

## History and Applications of Calculus Homework 4

1. A basic physics fact is that “work is force times distance.” This means that to pull something along with constant force  $F$  Newtons for a total distance of  $x$  meters takes  $Fx$  Joules of energy... In this question, you’ll also need Newton’s Law of Gravity: the gravitational force between mass  $M$  and mass  $m$  which are a distance  $r$  apart is  $F = GMm/r^2$  where  $G$  is the universal gravitational constant.

- (a) The constant  $G$  is roughly  $6.674 \times 10^{-11} \text{ m}^2 / \text{kg} \cdot \text{sec}^2$ . Also the radius of the earth is about 4000 miles. What is the mass of the earth?
- (b) Suppose you have a mass  $m$  distance  $r$  from another mass  $M$ . How much work do you need to do in order to move the mass  $m$  very very far away from the mass  $M$  (“to infinity”)?
- (c) In the lectures this week, we showed that the eccentricity of a planetary orbit around the sun satisfies the formula  $e = \left(\frac{v_0}{v_{\text{crit}}}\right)^2 - 1$  where  $v_{\text{crit}} = \sqrt{GM}/\sqrt{r_0}$ ,  $v_0$  is the initial velocity of a planet of mass  $m$  moving in the tangential direction,  $r_0$  is its initial distance from the sun, and  $M$  is the mass of the sun. How much kinetic energy does the planet have when  $v_0 = \sqrt{2}v_{\text{crit}}$ ?
- (d) Explain the connection between parts (b) and (c).

2. Recall that a function  $f(x)$  with domain  $\mathbb{R}$  is called *even* if  $f(x) = f(-x)$  and *odd* if  $f(x) = -f(-x)$  for all  $x$ .

- (a) Suppose that  $f(x)$  is *any* function with domain  $\mathbb{R}$ . Prove that there exist unique functions  $E(x)$  and  $O(x)$  such that  $f(x) = E(x) + O(x)$  for every  $x$ , with  $E(x)$  being even and  $O(x)$  being odd. These functions  $E(x)$  and  $O(x)$  are called the *even* and *odd* parts of  $f(x)$ , respectively.
- (b) Suppose that  $f(x) = e^x$  is the exponential function. Sketch the graphs of  $f(x)$ , and its even and odd parts  $E(x)$  and  $O(x)$ , all on the same axes.

3. The functions  $E(x)$  and  $O(x)$  in 2(b) are called the *hyperbolic functions*  $\cosh x$  and  $\sinh x$ , respectively. Explicitly:

$$\cosh x = \frac{1}{2}(e^x + e^{-x}), \quad \sinh x = \frac{1}{2}(e^x - e^{-x}).$$

- (a) Show that the derivative of  $\sinh x$  is  $\cosh x$  and the derivative of  $\cosh x$  is  $\sinh x$ .
- (b) Show that  $\cosh^2 x - \sinh^2 x = 1$ .
- (c) Let  $\tanh x := \frac{\sinh x}{\cosh x}$  and  $\operatorname{sech} x := \frac{1}{\cosh x}$ . Prove that the derivative of  $\tanh x$  is  $\operatorname{sech}^2 x$ .

4. Consider the differential equation

$$f''(x) - f(x) = 0.$$

- (a) Prove that the functions  $\cosh x$  and  $\sinh x$  both give solutions of this differential equation, hence, so does any linear combination of the form  $a \cosh x + b \sinh x$  for  $a, b \in \mathbb{R}$ . In fact, the latter is the *general solution* of this differential equation; you may assume this without proof in the remainder of the question.
- (b) Prove that  $\cosh(x + y) = \cosh x \cosh y + \sinh x \sinh y$ . In particular,  $\cosh 2x = \cosh^2 x + \sinh^2 x = 2 \cosh^2 x - 1 = 1 + 2 \sinh^2 x$ .
- (c) What is the analogous “double angle formula” for  $\sinh 2x$ ?

5. In the lectures this week, we used laws of motion to derive the differential equation

$$\frac{d^2 u}{d\theta^2} + u = \frac{f(r)r^2}{(v_0)^2(r_0)^2 m}$$

where  $u = \frac{1}{r}$  and  $f(r)$  was the central force,  $m$  was the mass of the planet, and  $v_0$  and  $r_0$  came from the initial conditions we were assuming (which implied  $u = \frac{1}{r_0}$  and  $\frac{du}{d\theta} = 0$  when  $\theta = 0$ ). Then we assumed the central force was given by Newton’s inverse square law  $f(r) = \frac{GMm}{r^2}$ , and solved this equation to obtain the final equation of motion; it was an ellipse, parabola or hyperbola according to the initial velocity  $v_0$ . I suggest you review this before attempting this question!

Taking the same general setup and initial conditions, suppose instead that we live in a parallel universe in which the central force satisfies the *inverse cube law*

$$f(r) = \frac{GMm}{r^3}$$

for some alien constant  $G$  (it even has different units compared to our  $G$ !). What can you say about orbits of planets in this universe?