}{2}\gamma t} + c_2 e^{-i\tilde{\omega}_0 t - \frac{1}{2}\gamma t}$$

$$y(t=0) = c_1 + c_2 \stackrel{!}{=} a$$

$$\begin{aligned} \dot{y}(t=0) &= c_1 (i\tilde{\omega}_0 - \frac{1}{2}\gamma) + (-i\tilde{\omega}_0 - \frac{1}{2}\gamma) c_2 \\ &= -\frac{\gamma}{2} a + (c_1 - c_2) i\tilde{\omega}_0 \stackrel{!}{=} 0 \end{aligned}$$

$$\begin{aligned} \rightarrow c_2 &= c_1 - \frac{\gamma}{2} a \frac{1}{i\tilde{\omega}_0} = c_1 + i a \frac{\gamma}{2\tilde{\omega}_0} \\ &= a - c_1 \end{aligned}$$

$$\rightarrow c_1 = \frac{a}{2} - i \frac{a}{4} \frac{\gamma}{\tilde{\omega}_0}, \quad c_2 = \frac{a}{2} + i \frac{a}{4} \frac{\gamma}{\tilde{\omega}_0}$$

$$\rightarrow \underline{y(t) = a \left[\cos \tilde{\omega}_0 t + \frac{\gamma}{2\tilde{\omega}_0} \sin \tilde{\omega}_0 t \right] e^{-\gamma t/2}}$$

$$\begin{aligned} \underline{\underline{\dot{y}(t)}} &= a \left[-\tilde{\omega}_0 \sin \tilde{\omega}_0 t + \frac{\gamma}{2\tilde{\omega}_0} \tilde{\omega}_0 \cos \tilde{\omega}_0 t \right] e^{-\gamma t/2} \\ &\quad - a \frac{1}{2} \gamma \left[\cos \tilde{\omega}_0 t + \frac{\gamma}{2\tilde{\omega}_0} \sin \tilde{\omega}_0 t \right] e^{-\gamma t/2} \\ &= -a \tilde{\omega}_0 \left[1 + \left(\frac{\gamma}{2\tilde{\omega}_0} \right)^2 \right] \sin \tilde{\omega}_0 t e^{-\gamma t/2} \end{aligned}$$

wh $\underline{\underline{\tilde{\omega}_0 = \omega_0 \sqrt{1 - (\gamma/2\omega_0)^2}}}$

(1)

(1)

2nd case: $\gamma = 2\omega_0$ critically damped

3rd case: $\gamma > 2\omega_0$ overdamped

d.5 §4.5 \rightarrow the final answer (to be confirmed self-consistently below) is $\gamma = \frac{e^2 \omega_0^2}{mc^2}$

$$\begin{aligned} \rightarrow \frac{\gamma}{\omega_0} &\approx \frac{e^2 \omega_0 / c}{mc^2} = \frac{2\pi e^2 / \lambda}{mc^2} && \text{with } \lambda \text{ the wavelength of} \\ & && \text{the radiation} \\ &\approx \frac{e^2 / c_0}{mc^2} \cdot \frac{2\pi c_0}{\lambda} && \text{with } c_0 \text{ the Bohr radius} \\ &\approx \frac{10 \text{ eV}}{5 \times 10^5 \text{ eV}} \cdot \frac{10 \text{ \AA}}{5,000 \text{ \AA}} \approx 4 \times 10^{-8} && \text{for electrons and} \\ & && \text{visible light} \end{aligned}$$

\rightarrow We always have the underdamped case, except for oscillators with a very large frequency ω_0 .

\rightarrow Consider the underdamped case only.

b) Now we have $v(t) = \dot{y}(t) = a \tilde{\omega}_0 \sin \tilde{\omega}_0 t e^{-\gamma t/2}$

with $\tilde{\omega}_0 = \tilde{\omega}_0 [1 + (\gamma/2\tilde{\omega}_0)^2]$

This yields (see d.5 §4.5)

$$v(\omega) = \frac{e \tilde{\omega}_0}{2} \left[\frac{1}{\omega - \tilde{\omega}_0 + i\gamma/2} - \frac{1}{\omega + \tilde{\omega}_0 + i\gamma/2} \right]$$

p-4.4.2-3

$$\begin{aligned} \Rightarrow \frac{du}{d\omega} &= \frac{ze^2}{35c^2} |\dot{v}(0)|^2 = \frac{ze^2 \tilde{\omega}_0^2 \omega^2}{65c^2} \left[\frac{1}{(\omega - \tilde{\omega}_0 + i\gamma/2)} - \frac{1}{\omega + \tilde{\omega}_0 + i\gamma/2} \right] \\ &\quad \times \left[\frac{1}{\omega - \tilde{\omega}_0 - i\gamma/2} - \frac{1}{\omega + \tilde{\omega}_0 - i\gamma/2} \right] \\ &= \frac{ze^2 \tilde{\omega}_0^2 \omega^2}{65c^2} \left[\frac{1}{(\omega - \tilde{\omega}_0)^2 + \gamma^2/4} + \frac{1}{(\omega + \tilde{\omega}_0)^2 + \gamma^2/4} - 2i\gamma \frac{1}{(\omega - \tilde{\omega}_0 + i\gamma/2)(\omega + \tilde{\omega}_0 - i\gamma/2)} \right] \\ &= \frac{ze^2 \tilde{\omega}_0^2 \omega^2}{65c^2} \left[\frac{1}{(\omega - \tilde{\omega}_0)^2 + \gamma^2/4} + \frac{1}{(\omega + \tilde{\omega}_0)^2 + \gamma^2/4} - 2i\gamma \frac{\omega^2 - \tilde{\omega}_0^2 + \gamma^2/4}{[(\omega - \tilde{\omega}_0)^2 + \gamma^2/4][(\omega + \tilde{\omega}_0)^2 + \gamma^2/4]} \right] \\ &= \frac{ze^2 \tilde{\omega}_0^2 \omega^2}{65c^2} \frac{2\omega^2 + 2\tilde{\omega}_0^2 + \gamma^2/2 - 2i\gamma(\omega^2 - \tilde{\omega}_0^2 - \gamma^2/4)}{[(\omega - \tilde{\omega}_0)^2 + \gamma^2/4][(\omega + \tilde{\omega}_0)^2 + \gamma^2/4]} \\ &= \frac{2}{35} \frac{ze^2 \tilde{\omega}_0^2 \omega^2}{c^2} \frac{\omega^2}{[(\omega - \tilde{\omega}_0)^2 + \gamma^2/4][(\omega + \tilde{\omega}_0)^2 + \gamma^2/4]} \end{aligned}$$

$$\begin{aligned} \Rightarrow u &= \int d\omega \frac{du}{d\omega} = \\ &= \frac{2}{35} \frac{ze^2 \tilde{\omega}_0^2 \omega^2}{c^2} \frac{1}{\tilde{\omega}_0} \int dx \frac{x^2}{[(x-1)^2 + \gamma^2/4\tilde{\omega}_0^2][(x+1)^2 + \gamma^2/4\tilde{\omega}_0^2]} \\ &= \frac{2}{35} \frac{ze^2 \tilde{\omega}_0^2 \omega^2}{c^2} \frac{\sigma}{\tilde{\omega}_0} \frac{2\tilde{\omega}_0}{\gamma} = \frac{2e^2 \tilde{\omega}_0^2 \omega^2}{35c^2 \gamma} \\ \frac{\tilde{\omega}_0 \omega^2}{\tilde{\omega}_0 \omega^2} &= \tilde{\omega}_0^4 [1 + (\gamma/2\tilde{\omega}_0)^2] = \omega_0^4 [1 - (\gamma/2\omega_0)^2]^2 [1 + \frac{(\gamma/2\omega_0)^2}{1 - (\gamma/2\omega_0)^2}] = \omega_0^4 \\ &= \frac{2e^2 c^2 \omega_0^4}{35c^2 \gamma} \end{aligned}$$

Compare with d 4 § 4.5 \Rightarrow The effects of the approximations in the mode is that calculation is correct!

$\Rightarrow \gamma/\omega_0$ is indeed as small as estimated in part a).

This justifies both using the underdamped case and the approximations made in d 5 § 4.5

4.5.1.7c) particle: rest mass: m_0

initial momentum, energy: \vec{p}_0, E_0

final momentum, energy: \vec{p}_f, E_f

energy-momentum relation:

$$E = \sqrt{m_0^2 c^4 + p^2 c^2}$$

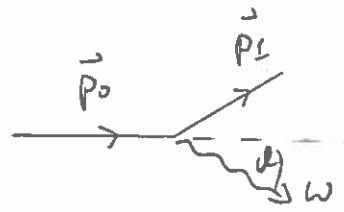
photon: frequency ω wave vector \vec{k}

momentum: $\vec{p}_\gamma = \hbar \vec{k}$, $p_\gamma = |\vec{p}_\gamma| = \hbar k = \hbar \omega / c_{\text{eff}}$

energy-momentum relation: $E_\gamma = \hbar \omega = c_{\text{eff}} p_\gamma$

medium: index of refraction n

speed of light: $c_{\text{eff}} = c/n$



energy conservation: $E_0 = E_f + \hbar \omega$ (1)

momentum conservation: $\vec{p}_0 = \vec{p}_f + \hbar \vec{k}$ (2)

(2) $\rightarrow p_0^2 + \hbar^2 k^2 - 2 p_0 \hbar k \cos \theta = p_f^2 = E_f^2 / c^2 - m_0^2 c^2$ (3)

(1) $\rightarrow E_f = c \sqrt{m_0^2 c^2 + p_0^2} - c_{\text{eff}} \hbar k$

$\rightarrow E_f^2 = c^2 (m_0^2 c^2 + p_0^2) + \hbar^2 k^2 c_{\text{eff}}^2 - 2 c_{\text{eff}} \hbar k E_0$ (4)

(4) \sim (3) $\rightarrow p_0^2 + \hbar^2 k^2 - 2 p_0 \hbar k \cos \theta = m_0^2 c^2 + p_0^2 + \hbar^2 k^2 \frac{c_{\text{eff}}^2}{c^2} - 2 \frac{\hbar k c_{\text{eff}}}{c^2} E_0 - m_0^2 c^2$

$\rightarrow 2 p_0 \hbar k \cos \theta = \hbar^2 k^2 \left(-\frac{c_{\text{eff}}^2}{c^2} + 1 \right) + 2 \hbar k \frac{c_{\text{eff}}}{c^2} E_0$

$\rightarrow 2 p_0 \cos \theta = \hbar k (1 - 1/n^2) + 2 E_0 / c n$

$\rightarrow \boxed{\cos \theta = \frac{E_0}{c p_0 n} + \frac{\hbar k}{p_0} \frac{n^2 - 1}{2 n^2}}$

$$b) \quad p_0 = mv_0 = m_0 v_0 \gamma \quad \text{with } \gamma = \frac{1}{\sqrt{1-v_0^2/c^2}}$$

$$E_0 = mc^2 = m_0 c^2 \gamma$$

$$\rightarrow \frac{wsd}{cp_0} = \frac{m_0 c^2 \gamma}{(m_0 v_0 \gamma) c} = 0(v^0) = \frac{c}{v_0} [1 + O(v/c)]$$

$$\textcircled{1} \quad d \approx 5.2 \checkmark$$

$$c) \quad E_0 = 1.42 \text{ MeV} = \sqrt{m_0^2 c^4 + p_0^2 c^2} \quad \frac{m_0 c^2 = 0.51 \text{ MeV}}{\text{for electrons}}$$

$$\rightarrow p_0^2 = \frac{1}{c^2} (E_0^2 - m_0^2 c^4)$$

$$\rightarrow cp_0 = \sqrt{1.42^2 - 0.51^2} \text{ MeV} = 1.33 \text{ MeV}$$

$$\rightarrow \frac{E_0}{cp_0} = \frac{1.42}{1.33} \cdot \frac{1}{1.33} = 0.81$$

$$\frac{t \lambda}{p_0} = \frac{2\pi}{\lambda} \frac{t}{p_0} = \frac{2\pi}{4 \times 10^2 \times 10^{-2}} \frac{10^{-27} \times 3 \times 10^{10}}{1.33 \times 10^6 \times 1.6 \times 10^{-12}} \approx 10^{-5} \times \frac{1}{4 \times 1.33 \times 1.6} = 10^{-6}$$

① \rightarrow The term of $O(v/c)$ is a negligible correction!

$$\rightarrow \frac{wsd}{cp_0} = 0.81$$

$$\underline{\underline{d = 36^\circ}}$$

①