

6. Show that the current proper distance to the particle horizon, the most distant place we can see, for a matter dominated, $k = 0$ universe with no cosmological constant, is given by $r_h R_o = 3ct_o$. Here, r_h is the comoving radial coordinate of the particle horizon, R_o is the current scale factor, and t_o is the current age of the universe. Why is this answer not simply ct_o ? Hint: light moves along null geodesics where $ds^2 = 0$ and the scale factor for the universe described in this problem is $R(t) = R_o(t/t_o)^{2/3}$.
7. A universe has curvature $k = 0$ and $\Omega_\Lambda = 1$. Let $R(0) = R_o$. Find the comoving coordinate of a galaxy that sits at the particle horizon, r_h , at time t . Comment on your result.
8. You are a two-dimensional creature living on a sphere with radius R . Find the circumference of a circle of radius r drawn on the sphere. Suppose you can only measure distances to ± 1 cm, how large of a circle would you need to draw to convince yourself the Earth is spherical in shape.
9. Suppose the energy density of the cosmological constant, ε_Λ , was equal to the critical energy density. How much energy is contained in the energy of the cosmological constant in a sphere of radius 1 Astronomical Unit? Compare this energy to the rest mass energy of the Sun. Does the energy in the cosmological constant play a large role in the dynamics of the Solar System? Explain your answer.
10. By making the substitutions

$$x = r \sin \theta \cos \phi \quad (1)$$

$$y = r \sin \theta \sin \phi \quad (2)$$

$$z = r \cos \theta \quad (3)$$

demonstrate that equations (3.12) and (3.13) represent the same metric.