

Chapter 1

Maxwell's equations

§1 The variational principle of classical electrodynamics

1.1 The Minkowski metric

exer 1: Spacetime can be described by a 4-dimensional Minkowski spacetime with $x^\mu = (ct, \vec{x})$, where t is time, $\vec{x} = (x^1, x^2, x^3)$ is position in real space, and c is a characteristic velocity.

remark: (1) We will choose the metric to be $g_{\mu\nu} = (+, -, -, -)$. We will use Greek indices, $\mu, \nu, \dots = 0, 1, 2, 3$, to denote coordinates in \mathbb{R}^4 , and Latin indices, $i, j, \dots = 1, 2, 3$, to denote coordinates in the 3-d Euclidean subspace.

exer 2: Empty space ("vacuum") supports a Minkowski-vector field $A^\mu(x)$ called the electromagnetic 4-vector potential

def. 1: The rank-2 antisymmetric field $F^{\mu\nu}$ constructed from gradients of A^μ as

$$F^{\mu\nu}(x) = \partial^\mu A^\nu(x) - \partial^\nu A^\mu(x)$$

is called electromagnetic field tensor

remark: (2) ~~60021~~ $F^{\mu\nu}(x)$ can be represented in terms of a \vec{E} , a Euclidean vector field and a Euclidean pseudovector field

exer 1: (c) The physical field configurations $A^\mu(x)$ we seek are those that minimize the action

$$S_{\text{vec}} = -\frac{1}{16\pi} \int d^4x F_{\mu\nu}(x) F^{\mu\nu}(x) \quad \text{with } \int d^4x \equiv c \int dt \int d^3x$$

- remark: (1) EdH is governed by a principle of least action, as is classical mechanics. However, in EdH we need to find field configurations $A^\mu(x)$; in mechanics we only need to find paths $\vec{x}(t)$.
- (2) $A^\mu(x)$ enters the action only in conjunction with gradients!
- (3) $F_{\mu\nu}F^{\mu\nu}$ is a Lorentz scalar \Rightarrow the theory is invariant under Lorentz boosts, but not under Galilean boosts!
- (4) Maxwell did not formulate the theory like us. Advantages of his formulation: Manifest Lorentz invariance, easy to generalize, analogous to other field theories (GR, particle physics).
- (5) Maxwell theory was the first unified field theory.

exer 1: (b) Matter is described by (among other things) a 4-vector $J^\mu(x)$ that couples to $A^\mu(x)$ via

$$S_{\text{int}} = -\frac{1}{c} \int d^4x J_\mu(x) A^\mu(x) \quad J^\mu \text{ is called } \underline{\text{4-current vectors}}$$

(7) The factor $1/c$ is conventional

remark: (6) This does not include the feedback from the field on the matter. For that, one needs another piece of the action that governs the matter.

def. 2: The dual field tensor (in physics name) is defined as

$$\hat{F}^{\mu\nu} := \epsilon^{\mu\nu\lambda\sigma} F_{\lambda\sigma} \quad \text{with } \epsilon^{\mu\nu\lambda\sigma} \text{ the 4-d Levi-Civita tensor}$$

with 1
and field tensor

$$\partial_\mu \hat{F}^{\mu\nu}(x) = 0 \quad (*) \quad \text{proof: Problem 1.1}$$

Remark: (*) The structure of $\hat{F}^{\mu\nu}$ is that of gradients of A^μ potentials (*).

1.2 Euler-Lagrange equations for fields

Recall classical mechanics:

Lagrangian: $L(q_1(t), q_2(t), \dots, q_f(t); \dot{q}_1(t), \dot{q}_2(t), \dots, \dot{q}_f(t))$

action: $S' = \int dt L(q(t), \dot{q}(t))$

extremals: $0 \stackrel{!}{=} \delta S' = \int dt \left[\sum_i \frac{\partial L}{\partial q_i} \delta q_i + \sum_i \frac{\partial L}{\partial \dot{q}_i} \delta \dot{q}_i \right]$

$\stackrel{\text{part. int.}}{=} \int dt \sum_i \left[\frac{\partial L}{\partial q_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} \right] \delta q_i \quad \forall \delta q_i(t)$

$\rightarrow \frac{\partial L}{\partial q_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} = 0$ by the fundamental lemma

Now consider field theory: discretize space \rightarrow a field

$$\phi(x) = \phi(\vec{x}, t) = q_i$$

can be considered a system with $f = \infty$ if we identify $\phi(\vec{x}_1, t) \equiv q_1(t)$, $\phi(\vec{x}_2, t) \equiv q_2(t)$, etc. The Lagrangian now becomes a

Lagrangian density: $\mathcal{L}(\phi(\vec{x}, t), \partial^\mu \phi(\vec{x}, t))$ that depends on spatial gradients in addition to time derivatives, etc

Lagrangian: $L = \int d\vec{x} \mathcal{L}(\phi(\vec{x}, t), \partial^\mu \phi(\vec{x}, t))$ is the spatial integral over \mathcal{L} .

discussion: derive the GL eqs, Problem 4, the way we derived (+):

$$S = S[\phi, \phi^\dagger, \vec{\nabla}\phi, \vec{\nabla}\phi^\dagger, \vec{A}, \partial_i t_j]$$

Extremize:

$$\begin{aligned} 0 &\stackrel{!}{=} \delta S = \int d\vec{x} \left(\frac{\partial \mathcal{L}}{\partial \phi^\dagger} \delta \phi^\dagger + \frac{\partial \mathcal{L}}{\partial \vec{\nabla} \phi^\dagger} \cdot \delta \vec{\nabla} \phi^\dagger + \text{c.c.} + \frac{\partial \mathcal{L}}{\partial \vec{A}} \delta \vec{A} + \frac{\partial \mathcal{L}}{\partial \partial_i t_j} \delta (\partial_i t_j) \right) \\ &= \int d\vec{x} \left(\left(\frac{\partial \mathcal{L}}{\partial \phi^\dagger} - \vec{\nabla} \cdot \frac{\partial \mathcal{L}}{\partial \vec{\nabla} \phi^\dagger} \right) \delta \phi^\dagger + \text{c.c.} + \int d\vec{x} \left(\frac{\partial \mathcal{L}}{\partial A_j} - \partial_i \frac{\partial \mathcal{L}}{\partial (\partial_i t_j)} \right) \delta t_j \right) \end{aligned}$$

\rightarrow One condition is

$$\boxed{\frac{\partial \mathcal{L}}{\partial \phi^\dagger} - \vec{\nabla} \cdot \frac{\partial \mathcal{L}}{\partial \vec{\nabla} \phi^\dagger} = 0} \quad (+)$$

Now, $\mathcal{L}_{GL} = r \phi \phi^\dagger + c [(\vec{\nabla} - iq\vec{A})\phi][(\vec{\nabla} + iq\vec{A})\phi^\dagger] + \lambda \phi^2 \phi^{\dagger 2} + \dots$ has λc don't depend on \vec{A}

$$\begin{aligned} &= r \phi \phi^\dagger + c (\vec{\nabla} \phi) (\vec{\nabla} \phi^\dagger) + c (\vec{\nabla} \phi) \cdot iq\vec{A} \phi^\dagger - c iq\vec{A} \phi \cdot \vec{\nabla} \phi^\dagger \\ &\quad + c q^2 \vec{A}^2 \phi \phi^\dagger + \lambda \phi^2 \phi^{\dagger 2} + \dots \end{aligned}$$

$$\rightarrow \frac{\partial \mathcal{L}}{\partial \phi^\dagger} = r \phi + c iq\vec{A} \cdot \vec{\nabla} \phi + c q^2 \vec{A}^2 \phi + 2\lambda \phi^2 \phi^\dagger$$

$$\frac{\partial \mathcal{L}}{\partial (\vec{\nabla} \phi^\dagger)} = c \vec{\nabla} \phi - c iq \vec{A} \phi$$

\rightarrow (+) reads explicitly

$$\begin{aligned} \underline{0} &= (r + 2\lambda |\phi|^2) \phi - c (\vec{\nabla} \phi - iq \vec{\nabla} \cdot (\vec{A} \phi)) - c iq \vec{A} \cdot \vec{\nabla} \phi - q^2 \vec{A}^2 \phi \\ &= (r + 2\lambda |\phi|^2) \phi - c [\vec{\nabla} - iq \vec{A}]^2 \phi \end{aligned}$$

which is the same eq. we obtain from $\frac{\delta S}{\delta \phi^\dagger(\vec{x})} \stackrel{!}{=} 0$, the way we obtain of Problem 4. The other GL eq. follows analogously.

The

action: $S = c \int dt L = \int dx^0 \int d^3x \mathcal{L}(\phi(x), \partial_\mu \phi(x)) = \int d^4x \mathcal{L}(\phi(x), \partial_\mu \phi(x))$

is defined as for $f < \infty$.

extremals: $0 \stackrel{!}{=} \delta S = \int d^4x \left[\frac{\partial \mathcal{L}}{\partial \phi} \delta \phi + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \delta (\partial_\mu \phi) \right]$
 $= \int d^4x \left[\frac{\partial \mathcal{L}}{\partial \phi} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right] \delta \phi + \delta \phi$

$$\Rightarrow \boxed{\partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} = \frac{\partial \mathcal{L}}{\partial \phi}} \quad (*)$$

remark: (1) (*) is the Euler-Lagrange equation for the scalar field ϕ . (This is the fast way of deriving it. In Problem 2 for a somewhat more detailed procedure.)

(2) Generalization to tensor fields is straightforward just add more (discrete) indices.

(3) In general, \mathcal{L} will depend on higher gradients. §1.1 axiom (2) postulates that the Maxwell \mathcal{L} has two gradients of $F^{\mu\nu}$, or two gradients of A^μ .

(4) The EL eqs for fields on PDEs, in contrast to the case $f < \infty$, often being on coupled ODEs!

1.3 The field equations

§1.1 \leadsto the Lagrangian density of Maxwell theory is

$$\mathcal{L}(A^\mu(x)) = \frac{-1}{16\pi} F_{\mu\nu}(x) F^{\mu\nu}(x) - \frac{1}{c} A_\mu(x) j^\mu(x)$$

§1.2 \leadsto The Euler-Lagrange equation reads

$$\partial_\lambda \frac{\partial \mathcal{L}}{\partial (\partial_\lambda A_\mu(x))} = \frac{\partial \mathcal{L}}{\partial A_\mu(x)}$$

In \mathcal{L} , the first term depends only on gradients of $A^\mu(x)$, and the second term is gradient-free.

Problem 2 partial derivative

Problem 3 Mini-words of

Problem 4 OLEPs

1/11/17

1/11/16

$$\begin{aligned}
\rightarrow \frac{\partial}{\partial(\partial_\lambda A_\kappa(x))} \mathcal{L} &= \frac{-1}{16\pi} \frac{\partial}{\partial(\partial_\lambda A_\kappa)} F^{\mu\nu} \int F^{\mu\nu} \int F_{12} \\
&= \frac{-1}{16\pi} \frac{\partial}{\partial(\partial_\lambda A_\kappa)} (\partial_\mu A_\nu - \partial_\nu A_\mu) (\partial_{12} A_\lambda - \partial_\lambda A_{12}) \int F^{\mu\nu} \int \\
&= \frac{-1}{16\pi} \int F^{\mu\nu} \int [(\delta^\lambda_\mu \delta^\kappa_\nu - \delta^\lambda_\nu \delta^\kappa_\mu) (\partial_{12} A_\lambda - \partial_\lambda A_{12}) \\
&\quad + (\delta^\lambda_{12} \delta^\kappa_\lambda - \delta^\lambda_\lambda \delta^\kappa_{12}) (\partial_\mu A_\nu - \partial_\nu A_\mu)] \\
&= \frac{-1}{16\pi} \left[\int A_{12} \int - \int A_{12} \int \right] (\partial_{12} A_\lambda - \partial_\lambda A_{12}) + \left[\int F^{\mu\nu} \int - \int F^{\mu\nu} \int \right] (\partial_\mu A_\nu - \partial_\nu A_\mu) \\
&= \frac{-1}{16\pi} \left[\partial^\lambda A^\kappa - \partial^\kappa A^\lambda - \partial^\kappa A^\lambda + \partial^\lambda A^\kappa + \partial^\lambda A^\kappa - \partial^\kappa A^\lambda - \partial^\kappa A^\lambda + \partial^\lambda A^\kappa \right] \\
&= \frac{-1}{4\pi} (\partial^\kappa A^\lambda - \partial^\lambda A^\kappa) = \frac{1}{4\pi} F^{\kappa\lambda}
\end{aligned}$$

$$\text{and } \frac{\partial}{\partial A_\kappa(x)} \mathcal{L} = \frac{\partial}{\partial A_\kappa} \left[\frac{1}{c} A_\mu \dot{\gamma}^\mu - \frac{1}{c} \delta^\kappa_\mu \dot{\gamma}^\mu \right] = -\frac{1}{c} \dot{\gamma}^\kappa$$

$$\rightarrow \text{The EL-eg. reads } \frac{1}{4\pi} \partial_\lambda F^{\lambda\kappa}(x) = -\frac{1}{c} \dot{\gamma}^\kappa(x)$$

$$\text{or } \boxed{\partial_\mu F^{\mu\nu}(x) = \frac{4\pi}{c} \dot{\gamma}^\nu(x)} \quad (*)$$

remark: (1) All physical field configurations must obey (*).

(2) $F^{\mu\nu}$ is defined in terms of A^μ , so (*) should be considered a differential eq. for $A^\mu(x)$. Alternatively we can interpret (*) by the proportionality for §1.1:

$$\boxed{\partial_\mu \varepsilon^{\mu\nu\alpha\beta} F_{\alpha\beta}(x) = 0} \quad (***)$$

which contains the structure of $F^{\mu\nu}$ in terms of gradients of A^μ . We can then consider (*) + (***) the field eqs for $F^{\mu\nu}(x)$, which is now the fundamental field.

[12] Conservation laws, ed gauge invariance

2.1 Continuity equation for the 4-current

Proposition: The 4-current vector obeys the continuity eq.

$$\boxed{\partial_\mu j^\mu(x) = 0} \quad (*)$$

Proof: $\int 1.3 (2) \rightarrow \partial_\nu j^\nu = \frac{c}{4\pi} \partial_\nu \partial_\mu F^{\mu\nu} = -\frac{c}{4\pi} \partial_\nu \partial_\mu F^{\nu\mu} = 0$

sym. antisym.

Remark: (1) The 4-vector $j^\mu = (c\rho, \vec{j})$ has a time-like component $j^0 = c\rho$ ed three space-like components $j^i = j^i$. ρ is called electric charge density ed \vec{j} is called electric current density.

(2) In terms of ρ ed \vec{j} , (*) reads $\partial_0 c\rho + \partial_i j^i = 0$
 that $\partial_0 = \frac{\partial}{\partial x^0} = \frac{1}{c} \partial_t$, ed $\partial_i = \frac{\partial}{\partial x^i} = \vec{\nabla}$ (w.r.t. of \vec{x})

$$\rightarrow \boxed{\partial_t \rho(\vec{x}, t) + \vec{\nabla} \cdot \vec{j}(\vec{x}, t) = 0} \quad (**)$$

is equivalent to (*).

(3) Integrate (**) over a spatial volume V with boundary surface

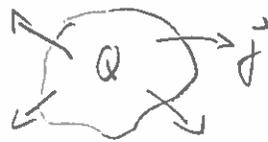
$$\partial_t \int_V d\vec{x} \rho(\vec{x}, t) = - \int_V d\vec{x} \vec{\nabla} \cdot \vec{j}(\vec{x}, t) = - \int_{(\partial V)} d\vec{s} \cdot \vec{j}(\vec{x}, t)$$

define $Q := \int_V d\vec{x} \rho(\vec{x}, t)$ to be the total charge within

$$\rightarrow \frac{d}{dt} Q = - \int_{(\partial V)} d\vec{s} \cdot \vec{j}(\vec{x}, t)$$

\rightarrow Q can change only via current

flowing in or out of the volume V , hence the name "continuity eq."



Week 1

Problem 1
(1, 2, 3, 4)

2.2 The energy-momentum tensor

def.: The tensor field

$$T^{\mu\nu}(x) := \frac{-1}{4\pi} F^{\mu\alpha}(x) F^{\nu}_{\alpha}(x) + \frac{1}{16\pi} \int^{\mu\nu} F_{\lambda\alpha}(x) F^{\lambda\alpha}(x) \quad (*)$$

is called the electromagnetic energy-momentum tensor.

remark: (1) It is not obvious what this has to do with energy and momentum, see Problem 5 for a hint, and LL for details.

properties: (1) $T^{\mu\nu}$ is symmetric,

$$T^{\mu\nu}(x) = T^{\nu\mu}(x)$$

(2) $T^{\mu\nu}$ is traceless,

$$T^{\mu}_{\mu}(x) = 0$$

proof: (1) $\int^{\mu\nu} = \int^{\nu\mu}$, and $F^{\mu\alpha} F^{\nu}_{\alpha} = \int^{\lambda\mu} F^{\lambda\alpha} \int^{\nu\lambda} F^{\alpha\lambda}$
 $= \delta^{\lambda\mu} F^{\lambda\alpha} F^{\nu\lambda} = F^{\nu\lambda} F^{\lambda\mu}$

$$(2) \quad -\frac{1}{4\pi} T^{\mu}_{\mu} = F^{\mu\alpha} F^{\mu}_{\alpha} - \frac{1}{4} \int^{\mu\nu} F_{\lambda\alpha} F^{\lambda\alpha} = 0 \quad \square$$

remark: (2) $\alpha \in \int^{\mu\nu} \Rightarrow T^{\mu\nu}$ can be decomposed into T^{00} plus a Euclidean vector T^{0i} plus a symmetric Euclidean tensor T^{ij} .

2.3 The continuity equation for the energy-momentum tensor

the absence of matter the

properties: the $\int^{\mu\nu}$ energy-momentum tensor obeys the continuity eq.

$$\partial_{\nu} T^{\mu\nu}(x) = 0$$

proof: $\int^{\mu\nu} (2.2) \Rightarrow \partial_{\nu} T^{\mu\nu} = \frac{1}{4\pi} \left[-\partial_{\nu} F^{\mu\alpha} F^{\nu}_{\alpha} + \frac{1}{4} \partial_{\nu} \delta^{\mu\nu} F_{\lambda\alpha} F^{\lambda\alpha} \right]$

$$= \frac{1}{4\pi} \left[-(\partial_\nu F_T^{\nu\kappa}) F^\nu_\kappa - F_T^{\nu\kappa} \partial_\nu F^\nu_\kappa + \frac{1}{2} \partial_T F_{\kappa\lambda} F^{\kappa\lambda} \right]$$

$$\partial_\nu F^\nu_\kappa = 0 \quad \text{by EL}$$

$$\begin{aligned} F_{\kappa\lambda} \partial_T F^{\kappa\lambda} &= \int_{\kappa\lambda} \int_{\lambda\sigma} F^{\lambda\sigma} \partial_T \int_{\sigma\epsilon} \int_{\epsilon\mu} F^{\mu\nu} = \delta_\gamma^\epsilon \delta_\delta^\nu F^{\lambda\sigma} \partial_T F_{\epsilon\nu} \\ &= F^{\lambda\sigma} \partial_T F_{\lambda\sigma} = (\partial_T F_{\kappa\lambda}) F^{\kappa\lambda} \rightarrow \partial_T F_{\kappa\lambda} F^{\kappa\lambda} - 2(\partial_T F_{\kappa\lambda}) F^{\kappa\lambda} \end{aligned}$$

$$= \frac{1}{4\pi} \left[-(\partial_\nu F_T^{\nu\kappa}) F^\nu_\kappa + \frac{1}{2} (\partial_T F_{\kappa\lambda}) F^{\kappa\lambda} \right]$$

$$\text{Problem 1} \rightarrow 0 = \partial_T F_{\kappa\lambda} + \partial_\kappa F_{\lambda T} + \partial_\lambda F_{T\kappa}$$

$$= \frac{1}{4\pi} \left[-(\partial_\nu F_T^{\nu\kappa}) F^\nu_\kappa - \frac{1}{2} (\partial_\kappa F_{\lambda T}) F^{\kappa\lambda} - \frac{1}{2} (\partial_\lambda F_{T\kappa}) F^{\kappa\lambda} \right]$$

$$= \frac{1}{4\pi} \left[-(\partial_\nu F_T^{\nu\kappa}) F^\nu_\kappa + \frac{1}{2} (\partial_\kappa F_{\lambda T}) F^{\kappa\lambda} + \frac{1}{2} (\partial_\lambda F_{T\kappa}) F^{\kappa\lambda} \right] = 0$$

Wolley: In the proof of matter the continuity eq. gets modified to

$$\partial_\nu \bar{T}_\mu^\nu(x) = -\frac{1}{c} F_T^\nu(x) j_\nu(x)$$

proof: Problem 6

remark: For any rank- $(n+1)$ tensor field $t^{\mu_1 \dots \mu_{n+1}}(x)$ the continuity eq. $\partial_T t^{\mu_1 \dots \mu_{n+1}}(x) = 0$ implies a conservation law for the rank- n tensor $t^{\mu_1 \dots \mu_n}(x)$ by the arguments from § 2.1. (2) ~ § 2.1 is the case $n=0$; the proposition above is the case $n=1$.

Problem 6

proof of Wolley

Problem 7

§ 2.1 laws for the stress-energy field

1/20/17

2.4 Gauge invariance

Let $\chi(x)$ be an arbitrary scalar fct. of spacetime.

def. 1: A transformation of the potential $A^\mu(x)$ according to

$$A^\mu(x) \rightarrow A'^\mu(x) = A^\mu(x) - \partial^\mu \chi(x)$$

is called a gauge transformation.

proposition: The action from [1.1 eqn.] is invariant under gauge transformations.

proof: $F^{\mu\nu}(x) = \partial^\mu A^\nu(x) - \partial^\nu A^\mu(x)$
 $\rightarrow \partial^\mu A^\nu(x) - \partial^\mu \partial^\nu \chi(x) - \partial^\nu A^\mu(x) + \partial^\nu \partial^\mu \chi(x)$
 $= F^{\mu\nu}(x) \rightarrow$ the field tensor is invariant
 \rightarrow S_{em} is invariant

$$S'_{\text{int}} = -\frac{1}{c} \int d^4x \int_{\mathcal{F}}(\lambda) A^\mu(x) \rightarrow S'_{\text{int}} + \frac{1}{c} \int d^4x \int_{\mathcal{F}}(\lambda) \partial^\mu \chi(x)$$

$$\stackrel{\text{part. int.}}{=} S_{\text{int}} - \frac{1}{c} \int d^4x \underbrace{(\partial^\mu \int_{\mathcal{F}}(\lambda)) \chi(x)}_{=0 \text{ by } \S 2.1} = S_{\text{int}}$$

\rightarrow S'_{int} is invariant \square

remark: (1) The potential is not unique. This is a result of the fact that $F^{\mu\nu}$ depends only on derivations of A^μ .

(2) We may choose a gauge before to impose a particular condition on A^μ ("fixing the gauge").

workey: A^μ can be chosen such that $\partial_\mu A^\mu(x) = 0$ "Lorentz gauge" \downarrow "no t.t.!"

proof: Choose χ such that it solves the PDE $\partial^\mu \partial_\mu \chi(x) = \partial_\mu A^\mu(x)$

$$\rightarrow \partial_\mu A'^\mu(x) = \partial_\mu A^\mu(x) - \partial_\mu \partial^\mu \chi(x) = 0 \quad \square$$

remark: (3) $\partial_\mu A^\mu$ is a Lorentz scalar \rightarrow the Lorentz gauge is Lorentz invariant

problem 8
order's gauge

Electric and magnetic fields

1.1 The field laws in terms of Euclidean vector fields

PHYS 610 dI §5.4 → As an antisymmetric Minkowski tensor, $F_{\mu\nu}$ has the form

$$F_{\mu\nu} = \begin{pmatrix} 0 & E_x & E_y & E_z \\ -E_x & 0 & -\vec{D}_t & \vec{D}_y \\ -E_y & \vec{D}_t & 0 & -\vec{D}_x \\ -E_z & -\vec{D}_y & \vec{D}_x & 0 \end{pmatrix}$$

$$= \begin{pmatrix} 0 & \vec{E} \\ -\vec{E} & \vec{D}_{ij} \end{pmatrix}$$

LL's notation:
 $\vec{E} = (E^1, E^2, E^3)$
 $\equiv (E_x, E_y, E_z)$
 $\Rightarrow (E_x, E_y, E_z)$
 same for \vec{D}

with $\vec{E}(x) = (E_x(x), E_y(x), E_z(x))$ a Euclidean vector field
 and $\vec{D}(x) = (D_x(x), D_y(x), D_z(x))$ a Euclidean pseudovector field

def. 1: (a) $\vec{E}(x)$ is called electric field, and $\vec{D}(x)$ is called magnetic field.

(b) The antisymmetric Euclidean tensor

$$D_{ij} = \begin{pmatrix} 0 & -\vec{D}_t & \vec{D}_y \\ \vec{D}_t & 0 & -\vec{D}_x \\ -\vec{D}_y & \vec{D}_x & 0 \end{pmatrix} = D^{ij} = -\epsilon_{ijk} D^k$$

is called magnetic field tensor.

Week 2 (short due w/br/2 to TLK by)

§5,6,7,8)

remark: (1) $F^{\mu\nu} = g^{\mu\alpha} g^{\nu\beta} F_{\alpha\beta} = \begin{pmatrix} + \\ - \\ - \\ + \end{pmatrix} \begin{pmatrix} + \\ - \\ - \\ + \end{pmatrix}$, $F_{\mu\nu} = \begin{pmatrix} 0 & -\vec{E} \\ \vec{E} & \vec{D}_{ij} \end{pmatrix}$
 (1') $\epsilon^{ijkl} = g^{ik} g^{jl} g^{\lambda\sigma} \epsilon_{\lambda\sigma\mu\nu} = -\epsilon_{ijkl} = -\epsilon_{ij\lambda} = \epsilon_{ij\lambda}$ etc.

PHYS 610 dI §5.4 → A^μ has the form (1'') $\epsilon^{ijkl} \epsilon_{iklm} = -(\delta^j_l \delta^i_k - \delta^j_k \delta^i_l)$ (NB the overall sign!)

$$A^\mu(x) = (\varphi(x), \vec{A}(x)) \quad (\varphi(x) \equiv A_0(x) = A^0(x))$$

and $\vec{A}(x)$ a Euclidean vector field

def. 2: $\varphi(x)$ is called scalar potential, and $\vec{A}(x)$ is called vector potential

remark: (2) $\gamma^\mu(x) = (c\varphi(x), \vec{A}(x))$ with γ the large chunk and \vec{A} the small chunk. see §2.1.

3.2 Maxwell's equations

With the field eqs from §1.3 in terms of \vec{E} and \vec{A} .

$$\S 1.3 (*) : \quad \partial_\mu \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta} = 0$$

$$\underline{\nu=0} : \quad \underline{0} = 2 \left(-\partial_1 F_{23} + \partial_2 F_{13} - \partial_3 F_{12} \right) = 2 \left(\partial_1 A_x + \partial_2 A_y + \partial_3 A_z \right) = \underline{\vec{\nabla} \cdot \vec{A}}$$

$$\underline{\nu=1} : \quad \underline{0} = 2 \left(\partial_0 F_{23} - \partial_2 F_{03} + \partial_3 F_{02} \right) = \frac{-1}{c} \partial_t A_x - \partial_2 E_z + \partial_3 E_y$$

$$= \underline{\frac{-1}{c} \partial_t A_x - (\vec{\nabla} \times \vec{E})_x}$$

and cyclic

Proposition 1: $\S 1.3 (*)$ is equivalent to

$\vec{\nabla} \cdot \vec{A}(x) = 0$	(1)
$\frac{1}{c} \partial_t \vec{A}(x) + \vec{\nabla} \times \vec{E}(x) = 0$	(2)

Remark: (1) Then on 4 homogeneous PDEs for the six field components $E_{x,y,z}, A_{x,y,z}$.

Now under the EC eq., $\S 1.3 (a)$:

$$\partial_\mu F^{\mu\nu} = \frac{4\pi}{c} j^\nu$$

$$\underline{\nu=0} : \quad \partial_0 F^{00} + \partial_i F^{i0} = \frac{4\pi}{c} j^0 \rightarrow \frac{\partial}{\partial x^i} E^i = \frac{4\pi}{c} \rho \rightarrow \underline{\vec{\nabla} \cdot \vec{E} = 4\pi \rho}$$

$$\underline{\nu=1} : \quad \partial_0 F^{01} + \partial_i F^{i1} = \frac{4\pi}{c} j^1 \rightarrow -\frac{1}{c} \partial_t E_x + \partial_2 A_z - \partial_3 A_y = \frac{4\pi}{c} j_x$$

and cyclic

$$\rightarrow \underline{-\frac{1}{c} \partial_t E_x + (\vec{\nabla} \times \vec{A})_x = \frac{4\pi}{c} j_x}$$

Proposition 2: $\S 1.3 (*)$ is equivalent to

$\vec{\nabla} \cdot \vec{E}(x) = 4\pi \rho(x)$	(3)
$-\frac{1}{c} \partial_t \vec{E}(x) + \vec{\nabla} \times \vec{A}(x) = \frac{4\pi}{c} \vec{j}(x)$	(4)

~~plz f.p~~ plz f.p

(5) units: we use cgs units: $[\text{charge}] = esu = g^{1/2} \text{cm}^{3/2} / s$

$$[\rho] = g^{1/2} / \text{cm}^{3/2} s, \quad [\vec{j}] = g^{1/2} / \text{cm}^{3/2} s^2$$

$$[\vec{E}] = g^{1/2} / \text{cm}^{1/2} s, \quad [\vec{D}] = g^{1/2} / \text{cm}^{1/2} s$$

in SI units, $[\text{charge}] = C, [\rho] = C/\text{m}^3, [\vec{j}] = C/\text{m}^2 s$

$$[\vec{E}] = N/C, \quad [\vec{D}] = N/\text{Am} \quad (\text{NB: } C = As)$$

and the Maxwell eqs (2)-(4) read

$\partial_t \vec{D} + \nabla \times \vec{E} = 0$	(2')
$\nabla \cdot \vec{E} = \frac{1}{\epsilon_0} \rho$	(3')
$-\partial_t \vec{E} + \frac{1}{\mu_0 \epsilon_0} \nabla \times \vec{D} = \frac{1}{\epsilon_0} \vec{j}$	(4')

where $\epsilon_0 = 8.85 \times 10^{-12} \text{ C}^2 / \text{Nm}^2$ "permittivity of vacuum"

$\mu_0 = 4\pi \times 10^{-7} \text{ N/A}^2$ "permeability of vacuum"

and $\mu_0 \epsilon_0 = 1/c^2 = 1.11 \times 10^{-17} \text{ s}^2 / \text{m}^2$

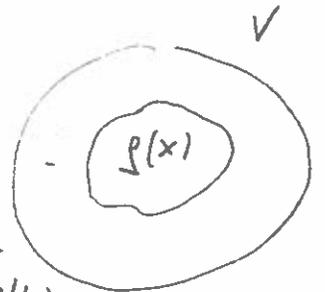
remark: (2) Eqs (1)-(4) are called Maxwell's eqs. Their solutions determine physical field configurations for given charge and current densities.

(3) Eqs (1)-(4) are equivalent to Eqs. (5), (6) in §1.3.

(4) \vec{E} and \vec{A} are Euclidean vector fields. \rightarrow The Lorentz invariance of (1)-(4) is for now obvious, whereas the Lorentz invariance of (5), (6) is.

3.3 Discussion of Maxwell's equations

Consider a localized charge density ρ inside a volume V with surface (V) .



$$\nabla \cdot \vec{E} \text{ (2)} \rightarrow \int_V d^3x \nabla \cdot \vec{E}(\vec{x}, t) = 4\pi \int_V d^3x \rho(\vec{x}, t) = \underline{\underline{4\pi Q(t)}} = \underline{\underline{\text{total charge}}}$$

$$\int_{(V)} d\vec{s} \cdot \vec{E}(\vec{x}, t) = \underline{\underline{\text{flux of } \vec{E} \text{ through the surface } (V)}}.$$

remark: (0) This is called Gauss's law.
(1) Electric charges are the sources of electric fields.

$$\nabla \cdot \vec{A} \text{ (1)} \rightarrow \underline{\underline{0}} = \int_V d^3x \nabla \cdot \vec{A}(\vec{x}, t) = \int_{(V)} d\vec{s} \cdot \vec{A}(\vec{x}, t) = \underline{\underline{\text{flux of } \vec{A} \text{ through any closed surface}}}$$

remarks: (2) The magnetic field has no sources! Equivalent statement: There is no magnetic charge (aka magnetic monopole).

(3) In the manifestly Lorentz-invariant formalism of §1.3 this symmetry manifests itself in the covariance of the field eqs:

$$\partial_\mu F^{\mu\nu} = \frac{4\pi}{c} j^\nu, \text{ but } \partial_\mu \tilde{F}^{\mu\nu} = 0$$

Consider a bounded surface S
with boundary (S) .



$$\begin{aligned} \text{M-eg. (2)} \rightarrow -\frac{1}{c} \int_S d\vec{s} \cdot \partial_t \vec{A}(\vec{x}, t) &= \int_S d\vec{s} \cdot (\vec{\nabla} \times \vec{E})(\vec{x}, t) \stackrel{\text{Stokes}}{\downarrow} \int_{(S)} d\vec{l} \cdot \vec{E}(\vec{x}, t) = \frac{\text{circulation}}{\text{of } \vec{E} \text{ around } (S)} \\ &= -\frac{1}{c} \frac{d}{dt} \int_S d\vec{s} \cdot \vec{A}(\vec{x}, t) = -\frac{1}{c} \frac{d}{dt} \Phi(t) = \frac{\text{time derivative of flux}}{\text{of } \vec{A} \text{ through the surface } S} \end{aligned}$$

Remark: (4) This is called Faraday's law of induction

(5) Consider a closed \vec{E} -field line



$$\text{Then } \int d\vec{l} \cdot \vec{E} > 0 \rightarrow \dot{\Phi}(t) < 0$$

\rightarrow In a static \vec{A} -field (i.e. perpendicular for $\vec{A} = 0$) there can be no closed \vec{E} -field lines!

$$\begin{aligned} \text{M-eg. (4)} \rightarrow \int_S d\vec{s} \cdot (\vec{\nabla} \times \vec{A})(\vec{x}, t) &= \frac{4\pi}{c} \int_S d\vec{s} \cdot \vec{j}(\vec{x}, t) + \frac{1}{c} \int_S d\vec{s} \cdot \partial_t \vec{E}(\vec{x}, t) \\ \text{circulation} &= \int_{(S)} d\vec{l} \cdot \vec{A}(\vec{x}, t) = \frac{4\pi}{c} \int \vec{j} + \frac{1}{c} \frac{d}{dt} \int_S d\vec{s} \cdot \vec{E}(\vec{x}, t) \\ \text{of } \vec{A} &= \frac{4\pi}{c} \times \text{total current} + \frac{1}{c} \times \text{displacement current} \end{aligned}$$

Remark: (6) This is called Ampere-Maxwell law

(7) Currents induce \vec{A} -fields, and vice versa

(8) For static fields, we have Ampere's law

$$\int d\vec{l} \cdot \vec{A} = \frac{4\pi}{c} \vec{j}$$

The displacement current was later added by Maxwell

3.4. Relation between fields and potentials

$$\begin{aligned} \S 3.1 \rightarrow E^i &= -F^{0i} = -\partial^0 A^i + \partial^i A^0 \\ &= -\partial_0 A^i - \partial_i A^0 = -\frac{1}{c} \partial_t A^i - \partial_i \varphi \end{aligned}$$

$$\rightarrow \boxed{\vec{E}(\vec{x}, t) = -\vec{\nabla} \varphi(\vec{x}, t) - \frac{1}{c} \partial_t \vec{A}(\vec{x}, t)}$$

Remark: (1) In general, both \vec{A} and φ determine \vec{E} .

$$(2) \text{ Gauge transformation: } A^\mu \rightarrow A^\mu - \partial^\mu \chi \quad \rightarrow \quad \begin{aligned} \varphi &\rightarrow \varphi - \frac{1}{c} \partial_t \chi \\ \vec{A} &\rightarrow \vec{A} + \vec{\nabla} \chi \end{aligned}$$

$$\rightarrow \vec{E} \rightarrow \vec{E} + \vec{\nabla} \frac{1}{c} \partial_t \chi - \frac{1}{c} \partial_t \vec{\nabla} \chi = \vec{E} \quad \underline{\vec{E} \text{ is unaffected}}$$

$$\S 3.1 \rightarrow F_{12} = (\partial_1 A_2 - \partial_2 A_1) = (\vec{\nabla} \times \vec{A})_3 = -(\vec{\nabla} \times \vec{A})^3$$

$$\dots \rightarrow -\vec{\nabla}_t = -\vec{\nabla}^3 \quad \text{and cyclic}$$

$$\rightarrow \boxed{\vec{\nabla}(\vec{x}, t) = \vec{\nabla} \times \vec{A}(\vec{x}, t)}$$

Remark: (3) $\vec{\nabla}$ invariant under gauge transform, since $\vec{\nabla} \times (\vec{\nabla} \chi) = 0$

3.5 Charges & electromagnetic fields

Attitude w.r. to: Field eqs determine fields for give charge and currents.

Question: For give fields, what is their influence on a point charge

Let a point particle with charge e be at point $\vec{y}(t)$ with velocity

$$\dot{\vec{y}}(t) =: \vec{v}(t)$$

$\rightarrow \rho(\vec{x}, t) = e \delta(\vec{x} - \vec{y}(t))$ charge density

$\vec{j}(\vec{x}, t) = \rho(\vec{x}, t) \vec{v}(t)$ current density

$\mathcal{J}^\Gamma = (\rho, \vec{j}), \mathcal{A}^\Gamma = (\phi, \vec{A})$ 4-current, 4-potential

$$\begin{aligned} \text{Eq. 1.1 ex. (2)} \rightarrow \underline{\mathcal{L}_{int}} &= -\frac{1}{c} \int d^4x \mathcal{J}^\Gamma(x) \mathcal{A}^\Gamma(x) \\ &= -\frac{1}{c} c \int dt \int d^3x \rho(\vec{x}, t) \phi(\vec{x}, t) + \frac{1}{c} c \int dt \int d^3x \vec{j}(\vec{x}, t) \cdot \vec{A}(\vec{x}, t) \\ &= -c \int dt e \phi(\vec{y}, t) + e \int dt \vec{v} \cdot \vec{A}(\vec{y}, t) \end{aligned}$$

Now consider the Lagrangian of the point particle, $L_{int}(\vec{y}, \dot{\vec{y}}, t)$,

which is related to \mathcal{L}_{int} via $\mathcal{L}_{int} = c \int dt L_{int}(\vec{y}, \dot{\vec{y}}, t)$

$$\rightarrow \underline{L_{int}(\vec{y}, \dot{\vec{y}}, t) = -e \phi(\vec{y}, t) + \frac{e}{c} \vec{v} \cdot \vec{A}(\vec{y}, t)} \quad \text{where } \underline{\vec{v} = \dot{\vec{y}}}$$

Remark: (1) ϕ and \vec{A} are examples of the scalar and vector potential, respectively, furnished from Maxwell's, see § 3.5a.

(2) § 1.1 ex. (2) is consistent with the action of Maxwell's

(3) L_{int} must be augmented by the Lagrangian L_0 for a free particle. Field eqs on Lorentz invariant $\rightarrow L_0$ must be the Einstein L_0 for consistency (although Galilean L_0 is okay approx. for $v \ll c$).

table 9
particle in \vec{E} and \vec{B} fields
table 10
non-rel. in \vec{E} -field
table 11
... ..

1.5c Misner: Potentials in Mechanics, and Newton's 2nd Law

Let $L_0(\vec{v})$ be the Lagrangian for a free point particle. Consider the interaction of the particle with its environment.

exer: The influence of the environment on the particle is described

by (a) a scalar potential $U(\vec{x}, t)$

and (b) a vector potential $\vec{V}(\vec{x}, t)$

and let the Lagrangian is

$$L(\vec{x}, \vec{v}, t) = L_0(\vec{v}) - U(\vec{x}, t) + \vec{v} \cdot \vec{V}(\vec{x}, t)$$

remark: (1) L_0 may describe a Galilean particle, $L_0^G = \frac{m}{2} \vec{v}^2$, or a Einsteinian one, $L_0^E = -mc^2 \sqrt{1 - \vec{v}^2/c^2}$.

(2) U and \vec{V} are determined either by experiment, or by another theory, not by Mechanics

example: (1) Particle in a gravitational field. U determined by experiment (Laplace) or a more general theory (GR, Einstein). $\vec{V} = 0$.

(2) Charged particle in an electromagnetic field. U and \vec{V} determined by a different theory (Maxwell), see § 1.5: $U(\vec{x}, t) = -e\phi(\vec{x}, t)$, $\vec{V}(\vec{x}, t) = \frac{e}{c} \vec{A}(\vec{x}, t)$.

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def.: In the presence of a vector potential we define the particle momentum by $\vec{p} = \partial L_0 / \partial \vec{v}$

remark: (2) $\vec{\pi} := \partial L / \partial \vec{v}$ is sometimes called generalized momentum

Ans:

$$F_{\lambda\sigma} F^{\lambda\sigma} = -2(\vec{E}^2 - \vec{D}^2)$$

Proof:

$$F_{\lambda\sigma} F^{\lambda\sigma} = F_{00} F^{00} + F_{0i} F^{0i} + F_{i0} F^{i0} + F_{ij} F^{ij}$$

$$= 0 + E_i E^i + E_i E^i + D_{ij} D^{ij}$$

$$= -2\vec{E}^2 + \epsilon_{ij\lambda} D^\lambda \epsilon^{ij\lambda} D_\lambda$$

∫ ∫ ∫ over (V)

$$= -2\vec{E}^2 - (\delta_j^i \delta_\lambda^k - \delta_j^k \delta_\lambda^i) D^\lambda D_k$$

$$= -2\vec{E}^2 - (D^\lambda D_\lambda - D_\lambda D^\lambda)$$

$$= -2\vec{E}^2 - 2 D^\lambda D_\lambda$$

$$= -2\vec{E}^2 + 2\vec{D}^2$$

ik where

$$F_{0i} = E^i = -E_i$$

$$F_{i0} = -E^i = E_i$$

proposition: The equation of motion takes the form

$$\frac{d}{dt} \vec{p}(\vec{x}, t) = \vec{F}^{(1)}(\vec{x}, t) + \vec{F}^{(2)}(\vec{x}, \vec{v}, t) \quad \text{Newton's 2nd law}$$

vill	$\vec{F}^{(1)}(\vec{x}, t) = -\vec{\nabla} U(\vec{x}, t) - \partial_t \vec{V}(\vec{x}, t)$	a velocity-independent force
al	$\vec{F}^{(2)}(\vec{x}, t) = \vec{v} \times (\vec{\nabla} \times \vec{V}(\vec{x}, t))$	a velocity-dependent force

proof: Euler-Lagrange \rightarrow

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial v_i} = \frac{d}{dt} p_i = \frac{d}{dt} p_i + \frac{d}{dt} V_i \stackrel{EL}{=} \frac{\partial \mathcal{L}}{\partial x_i} = -\partial_i U + v_j \partial_i v_j$$

$$\rightarrow \frac{d}{dt} p_i = \underbrace{-\partial_i U - \partial_t V_i}_{= F_i^{(1)}} - (\partial_j v_j) v_i + v_j \partial_i v_j$$

$$\begin{aligned} (\vec{v} \times (\vec{\nabla} \times \vec{V}))_i &= \epsilon_{ijl} v_j \epsilon_{lmn} \partial_l V_n \\ &= (\delta_{in} \delta_{jl} - \delta_{il} \delta_{jn}) v_j \partial_l V_n \\ &= v_j \partial_i v_j - v_j \partial_j v_i \end{aligned}$$

$$= -\partial_i U - \partial_t V_i + (\vec{v} \times (\vec{\nabla} \times \vec{V}))_i = \underline{\underline{F_i^{(1)} + F_i^{(2)}}} \quad \square$$

Week 3

Problem 3
(9, 10, 11)

1.6 Poynting's theorem

Consider the continuity eq. for \vec{p} for $\mu=0$:

$$\begin{aligned} \underline{\underline{T^{00}}} &= \frac{1}{4\pi} F^{0\alpha} F^0_{\alpha} + \frac{1}{16\pi} F^{\alpha\lambda} F^{\lambda\alpha} \stackrel{1}{=} \frac{1}{4\pi} F^{0i} F^0_i - \frac{1}{8\pi} (\vec{E}^2 - \vec{B}^2) = \frac{1}{4\pi} F^{0i} F^{0\alpha} \eta_{\alpha i} - \frac{1}{8\pi} (\vec{E}^2 - \vec{B}^2) \\ &= \frac{1}{4\pi} (-E^i)(-E_i) - \frac{1}{8\pi} (\vec{E}^2 - \vec{B}^2) = \frac{1}{4\pi} \vec{E}^2 - \frac{1}{8\pi} (\vec{E}^2 - \vec{B}^2) = \underline{\underline{\frac{1}{8\pi} (\vec{E}^2 + \vec{B}^2)}} \end{aligned}$$

$$\underline{\underline{T^{0i}}} = \frac{1}{4\pi} F^{0\alpha} F^i_{\alpha} = \frac{1}{4\pi} F^{0j} F^i_j = \frac{1}{4\pi} (-E_j \partial_t + E_t \partial_j) = \underline{\underline{\frac{1}{4\pi} (\vec{E} \times \vec{B})^i}} \quad \text{ad cyclic}$$

$$\rightarrow \partial_\nu T^{0\nu} = \frac{1}{c} \partial_t \frac{1}{8\pi} (\vec{E}^2 + \vec{B}^2) + \vec{\nabla} \cdot \frac{1}{4\pi} (\vec{E} \times \vec{B}) \quad \text{this is the lhs of the continuity eq}$$

$$\underline{\underline{F^{0\nu} j_\nu}} = \underline{\underline{F^{0i} j_i}} = \underline{\underline{\vec{E} \cdot \vec{j}}} \quad \text{this is the rhs}$$

Define $u(\vec{x}, t) := \frac{1}{8\pi\epsilon_0} (\vec{E}^2(\vec{x}, t) + \vec{B}^2(\vec{x}, t))$ energy density of the field (see below)

$\vec{P}(\vec{x}, t) := \frac{c}{4\pi} \vec{E}(\vec{x}, t) \times \vec{B}(\vec{x}, t)$ "Poynting vector" = energy current density (see below)

\rightarrow In terms of \vec{E} and \vec{B} the continuity eq. takes the form

$$\partial_t u + \vec{\nabla} \cdot \vec{P} = - \vec{E} \cdot \vec{j}$$

$\partial_t u(\vec{x}, t) + \vec{\nabla} \cdot \vec{P}(\vec{x}, t) = - \vec{E}(\vec{x}, t) \cdot \vec{j}(\vec{x}, t)$
"Poynting's theorem"

Remark: (1) For $\vec{j} = 0$, this expresses local energy conservation. It is analogous to §2.1 with $\rho \rightarrow u$ (energy density) and $\vec{j} \rightarrow \vec{P}$ (energy current density).

(2) §2.5 $\rightarrow \vec{j} \cdot \vec{E} = e \vec{v} \cdot \vec{E}$ = work per time done by the fields on a charge e .

$\rightarrow \vec{j} \cdot \vec{E} = \frac{1}{V} e \vec{v} \cdot \vec{E}$ = work done per time and volume = power density

\rightarrow For $\vec{j} \neq 0$, P's theorem still expresses energy conservation.

$\text{energy change} = - \text{energy transported by energy current} - \text{work done by fields on charges}$

(3) We still need to show that $u(\vec{x}, t)$ can be naturally interpreted as the energy density of the field. Let \vec{j} be the current density due to just one particle,

do!

as in §3.5 (for many particles, just run over them)

$$\rightarrow \int d\vec{x} \vec{j} \cdot \vec{E} = \int d\vec{x} \vec{E}(\vec{x}, t) \cdot e\vec{v} \cdot \delta(\vec{x} - \vec{y}) = e\vec{v} \cdot \vec{E}(\vec{y})$$

where \vec{y} is the position of the particle.

Consider a nonrelativistic particle for simplicity (the argument also holds in the relativistic case, see Prob 12)

$$\rightarrow E_{\text{kin}} = \frac{m}{2} v^2 \quad \text{kinetic energy}$$

$$\rightarrow \frac{d}{dt} E_{\text{kin}} = m\vec{v} \cdot \frac{d\vec{v}}{dt} = \vec{v} \cdot \frac{d\vec{p}}{dt} \stackrel{\text{§3.5, eq. (4)}}{=} e\vec{v} \cdot \vec{E}$$

Now integrate Poynting's theorem over all of space:

$$\begin{aligned} \frac{d}{dt} \int d\vec{x} u(\vec{x}, t) + \underbrace{\int d\vec{x} \vec{\nabla} \cdot \vec{P}(\vec{x}, t)}_{\substack{= \int d\vec{s} \vec{P} = 0 \\ \text{via } \vec{P} = 0 \text{ at } \infty}} &= - \int d\vec{x} \vec{j} \cdot \vec{E} = -e\vec{v} \cdot \vec{E} \\ &= -\frac{d}{dt} E_{\text{kin}} \end{aligned}$$

$$\rightarrow \frac{d}{dt} (U + E_{\text{kin}}) = 0 \quad \text{where } U = \int d\vec{x} u(\vec{x}, t)$$

\rightarrow U must be the field energy, since the energy of the particle plus the energy of the field must be conserved

$\rightarrow u(\vec{x}, t)$ must be the energy density of the field

(4) Integrate over a finite volume \rightarrow Energy may change due to a energy current across the volume boundary $\rightarrow \vec{P}$ should be interpreted as the energy current density of the field.

(5) The remaining components of the continuity eq. for §2.2,

$$\partial_\nu T^{\mu\nu} = -\frac{1}{c} F^{\mu\nu} j_\nu$$

express the fact that the energy current density is also conserved.

Problem 12

energy density

DO

1/30/17

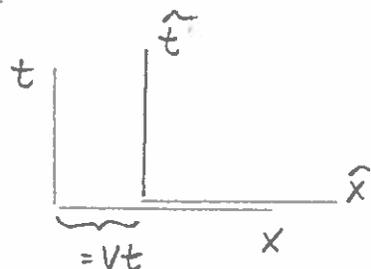
§4 Lorentz transformations of the fields

4.1 Physical interpretation of a Lorentz boost

Consider two inertial frames, cs and \tilde{cs}

Let \tilde{cs} move with respect to cs with

a constant velocity $\vec{V} = (V, 0, 0)$.



Problem: 610/21, 23 \rightarrow The transformation from cs to \tilde{cs} is accomplished by a Lorentz boost

$$c\tilde{t} = ct \cosh \phi - x \sinh \phi$$

$$\tilde{x} = ct \sinh \phi - x \cosh \phi$$

Consider the origin of \tilde{cs} as viewed in cs . Then $\tilde{x}' = 0$

$$\rightarrow x \cosh \phi = -ct \sinh \phi$$

$$\rightarrow v = x/t = -c \tanh \phi$$

$$\rightarrow \sinh \phi = \frac{\tanh \phi}{\sqrt{1 - \tanh^2 \phi}} = \frac{v/c}{\sqrt{1 - v^2/c^2}}, \quad \cosh \phi = \frac{1}{\sqrt{1 - v^2/c^2}}$$

Remark: (1) For $c \rightarrow \infty$ we recover the Galilean boost

$$\tilde{x} = x + vt, \quad \tilde{t} = t$$

\rightarrow A Lorentz boost along the x -axis is given by

$$\Lambda_{\nu}^{\mu} = \begin{pmatrix} \cosh \phi & \sinh \phi & 0 & 0 \\ \sinh \phi & \cosh \phi & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

with $\cosh \phi = \gamma$, $\sinh \phi = \frac{v}{c} \gamma$, where $\gamma = \frac{1}{\sqrt{1 - v^2/c^2}}$

Problem 13

Addition of velocities

Let \vec{u}

4.2 Transformation of \vec{E} and \vec{B} under a Lorentz boost

Consider the field tensor $F^{\mu\nu}$ in CS. The field tensor $\tilde{F}^{\mu\nu}$ in \tilde{CS} is

$$\tilde{F}^{\mu\nu} = \Lambda^\mu{}_\alpha \Lambda^\nu{}_\beta F^{\alpha\beta} \quad \text{and} \quad \tilde{x}^\mu = \Lambda^\mu{}_\nu x^\nu$$

Now let $\Lambda^\mu{}_\nu$ be a Lorentz boost.

$$\rightarrow \tilde{F}^{\mu\nu} = (\Lambda F \Lambda^T)^{\mu\nu} = \begin{pmatrix} \omega \sin\phi & \omega \cos\phi & 0 & 0 \\ \omega \cos\phi & \omega \sin\phi & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & -\tilde{B}_z & \tilde{B}_y \\ E_y & \tilde{B}_z & 0 & -\tilde{B}_x \\ E_z & -\tilde{B}_y & \tilde{B}_x & 0 \end{pmatrix} \begin{pmatrix} \omega \cos\phi & \omega \sin\phi & 0 & 0 \\ \omega \sin\phi & \omega \cos\phi & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$= \begin{pmatrix} E_x \omega \cos\phi & -E_x \omega \sin\phi & -E_y \omega \cos\phi - \tilde{B}_z \omega \sin\phi & -E_z \omega \cos\phi + \tilde{B}_y \omega \sin\phi \\ E_x \omega \sin\phi & -E_x \omega \cos\phi & -E_y \omega \sin\phi - \tilde{B}_z \omega \cos\phi & -E_z \omega \sin\phi + \tilde{B}_y \omega \cos\phi \\ E_y & \tilde{B}_z & 0 & -\tilde{B}_x \\ E_z & -\tilde{B}_y & \tilde{B}_x & 0 \end{pmatrix} \begin{pmatrix} \omega \cos\phi & \omega \sin\phi & 0 & 0 \\ \omega \sin\phi & \omega \cos\phi & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$= \begin{pmatrix} 0 & \cdot & \cdot & \cdot \\ \tilde{E}_x & 0 & \cdot & \cdot \\ E_y \omega \cos\phi + \tilde{B}_z \omega \sin\phi & E_y \omega \sin\phi + \tilde{B}_z \omega \cos\phi & 0 & \cdot \\ E_z \omega \cos\phi - \tilde{B}_y \omega \sin\phi & E_z \omega \sin\phi - \tilde{B}_y \omega \cos\phi & \tilde{B}_x & \cdot \end{pmatrix}$$

$$\equiv \begin{pmatrix} 0 & -\tilde{E}_x & -\tilde{E}_y & -\tilde{E}_z \\ \tilde{E}_x & 0 & -\tilde{B}_z & \tilde{B}_y \\ \tilde{E}_y & \tilde{B}_z & 0 & -\tilde{B}_x \\ \tilde{E}_z & -\tilde{B}_y & \tilde{B}_x & 0 \end{pmatrix}$$

\rightarrow

$$\begin{aligned} \tilde{E}_x &= E_x \\ \tilde{E}_y &= E_y \omega \cos\phi + \tilde{B}_z \omega \sin\phi \\ \tilde{E}_z &= E_z \omega \cos\phi - \tilde{B}_y \omega \sin\phi \end{aligned}$$

(*)

$$\tilde{B}_x = \tilde{B}_x$$

$$\tilde{B}_y = \tilde{B}_y \omega \cos\phi - E_z \omega \sin\phi$$

$$\tilde{B}_z = \tilde{B}_z \omega \cos\phi + E_y \omega \sin\phi$$

Remark: (1) The field eqs were formulated in terms of Minkowski tensors \rightarrow

Their Lorentz invariance is

guaranteed. (*) reflects the Lorentz invariance of M's eqs, which are required to hold in all inertial frames.

remark: (2) let $v/c \ll 1$ and keep terms to $O(v/c)$.

$$\rightarrow \text{wrt } \phi \approx 1, \text{ wrt } \phi \approx v/c$$

$$\rightarrow \underline{\underline{\hat{\vec{E}} \approx \vec{E} - \frac{1}{c} \vec{v} \times \vec{B} + O(v^2/c^2), \quad \hat{\vec{B}} \approx \vec{B} + \frac{1}{c} \vec{v} \times \vec{E} + O(v^2/c^2)}}$$

(3) let $\vec{E} = 0$, i.e., no \vec{E} -field in CS.

$$\rightarrow \hat{\vec{E}} = -\frac{1}{c} \vec{v} \times \vec{B} \rightarrow \text{in CS there is an } \vec{E}\text{-field, as long as } \vec{B} \neq 0!$$

4.2 Lorentz invariants

From the field tensor $F^{\mu\nu}$ we can form two Lorentz scalars:

$$\boxed{\mathcal{F}^{(S)} := -\frac{1}{2} F^{\mu\nu} F_{\mu\nu}, \quad \mathcal{F}^{(A)} := \frac{1}{8} \epsilon^{\mu\nu\alpha\beta} F_{\mu\nu} F_{\alpha\beta}}$$

remark: (1) $\mathcal{F}^{(S)}$ is a scalar $\rightarrow \mathcal{F}^{(S)} = \mathcal{F}^{(S)}$ in all inertial frames

$\mathcal{F}^{(A)}$ is a pseudoscalar $\rightarrow |\mathcal{F}^{(A)}| = \text{const}$ in all inertial frames but $\mathcal{F}^{(A)} \rightarrow -\mathcal{F}^{(A)}$

§ 3.6 done $\rightarrow \underline{\underline{\mathcal{F}^{(S)}}} = -\frac{1}{2} (-2) (\vec{E}^2 - \vec{B}^2)$
 $= \underline{\underline{\vec{E}^2 - \vec{B}^2}}$

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$$\underline{\underline{\mathcal{F}^{(A)}}} = \frac{1}{8} \left[\begin{array}{l} \epsilon^{0123} F_{01} F_{23} + \epsilon^{0132} F_{01} F_{32} \\ + \epsilon^{1023} F_{10} F_{23} + \epsilon^{1032} F_{10} F_{32} \\ + \epsilon^{2103} F_{21} F_{03} + \epsilon^{2130} F_{21} F_{30} \\ + \epsilon^{3102} F_{31} F_{02} + \epsilon^{3120} F_{31} F_{20} \\ + (\text{3} \times \text{6} = \text{18 other terms}) \end{array} \right] = \frac{1}{4} \left[\begin{array}{l} \epsilon^{0123} F_{01} F_{23} + \epsilon^{0132} F_{01} F_{32} \\ + \epsilon^{1023} F_{10} F_{23} + \epsilon^{1032} F_{10} F_{32} \\ + \epsilon^{2103} F_{21} F_{03} + \epsilon^{2130} F_{21} F_{30} \\ + \epsilon^{3102} F_{31} F_{02} + \epsilon^{3120} F_{31} F_{20} \\ + \text{9 other terms} \end{array} \right]$$

$$= \frac{1}{4} [-E_x \vec{u}_x - E_y \vec{u}_y - E_z \vec{u}_z] \times 4 = \underline{\underline{-\vec{E} \cdot \vec{u}}}$$

proposition. The field variations

of a pseudoscalar ($\mathcal{F}^{(1)}$),
respectively,

$$\mathcal{F}^{(1)} = \vec{E}^2 - \vec{u}^2, \quad \mathcal{F}^{(2)} = \vec{E} \cdot \vec{u}$$

transform as a scalar ($\mathcal{F}^{(1)}$) under Lorentz transformations, i.e., their observed values have the same value in all inertial frames.

remark: (1) $\exists \vec{E} \perp \vec{u}$ in some inertial frame, then $\vec{E} \perp \vec{u}$ in all other inertial frames.

problem 15

15.1 Problem: Lorentz invariance of fields

(1) $\vec{E}^2 = \vec{u}^2$ in some frame

§5 The superposition principle of Maxwell theory

5.1 field solutions

proposition: Let $\rho^{(1)}(x), \vec{j}^{(1)}(x)$ ($\lambda=1,2$) be two charge and current densities. Let $\vec{E}^{(1)}(x), \vec{u}^{(1)}(x)$ be solutions of M's eqs for $\rho^{(1)}, \vec{j}^{(1)}$, and let $\lambda^{(1)} \in \mathbb{R}$. Then

$$\vec{E} = \lambda^{(1)} \vec{E}^{(1)} + \lambda^{(2)} \vec{E}^{(2)}, \quad \vec{u} = \lambda^{(1)} \vec{u}^{(1)} + \lambda^{(2)} \vec{u}^{(2)}$$

are solutions for $\lambda^{(1)} \rho^{(1)} + \lambda^{(2)} \rho^{(2)}, \lambda^{(1)} \vec{j}^{(1)} + \lambda^{(2)} \vec{j}^{(2)}$

$$\text{proof: } \vec{\nabla} \cdot \vec{E} - \frac{1}{c} \rho = \underbrace{\vec{\nabla} \cdot \vec{E}^{(1)} - \frac{1}{c} \rho^{(1)}}_{=0} + \underbrace{\vec{\nabla} \cdot \vec{E}^{(2)} - \frac{1}{c} \rho^{(2)}}_{=0} = 0$$

etc.

remark: (1) This is obviously true since the theory is linear!

(2) \exists if the action contained terms higher than quadratic in $F_{\mu\nu}$ it would not be true.

(3) A field theory that leads to linear field eqs is called Gaussian or linear

Wolley 1: Let $\vec{E}^{(\lambda)}(x), \vec{J}^{(\lambda)}(x)$ be solutions for $\vec{g}^{(\lambda)}(x), \vec{j}^{(\lambda)}(x)$ with $\lambda \in \mathbb{R}$, and let $\lambda(\lambda) \in \mathbb{R}$ be a sufficiently well behaved fct. of λ . Then

$$\vec{E}(x) = \int d\lambda \lambda(\lambda) \vec{E}^{(\lambda)}(x), \quad \vec{J}(x) = \int d\lambda \lambda(\lambda) \vec{J}^{(\lambda)}(x)$$

are solutions for $\vec{g}(x) = \int d\lambda \lambda(\lambda) \vec{g}^{(\lambda)}(x), \vec{j}(x) = \int d\lambda \lambda(\lambda) \vec{j}^{(\lambda)}(x)$

proof: Generalize prop 1 to $\lambda = 1, \dots, N$ and let $N \rightarrow \infty$.

Remark: (4) This can obviously be generalized to $\vec{E}^{(\vec{\lambda})}(x)$ with $\vec{\lambda} \in \mathbb{R}^3$.

Wolley 2: The most general solution of Π 's eqs. is obtained as

$$\vec{E}(x) = \vec{E}^{(0)}(x) + \vec{E}^{(p)}(x), \quad \vec{J}(x) = \vec{J}^{(0)}(x) + \vec{J}^{(p)}(x)$$

where $\vec{E}^{(0)}, \vec{J}^{(0)}$ are the most general solution in vacuum (i.e., for $\vec{g} = \vec{j} = 0$) and $\vec{E}^{(p)}, \vec{J}^{(p)}$ is a particular solution of the eqs. in the presence of \vec{g}, \vec{j} .

proof: Let \vec{E}, \vec{J} be any solution for \vec{g}, \vec{j} , and let $\vec{E}^{(p)}, \vec{J}^{(p)}$ be a particular solution Prop. 1 \Rightarrow

$$\vec{E}^{(0)} := \vec{E} - \vec{E}^{(p)}, \quad \vec{J}^{(0)} := \vec{J} - \vec{J}^{(p)}$$

is a solution for $\vec{g} = 0, \vec{j} = 0$.

Conversely, if $\vec{E}^{(0)}$ is a solution for $\vec{g} = \vec{j} = 0$, and $\vec{E}^{(p)}, \vec{J}^{(p)}$ is some solution for \vec{g}, \vec{j} , then $\vec{E} = \vec{E}^{(0)} + \vec{E}^{(p)}, \vec{J} = \vec{J}^{(0)} + \vec{J}^{(p)}$ is a solution for \vec{g}, \vec{j} . \square

5.2 Complex solutions of Π 's eqs consist of real fields \vec{E} and \vec{J} .

All physical solutions of Π 's eqs, however, it sometimes is convenient to find complex solutions and take the real part.

Proposition 1: Let \vec{E}, \vec{D} be complex solutions for complex \vec{S}, \vec{J} . Then \vec{E}^*, \vec{D}^* are solutions for \vec{S}^*, \vec{J}^* .

Proof: $\vec{\nabla} \cdot \vec{E} = \vec{\nabla} \cdot \text{Re } \vec{E} + i \vec{\nabla} \cdot \text{Im } \vec{E} = 4\pi \text{Re } \vec{S} + i 4\pi \text{Im } \vec{S}$
 $\rightarrow \left. \begin{array}{l} \vec{\nabla} \cdot \text{Re } \vec{E} = 4\pi \text{Re } \vec{S} \\ \vec{\nabla} \cdot \text{Im } \vec{E} = 4\pi \text{Im } \vec{S} \end{array} \right\} \rightarrow \vec{\nabla} (\text{Re } \vec{E} - i \text{Im } \vec{E}) = 4\pi (\text{Re } \vec{S} - i \text{Im } \vec{S})$
 etc.

Remark: (1) This again is true because of linearity.

Corollary 1: Let \vec{E}, \vec{D} be complex solutions for real (i.e., physical) sources \vec{S}, \vec{J} . Then $\text{Re } \vec{E}, \text{Re } \vec{D}$ are also solutions for \vec{S}, \vec{J} .

Proof: Corollary 1 $\rightarrow \text{Re } \vec{E}, \text{Re } \vec{D}$ are solutions for $\text{Re } \vec{S} = \vec{S}, \text{Re } \vec{J} = \vec{J}$.

Remark: (2) In this case, $\text{Im } \vec{E}, \text{Im } \vec{D}$ are solutions in the absence of sources (via $\text{Im } \vec{S} = \text{Im } \vec{J} = 0$).

Week 4

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