

Physics 353: Problem Set 8 – SOLUTIONS

1 Mean field Ising magnet revisited.

a.

$$\tanh x = \tanh 0 + \left. \frac{d}{dx} \tanh x \right|_{x=0} x + \frac{1}{2} \left. \frac{d^2}{d^2 x} \tanh x \right|_{x=0} x^2 + \frac{1}{6} \left. \frac{d^3}{d^3 x} \tanh x \right|_{x=0} x^3 + \dots, \text{ Taylor expanding}$$

about  $x = 0$ .

$$\frac{d}{dx} \tanh x = \operatorname{sech}^2 x \text{ (i.e. } \frac{1}{\cosh^2 x} \text{)}$$

$$\frac{d^2}{d^2 x} \tanh x = \frac{d}{dx} \frac{1}{\cosh^2 x} = -2 \frac{1}{\cosh^3 x} \sinh x = -2 \tanh x \operatorname{sech}^2 x$$

$$\frac{d^3}{d^3 x} \tanh x = -2 \operatorname{sech}^2 x \frac{d}{dx} \tanh x + -2 \tanh x \frac{d}{dx} \operatorname{sech}^2 x = -2 \operatorname{sech}^4 x + 4 \tanh^2 x \operatorname{sech}^2 x$$

Noting that  $\tanh(0) = 0$  and  $\operatorname{sech}(0) = 1$ ,

$$\tanh x = 0 + 1x + 0x^2 + \frac{1}{6}(-2)x^3 + \dots$$

$$\tanh x \approx x + \frac{1}{3}x^3.$$

b. From our examination of the mean field Ising model:

$$\langle s \rangle = \tanh \left( \frac{JN_n}{\tau} \langle s \rangle \right) = \tanh \left( \frac{\tau_c}{\tau} \langle s \rangle \right), \text{ using } \tau_c = JN_n. \text{ Near the phase transition, } \langle s \rangle \text{ is small.}$$

$$\text{Therefore } \langle s \rangle \approx \left( \frac{\tau_c}{\tau} \langle s \rangle \right) - \frac{1}{3} \left( \frac{\tau_c}{\tau} \langle s \rangle \right)^3.$$

$$1 \approx \frac{\tau_c}{\tau} - \frac{1}{3} \left( \frac{\tau_c}{\tau} \right)^3 \langle s \rangle^2, \text{ so } \langle s \rangle \approx \left[ 3 \left( \frac{\tau}{\tau_c} \right)^3 \left( \frac{\tau_c}{\tau} - 1 \right) \right]^{1/2}, \text{ and so } \langle s \rangle \approx \left[ 3 \left( \frac{\tau}{\tau_c} \right)^2 \left( \frac{\tau_c - \tau}{\tau_c} \right) \right]^{1/2}. \text{ We see}$$

that  $\langle s \rangle \propto (\tau_c - \tau)^{1/2}$ , i.e. the critical exponent  $\beta = 1/2$ .

(In case you're worried about the other factors of  $\tau$ , try writing  $t \equiv \tau_c - \tau$ , with which

$$\langle s \rangle \approx \sqrt{3} \left( 1 - \frac{t}{\tau_c} \right) \left( \frac{t}{\tau_c} \right)^{1/2}. \text{ For } \tau \text{ close to } \tau_c, \text{ the first factor is just something near 1. The second}$$

factor changes sharply with  $t$ , like  $t^{1/2}$ .

c. Considering nonzero magnetic field. In the mean field treatment,

$$E_i = -\lambda B s_i - J s_i \langle s \rangle = -(\lambda B + J \langle s \rangle) s_i. \text{ Note that the magnetic field just adds a term that}$$

“combines” with our coupling factor. Therefore the solution to  $\langle s \rangle$  is simply

$$\langle s \rangle = \tanh\left(\frac{JN_n \langle s \rangle + \lambda B}{\tau}\right) = \tanh\left(\frac{\tau_c \langle s \rangle + \lambda B}{\tau}\right). \text{ I'll write two solutions.}$$

**Approach 1:** Again considering small  $\langle s \rangle$  and  $B$  near the critical point,

$$\langle s \rangle \approx \left(\frac{\tau_c \langle s \rangle + \lambda B}{\tau}\right) - \frac{1}{3} \left(\frac{\tau_c \langle s \rangle + \lambda B}{\tau}\right)^3. \text{ Differentiating both sides with respect to } B,$$

$$\chi = \frac{d\langle s \rangle}{dB} = \left(\frac{\tau_c}{\tau} \chi + \frac{\lambda}{\tau}\right) - \left(\frac{\tau_c \langle s \rangle + \lambda B}{\tau}\right)^2 \left(\frac{\tau_c}{\tau} \chi + \frac{\lambda}{\tau}\right) \text{ (Chain rule !)}$$

$$\text{Evaluating this as } B \rightarrow 0 \text{ and } \langle s \rangle \rightarrow 0, \chi = \frac{\tau_c}{\tau} \chi + \frac{\lambda}{\tau}, \text{ so } \chi = \frac{\lambda}{\tau - \tau_c}.$$

$$\text{We see that } \chi = \frac{d\langle s \rangle}{dB} \propto (\tau - \tau_c)^{-\gamma} \text{ with } \gamma = 1.$$

**Approach 2:** Avoiding Taylor expansion. From above  $\langle s \rangle = \tanh\left(\frac{\tau_c \langle s \rangle + \lambda B}{\tau}\right)$ . Differentiate:

$$\chi = \frac{d\langle s \rangle}{dB} = \text{sech}^2\left(\frac{\tau_c \langle s \rangle + \lambda B}{\tau}\right) \left[\frac{\tau_c}{\tau} \chi + \frac{\lambda}{\tau}\right] \text{ (Again, Chain rule !). Evaluate as } B \rightarrow 0 \text{ and}$$

$$\langle s \rangle \rightarrow 0, \text{ noting that } \text{sech}^2(0) = 1:$$

$$\chi = 1 \left[\frac{\tau_c}{\tau} \chi + \frac{\lambda}{\tau}\right], \text{ so } \chi \left(1 - \frac{\tau_c}{\tau}\right) = \frac{\lambda}{\tau}, \text{ and therefore } \chi = \frac{\lambda}{\tau - \tau_c}.$$

$$\text{We see that } \chi = \frac{d\langle s \rangle}{dB} \propto (\tau - \tau_c)^{-\gamma} \text{ with } \gamma = 1.$$

## 2 Miscibility transition.

$$u(x) = 0 \text{ @ } x=0, 1 ; \quad u(x) = u_0 N \text{ @ } x = \frac{1}{2}$$

parabolic, centered @  $x = \frac{1}{2}$ , concave down

$$\Rightarrow u(x) = 4u_0 \left[ -\left(x - \frac{1}{2}\right)^2 + \frac{1}{4} \right] N$$

$$u(x) = Nu_0 \left( 1 - 4\left(x - \frac{1}{2}\right)^2 \right)$$

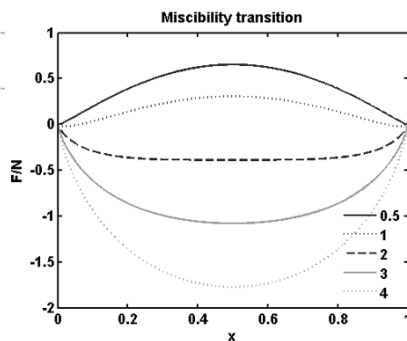
$$\text{note } x=0 \Rightarrow u=0, \quad x=1 \Rightarrow u=0, \quad x=\frac{1}{2} \Rightarrow u=u_0 N$$

$$\text{Entropy of mixing } \sigma_{\text{mix}} = -N [x \ln x + (1-x) \ln(1-x)]$$

$$\text{Free energy } F \equiv u - T\sigma$$

$$F = Nu_0 \left[ 1 - 4\left(x - \frac{1}{2}\right)^2 \right] + T N [x \ln x + (1-x) \ln(1-x)]$$

Plot



Phase separation if  $F(x)$  is concave-down, i.e.  $\frac{d^2 F}{dx^2} < 0$

$$\frac{dF}{dx} = Nu_0 \left[ -8\left(x - \frac{1}{2}\right) \right] + T N \left[ 1 + \ln x + -1 - \ln(1-x) \right]$$

$$= Nu_0 [-8x + 4] + T N \left[ \ln\left(\frac{x}{1-x}\right) \right]$$

$$\frac{d^2 F}{dx^2} = Nu_0 (-8) + T N \left( \left(\frac{1-x}{x}\right) \left(\frac{1}{(1-x)^2}\right) + \left(\frac{1}{1-x}\right) \right)$$

$$= -8Nu_0 + T N \left[ \frac{1}{1-x} + \frac{1}{x} \right]$$

$$= -8Nu_0 + T N \left( \frac{1}{x(1-x)} \right)$$

just consider  $x = \frac{1}{2}$

$$\left. \frac{d^2 F}{dx^2} \right|_{x=\frac{1}{2}} = \frac{-8Nu_0 + T N}{x(1-x)} \Big|_{x=\frac{1}{2}} = -8Nu_0 + 4NT$$

If  $-8Nu_0 + 4NT \begin{cases} < 0, \text{ phase sep.} \\ > 0, \text{ miscible (mixed)} \end{cases}$

$\Rightarrow$  phase sep. for  $-2u_0 + T < 0 \Rightarrow \boxed{T < 2u_0}$

critical temperature  $T_c = 2u_0$

### 3 Clausius-Clapeyron relation.

- a. Melting ice at  $-1\text{ }^\circ\text{C}$ .  $\frac{dP}{d\tau} = \frac{L}{\tau\Delta v}$ , where  $L$  is the latent heat per particle and  $\Delta v$  is the volume change per particle. We need to understand all the variables and convert them to appropriate units. Note that  $dT$ , one degree, is much smaller than  $T$ ,  $273\text{ K}$ , so we can write that the increase in pressure  $dP = \frac{L}{\Delta v} \frac{d\tau}{\tau}$ .

A long route: The density of water is  $1000\text{ kg/m}^3$ , and each water molecule is  $18 \times 1.67 \times 10^{-27}\text{ kg} = 3.01 \times 10^{-26}\text{ kg}$ , so the volume per water molecule is  $3.01 \times 10^{-26} / 1000 = 3.01 \times 10^{-29}\text{ m}^3$ . Similarly the volume per ice molecule is  $3.0 \times 10^{-26} / 917 = 3.28 \times 10^{-29}\text{ m}^3$ , so  $\Delta v = -0.27 \times 10^{-29}\text{ m}^3$ .  $L = 334\text{ kJ/kg}$  in units of heat per unit mass. Per particle,  $L$  is therefore  $334 \times 10^3 \times 3.01 \times 10^{-26} = 1.01 \times 10^{-20}\text{ J}$ . The temperature  $\tau = k_B \times 273\text{ K} = 3.77 \times 10^{-21}\text{ J}$ .

Therefore  $\frac{dP}{d\tau} = \frac{1.01 \times 10^{-20}}{3.77 \times 10^{-21} (-0.27 \times 10^{-29})} = -9.92 \times 10^{29} \frac{\text{Pa}}{\text{J}}$ . In conventional temperature

units, multiplying by Boltzmann's constant,  $\frac{dP}{dT} = -1.37 \times 10^7 \frac{\text{Pa}}{\text{K}}$ .

**More succinctly:** write  $\rho$  = density,  $m$  = mass of a water molecule, so volume per particle

$$v = (\rho/m)^{-1}, \text{ i.e. } 1/\text{concentration. } \Delta v = \frac{m}{\rho_{\text{water}}} - \frac{m}{\rho_{\text{ice}}} = m \left( \frac{1}{\rho_{\text{water}}} - \frac{1}{\rho_{\text{ice}}} \right) = m \left( \frac{\rho_{\text{ice}} - \rho_{\text{water}}}{\rho_{\text{water}} \rho_{\text{ice}}} \right).$$

The latent heat in units of energy is  $L = L'/m$ , where  $L'$  is the latent heat in the given units of energy / mass. Therefore:  $dP = \frac{L}{\Delta v} \frac{d\tau}{\tau} = \frac{dT}{T} \frac{L}{\Delta v} = \frac{dT}{T} \frac{L' \rho_{\text{water}} \rho_{\text{ice}}}{(\rho_{\text{ice}} - \rho_{\text{water}})}$ . Note that the  $m$ 's and

$$k_B \text{'s cancel! Plugging in numbers (SI units): } dP = \frac{-1}{273} \frac{334 \times 10^3 \times 917 \times 1000}{(917 - 1000)} = 1.35 \times 10^7 \text{ Pa.}$$

**Therefore:** to lower the melting temperature by one degree, you'd need to apply **13.7 million Pascal** of pressure above normal atmospheric pressure (about  $10^5\text{ Pa}$ ).

- b. Consider a glacier... Using part (a), we need a pressure of 13.7 million Pascal. This pressure is provided by the weight of the ice  $\rho g A h$ , where  $\rho$  is the density and  $h$  is the height, divided by its area  $A$ . Using  $\rho$  from above, we therefore need  $\rho g A h / A = \rho g h = 1.37 \times 10^7\text{ Pa/K}$ , so  $h = 1.53 \times 10^3$  meters, or about **one-and-a-half kilometers**.

- c. A 50 kg person skating on an ice skate with a blade with dimensions around  $20\text{ cm} \times 0.1\text{ mm}$  applies a pressure of  $2.5 \times 10^7\text{ Pa}$ . This is certainly enough to lower the melting temperature of ice by 1 degree. Applying  $\frac{dP}{dT} = -1.37 \times 10^7 \frac{\text{Pa}}{\text{K}}$  from above, it will lower the melting temperature by about  $2\text{ }^\circ\text{C}$ . However, one can (and usually does) ice skate at temperatures considerably lower than  $-2\text{ }^\circ\text{C}$ . When skating on  $-10\text{ }^\circ\text{C}$  ice it doesn't matter that your weight makes the melting temperature  $-2\text{ }^\circ\text{C}$  – the ice is still solid! Ice skating is not made possible by the Clausius-Clapeyron relation, but rather by heat generated by friction.

4 The one-dimensional Ising model.

$$a) \sum_{S_N} e^{\beta J S_{N-1} S_N}$$

$$= e^{\beta J S_{N-1}} + e^{-\beta J S_{N-1}}$$

$$\text{if } S_{N-1} = 1, \text{ then } \sum_{S_N} e^{\beta J S_{N-1} S_N} = e^{\beta J} + e^{-\beta J} = 2 \cosh(\beta J)$$

$$\text{if } S_{N-1} = -1, \text{ then } \sum_{S_N} e^{\beta J S_{N-1} S_N} = e^{-\beta J} + e^{\beta J} = 2 \cosh(\beta J)$$

$$(b) \sum_{S_2=\pm 1} \sum_{S_3=\pm 1} \dots \sum_{S_N=\pm 1} e^{\beta J S_1 S_2} e^{\beta J S_2 S_3} \dots e^{\beta J S_{N-1} S_N}$$

$$= 2 \cosh(\beta J) \sum_{S_2=\pm 1} \dots \sum_{S_N=\pm 1} e^{\beta J S_1 S_2} \dots e^{\beta J S_{N-2} S_{N-1}}$$

$$= (2 \cosh(\beta J))^2 \sum_{S_2=\pm 1} \dots \sum_{S_{N-2}=\pm 1} e^{\beta J S_1 S_2} \dots e^{\beta J S_{N-3} S_{N-2}}$$

$$= (2 \cosh(\beta J))^{N-1}$$

$$(c) \therefore Z = \sum_{S_1=\pm 1} \left( \sum_{S_2=\pm 1} \dots \sum_{S_N=\pm 1} e^{\beta J S_1 S_2} e^{\beta J S_2 S_3} \dots e^{\beta J S_{N-1} S_N} \right)$$

$$= \sum_{S_1=\pm 1} 2^{N-1} [\cosh(\beta J)]^{N-1}$$

$$= 2^N [\cosh(\beta J)]^{N-1}$$

for large  $N$ ,  $N-1 \approx N$

$$\therefore Z \approx [2 \cosh(\beta J)]^N$$

$$(d) U = -\frac{\partial}{\partial \beta} \ln Z = -\frac{\partial}{\partial \beta} (N (\ln 2 + \ln \cosh(\beta J)))$$

$$= -N \frac{\partial}{\partial \beta} [\ln \cosh(\beta J)]$$

$$= +NJ \frac{e^{\beta J} - e^{-\beta J}}{e^{\beta J} + e^{-\beta J}} = NJ \tanh(\beta J)$$

It's a continuous function, so there is no phase transition in one dimension.

### 5 Free energy of the mean field Ising model.

From class,  $U(\langle s \rangle) = -\frac{N}{2} J N_n \langle s \rangle^2$

$\sigma(s) = N \ln 2 - \frac{N}{2} \langle s \rangle^2 - \alpha N \langle s \rangle^4$  ,  $F \equiv U - T\sigma$

$F(\langle s \rangle) = \frac{N}{2} \left\{ (-J N_n + T) \langle s \rangle^2 - 2T \ln 2 + 2T\alpha \langle s \rangle^4 \right\}$

$\frac{\partial F}{\partial \langle s \rangle} = \frac{N}{2} \left\{ 2(-J N_n + T) \langle s \rangle + 8T\alpha \langle s \rangle^3 \right\}$

$= 0 \Rightarrow \langle s \rangle = 0$  or

$(-J N_n + T) + 4T\alpha \langle s \rangle^2 = 0$

$\Rightarrow \langle s \rangle^2 = \frac{-T + J N_n}{4T\alpha}$

If  $T > J N_n$ ,  $\langle s \rangle^2 < 0$ ,  
not possible so  $\langle s \rangle = 0$  is the  
only solution.

If  $T \leq J N_n$ ,  $\langle s \rangle^2$  has a solution  
 $\Rightarrow T_c = J N_n$  (as seen before).

$\Rightarrow \langle s \rangle^2 = \frac{T_c - T}{4\alpha T}$  for  $T \leq T_c$

$\langle s \rangle = \frac{1}{2\sqrt{\alpha}} \left( \frac{T_c}{T} - 1 \right)^{1/2}$  for  $T \leq T_c$

$\langle s \rangle = 0$  for  $T > T_c$

Smoothly increasing as  $T$  drops below  $T_c$ .