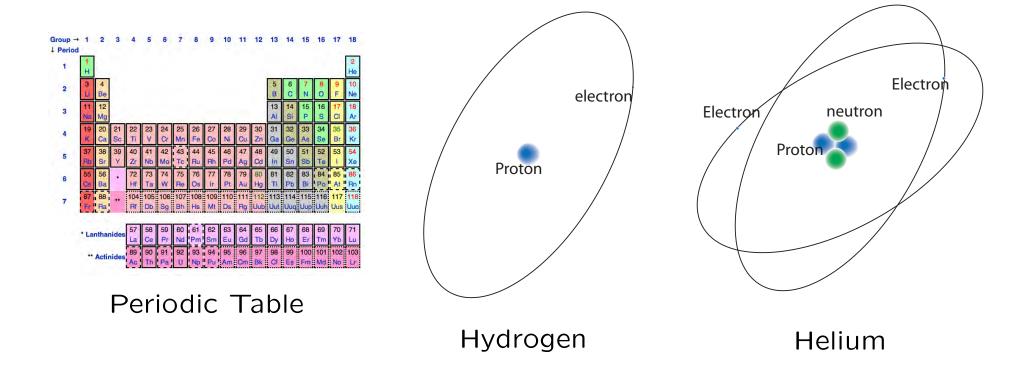
# Dark Matter and the ATLAS Experiment at the Large Hadron Collider



#### Outline:

- 1. A little bit of particle physics
- 2. What particle properties does dark matter have?
- 3. An example of dark matter model we can test at the LHC.
- 4. The LHC and the ATLAS detector

## Quick review of matter:



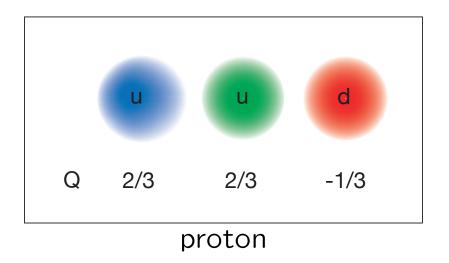
 $\Rightarrow$  All elements consist of electrons circulating around a nucleus made of protons and neutrons

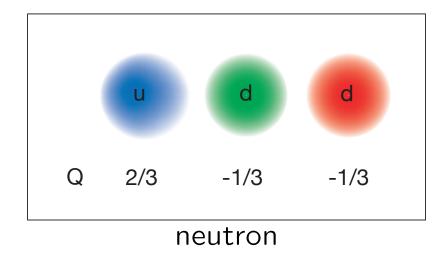
## All of chemistry from three particles!

What are the forces that act on these particles:

- **Gravity** Newton realized that gravity on the Earth and in the heavens are same.
- **Electromagnetism** Faraday, Maxwell and others realized the magnetism and electricity are connected and give rise to radio waves and light .
- **Strong aka QCD** There must be some strong force to hold the protons together in the nucleus!

We now know that protons (and neutrons) are made of three quarks\*:

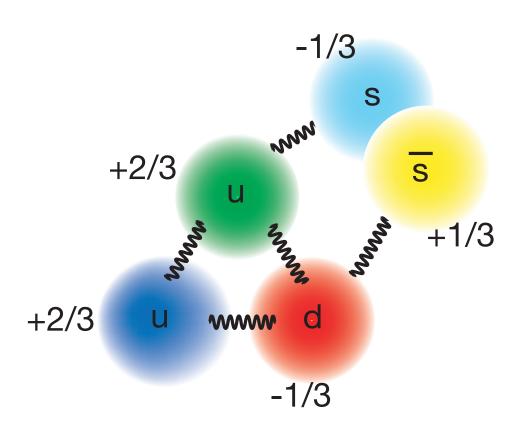




- The quarks interact with each via particles (similar to photons) that are **gluons**.
- These gluons hold the proton together and responsible for the mechanism that binds protons and neutrons.
- The theory that explains strong force is call Quantum ChromoDynamics or QCD.

<sup>\*</sup> Quark model is due to Gell-Mann, name based on fragment of Joyce's Finnegans Wake — "Three quarks for Muster Mark"

Previous particle physics experiments tell us that real protons are more complicated:



In collisions between protons, for example at the LHC, interactions interactions between constituents in the proton are important.

Can we make the Sun shine with just three forces? No.

Example

$$p + p \rightarrow pn \ e^{+} \ \nu_{e}$$

p = proton

n = neutron

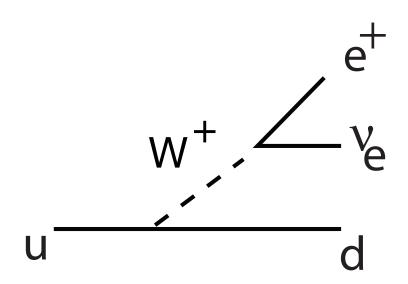
 $e^+ = postiron$ 

 $\nu_e$  = electron neutrion

pn = deuteron

Changing the protons into neutrons requires the weak force!

The force that can change the quark flavor is called the weak force:



- The neutrino was introduced by Pauli to save energy conservation in some kinds of radioactive decay and theory was developed by Fermi. Neutrinos were first "seen" in the early 1950s.
- Fermi's original theory had no "W" particle. This was introduce to save the theory and discovered in accelerators in the early 80s.

Matter

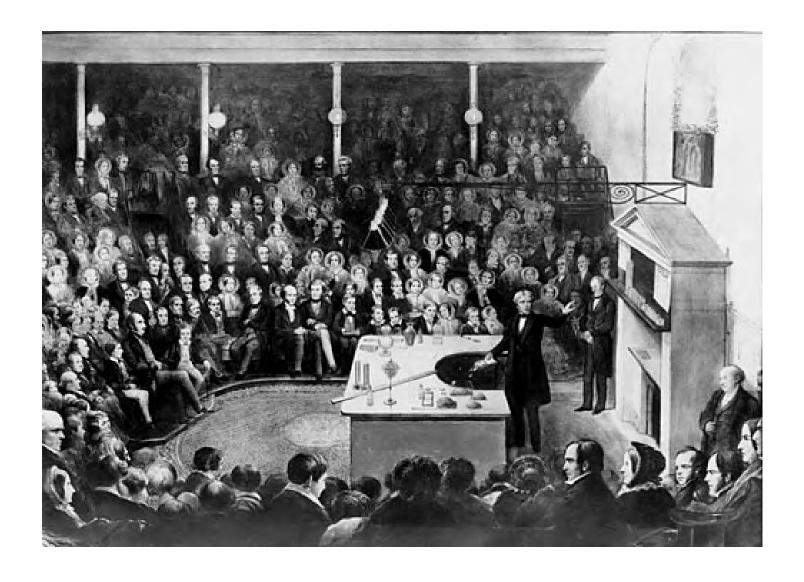
				Charge
Quarks*	u	С	t	2/3
	d	S	b	-1/3
Leptons	е	$\mu$	au	-1
	$ u_e$	$ u_{\mu}$	$ u_{ au}$	O

Forces

Gravity	graviton
Weak	$W^{\pm}$ , $Z^{0}$
Electromagnetic	$\gamma$
Strong	gluon

- Particles important in everyday life are shown in blue
- Each of the matter particles has a corresponding anti-particle
- There are no "anti-particles" for the Force Carriers

# Paradigm of Modern physics started by Michael Faraday



Michael Faraday  $\Rightarrow$  Electricity + Magnetism = radio waves?

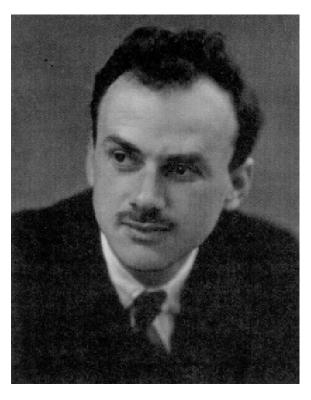
Gladstone: What use are your discoveries?

Faraday: I do not know, but some day you will tax them.

There are very few aspects of modern life not touched by the unification of Electricity and Magnetism:

- Radio-waves
- Electric motors
- Electric power generation

Another application of this paradigm is antimatter ....



P.A.M. Dirac

• Dirac combined Einstein's Relativity with Quantum Mechanics to predict anti-mater (1928)

• In 1932 Anderson discovers Positrons (anti-electrons) in cosmic rays

• Today, main use of anti-mater today is for medical diagnostics (PET)



- **Combining** the Weak and Electromagnetic forces forces us to postulate the famous **Higgs** boson perhaps soon to be discovered at the LHC.
- Trying to combine the the other forces (Weak, Electromagnetism, Strong (and perhaps gravity) leads to theoretical troubles called the **Hierarchy** problem.

Example solutions to this problem:

• **Supersymmetry** (postulates partners for all known particles), e.g.

quarks 
$$ightarrow$$
 squarks  $q 
ightarrow ilde{q}$ 

Extra dimensions

Many variants of these theory have particles that could be the cosmological dark matter

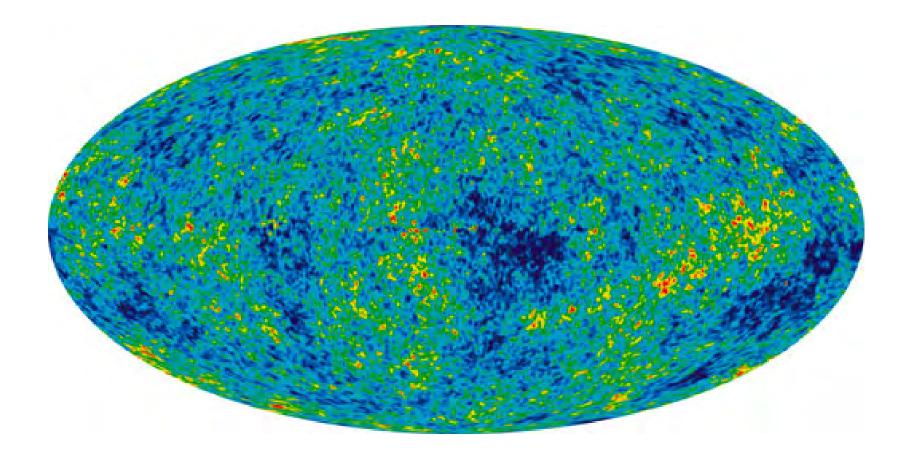
Requirements of dark matter particles:

- Must be very weakly interacting for example dark matter particles must pass through planets and starts without interacting.
- Dark matter particles must be stable on time scales of the age Universe
- Dark matter must have the correct mass and annihilation cross section so that 20% of the closure density of the universe is from dark matter  $\Rightarrow$  These cross section are "coincidentally" the same as weak interaction cross section, hence the term :

WIMP = Weakly Interacting Massive Particle

Must be cold – rules out neutrinos as the dark matter

Results from the Wilkinson Microwave Anisotropy Probe — show fluctuations in the microwave background that can only be explained by cold dark matter:

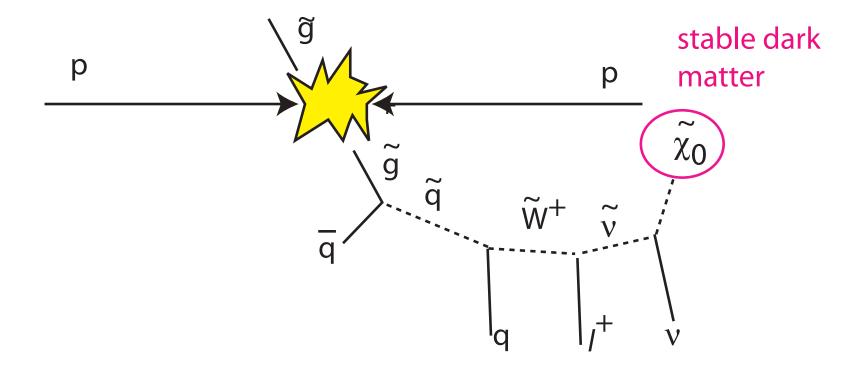


http://map.gsfc.nasa.gov/

#### Neutralino Dark Matter

- In SuperSymmetry (SUSY) the concept of R-parity plays a role similar to charge:
  - ⇒ SUSY particles are always produced in pairs (one SUSY particle and one anti-SUSY particle)
  - ⇒ SUSYness is conserved (similar to charge conservation)
  - ⇒ Lightest SUSY particle must be stable!

### Possible reaction at the LHC



Signature: multiple leptons, missing energy  $(\chi, \nu)$ , jets from quarks (q)

supersymmetric particles have  $\sim$ 

What do we need to observe such events?

- Collisions of very energetic protons (beam with protons of energies
   3.5 to 7 TeV each)
- i.e. The LHC Accelerator

These collisions have a center-of-mass energy of 7 to 14 TeV. 14 TeV is equivalent of rest mass of 14,000 protons

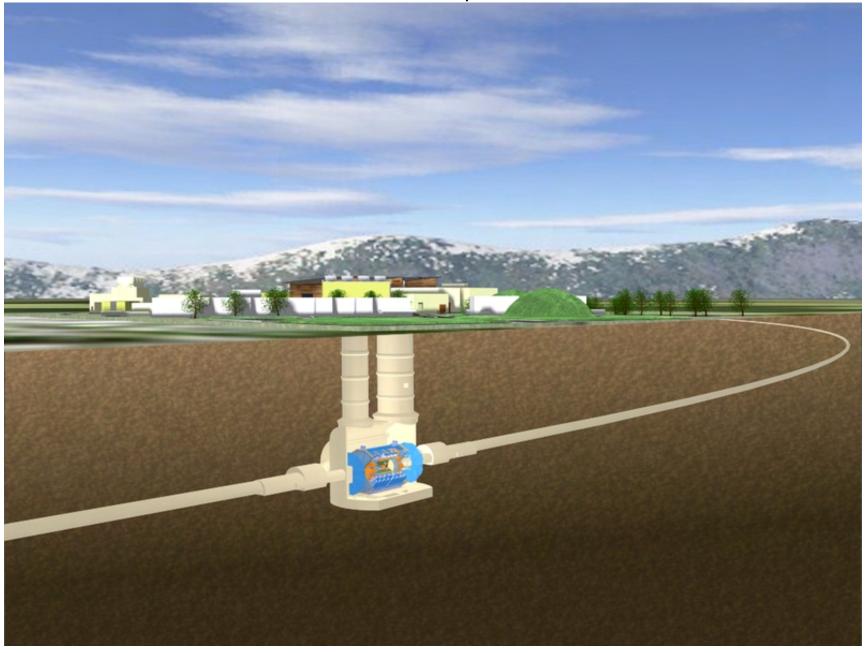
- Excellent detection of muons and electrons (leptons)
- Good detection of jets from quarks
- Hermetic detectors so that missing energy from invisible dark matter particles can be observed.

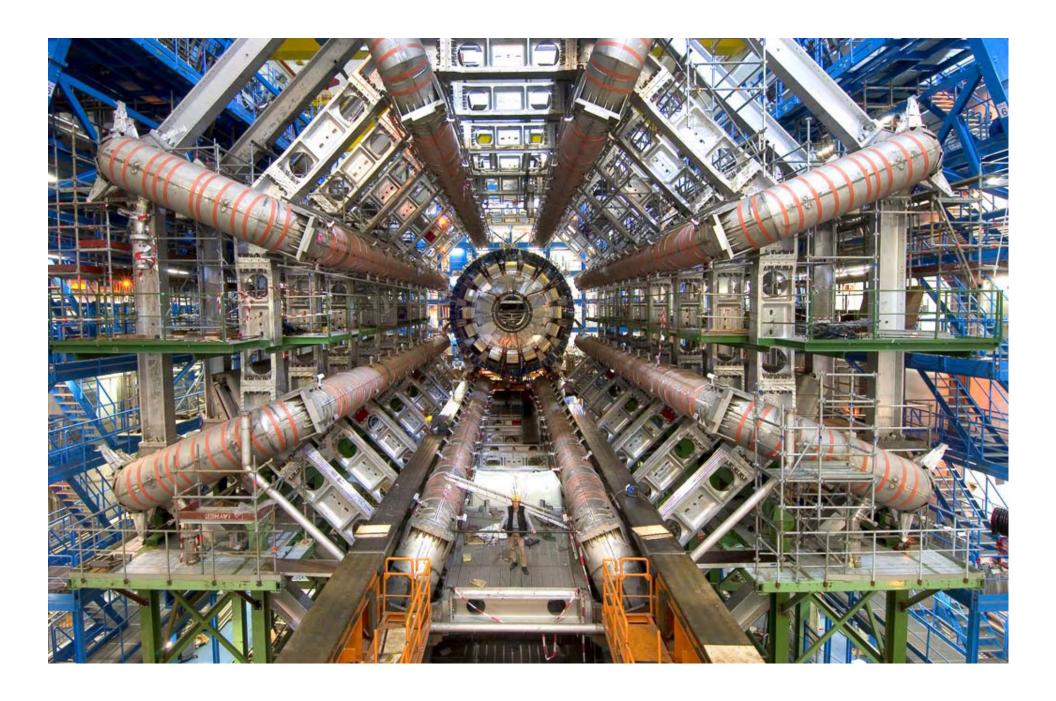
The LHC accelerator – near Geneva, Switzerland

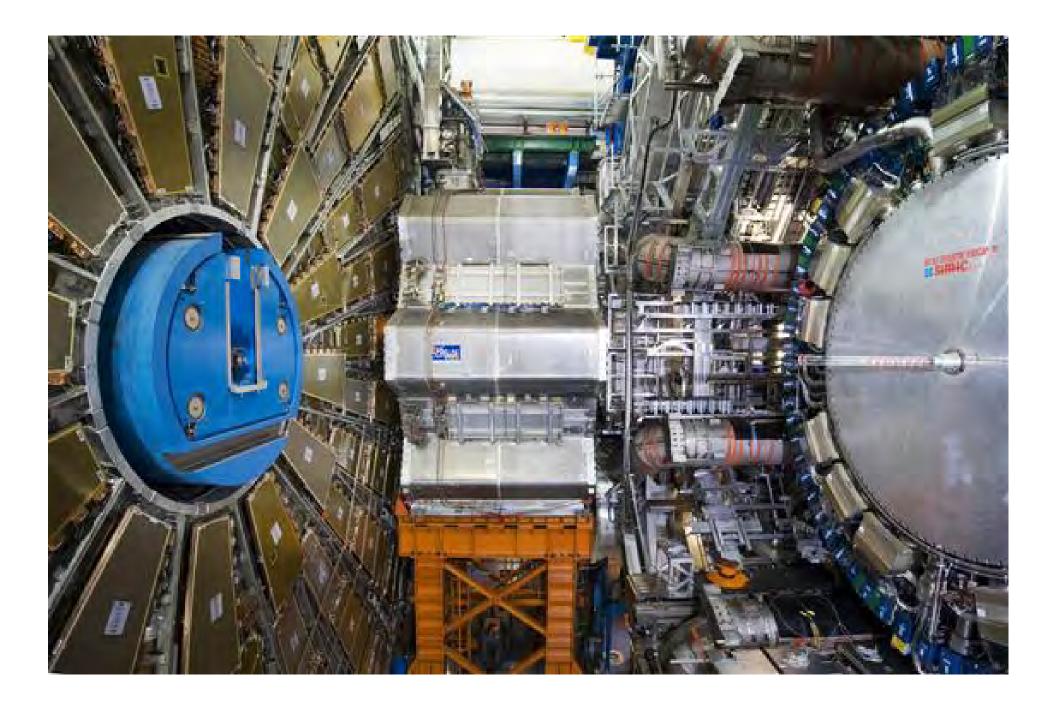


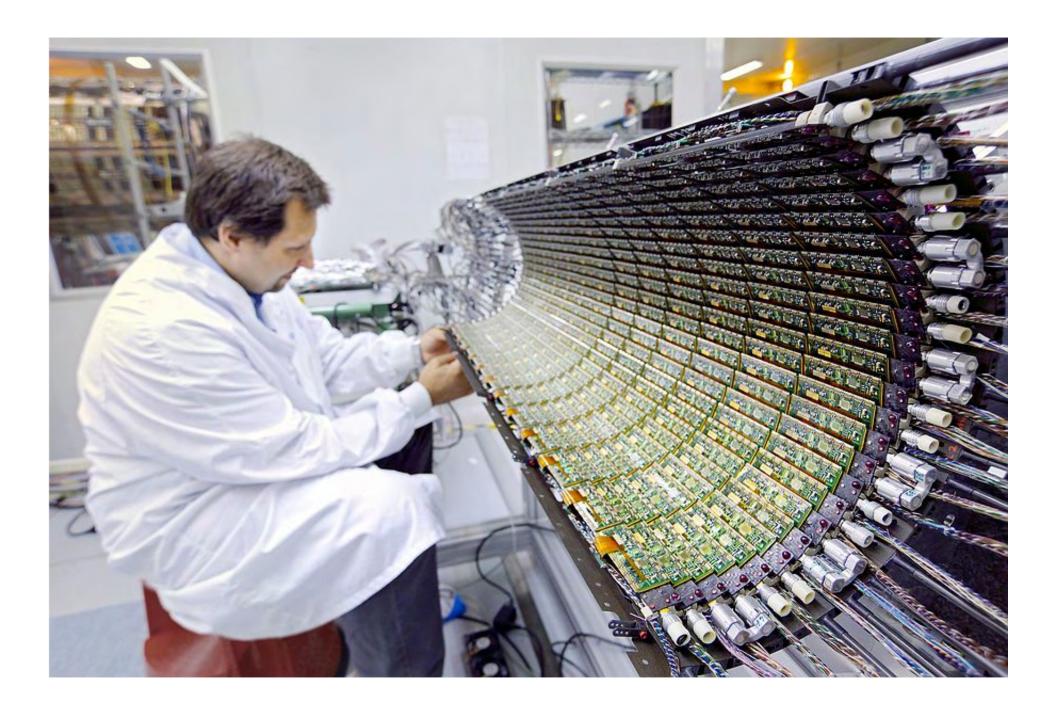


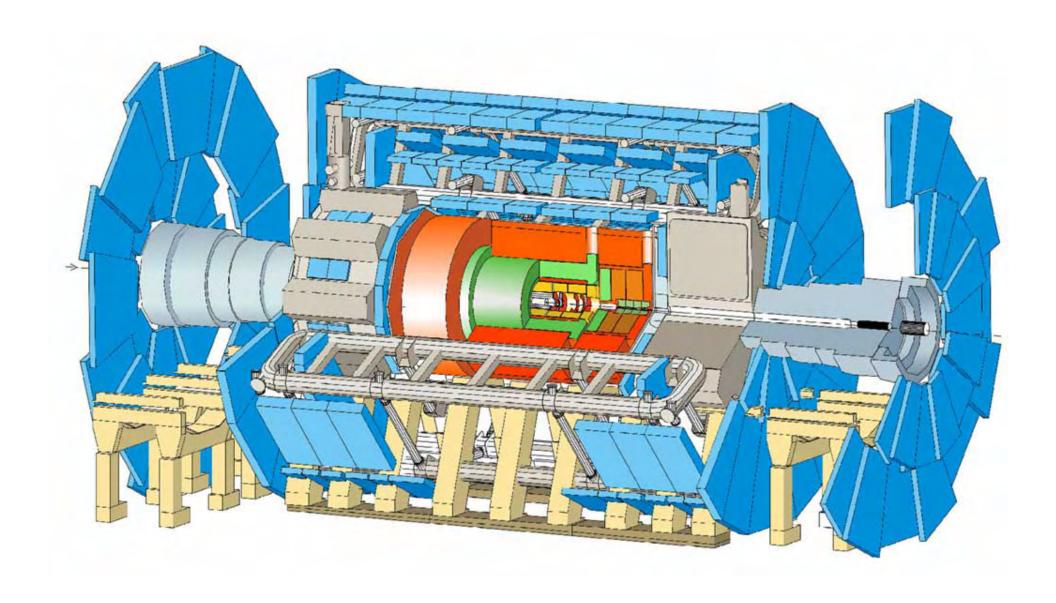
The Atlas Experiment











- http://www.atlas.ch/multimedia/index.html to see detector constuction
- http://atlas.web.cern.ch/Atlas/public/EVTDISPLAY/events.html shows an animation of an event display
- The following candidate events display show examples of events of the type

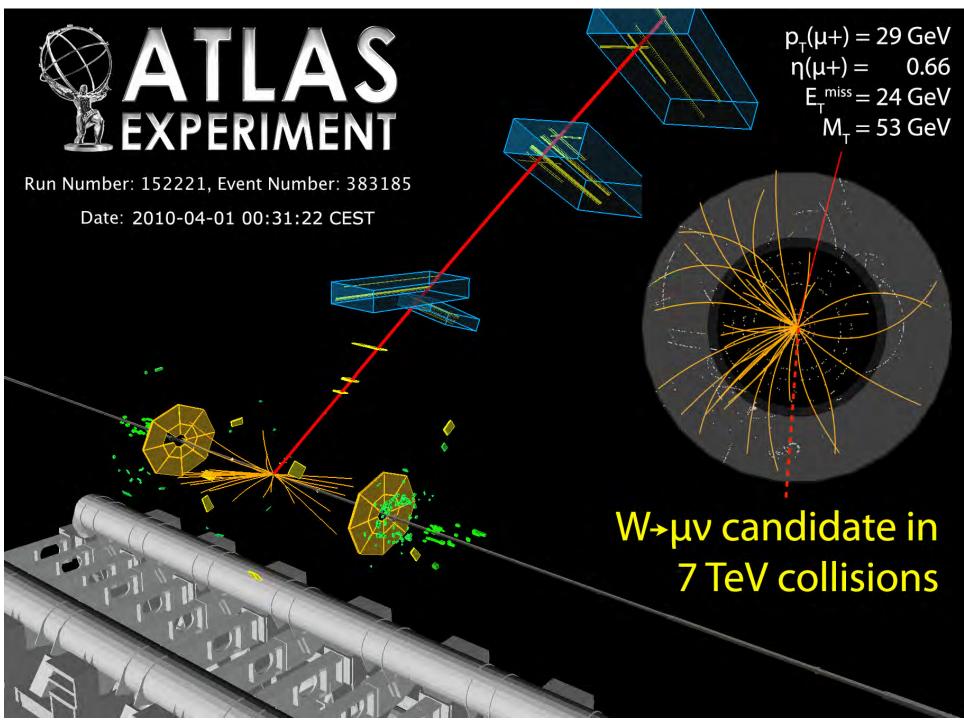
$$p + p \Rightarrow XW^+$$

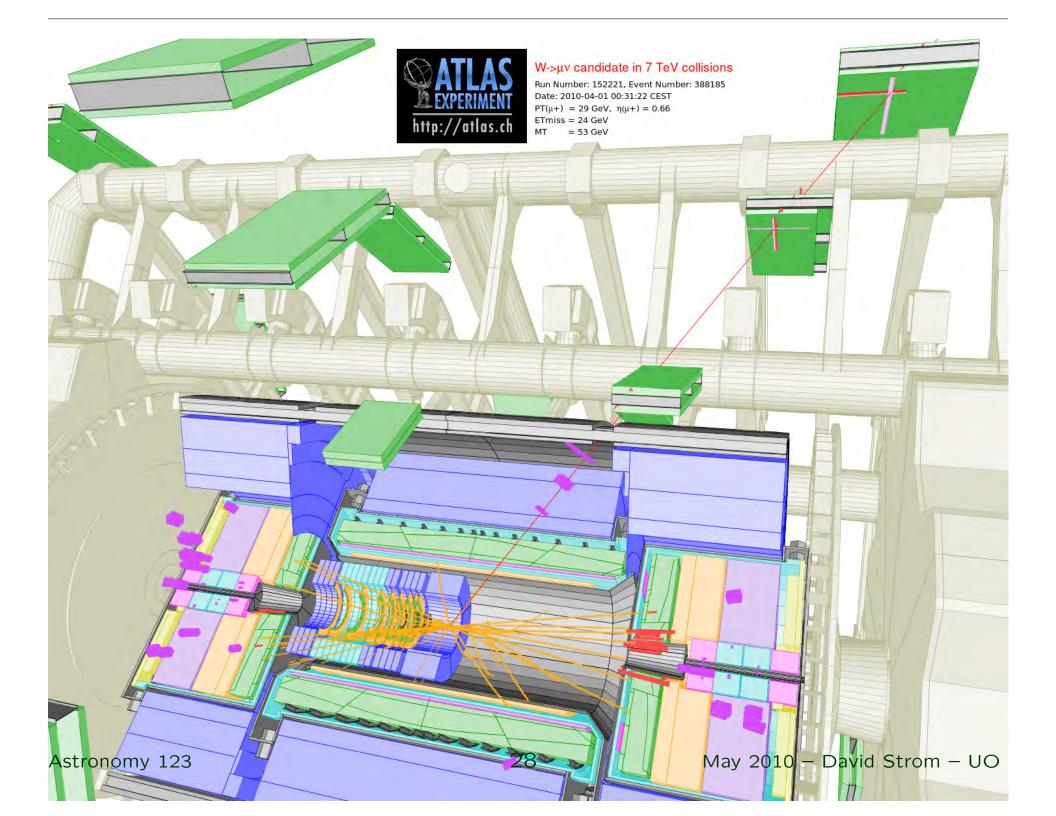
with the W decaying to either

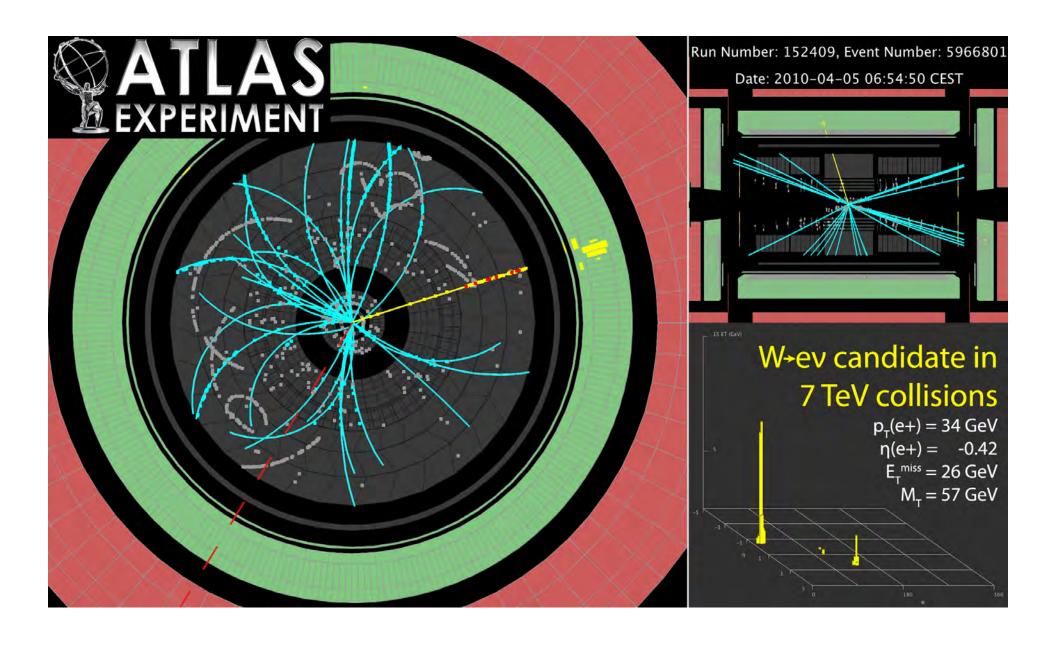
$$W^+ \rightarrow \mu^+ \nu_\mu$$

or

$$W^+ \to e^+ \nu_e$$



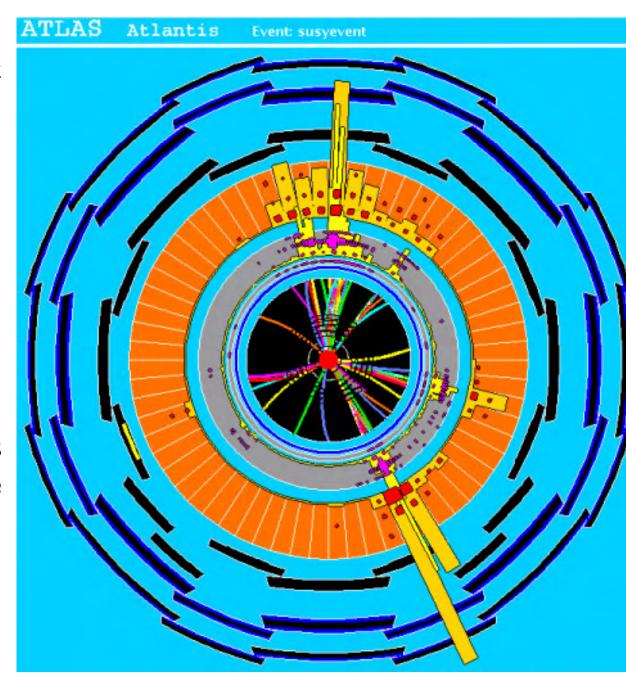




How a supersymmetry event might look in the ATLAS detector

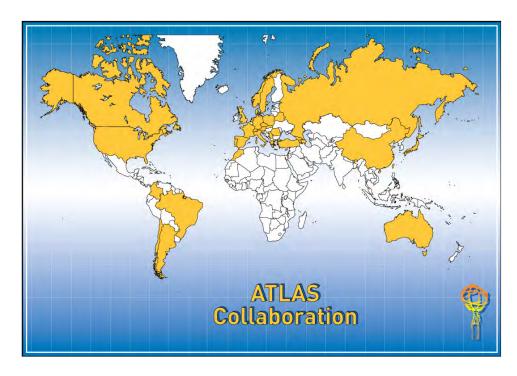
• A schematic end-view of an event

 Asymmetric form indicates missing particles – possible dark matter candidate



## How does Oregon Fit in?

- Four UO experimental faculty
- Eventually we will have four graduate students and two postdoctorial fellows at CERN
- There are five theoretical faculty at UO working on problems related to the LHC



 $\sim$  3000 Physicists

The Oregon experimental group is working mainly on the "Trigger" for the ATLAS experiment.

# The "Trigger"

- There will be more than 40 million proton-proton collisions per seconds
- Each compressed ATLAS event is 25 megabytes

In one second this is

1 petabyte = 1000 000 000 000 000 Bytes

In one year of running this corresponds to

10 zettabytes = 10 000 000 000 000 000 000 000 Bytes

The total amount of stored data in the world is expected to reach a ZettaByte/year in 2010

- To handle the enormous data rate, we use thousands of computers to pick out 100 to 200 events per second for permanent storage.
- At UO, we are developing the software that selects the events to be saved.



We will only get one chance to save the interesting events!

# Looking towards the future

- Luminosity is currently rapidly increasing (currently a factor of 10<sup>6</sup> below final value).
- In the next few months ATLAS will concentrate on what we call Standard Model processes, e.g. Z and W production
- We will have the first chance to discover SUSY this summer.
- In approximately 2 years the energy of LHC will be doubled from 7 TeV
   to 14 TeV
- Full exploitation of the LHC will take us at least another 10 years